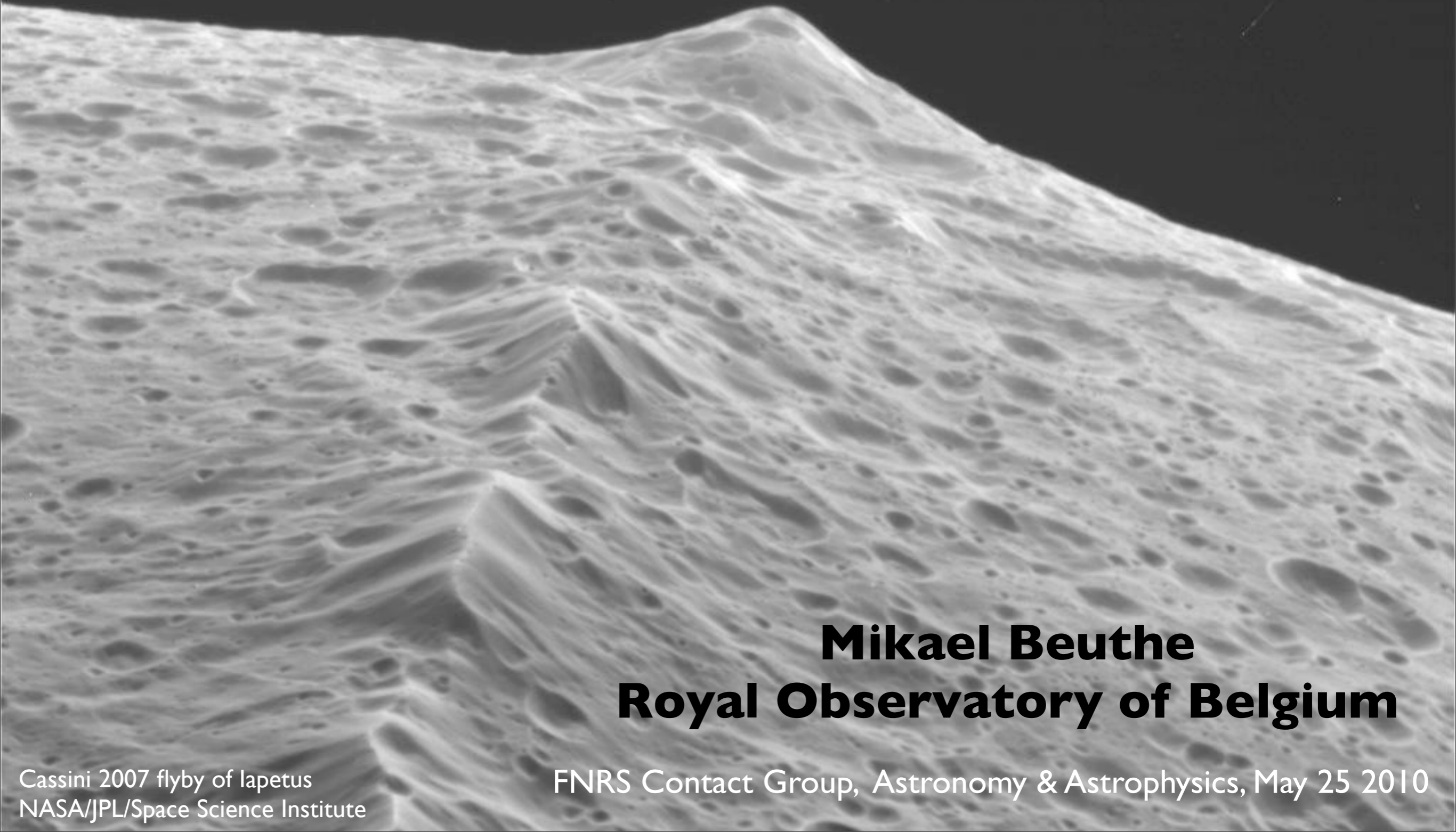
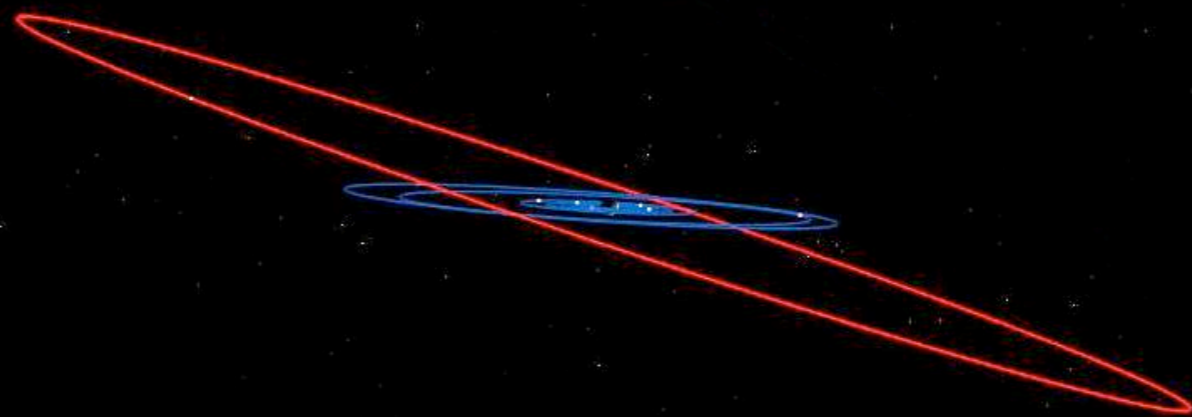
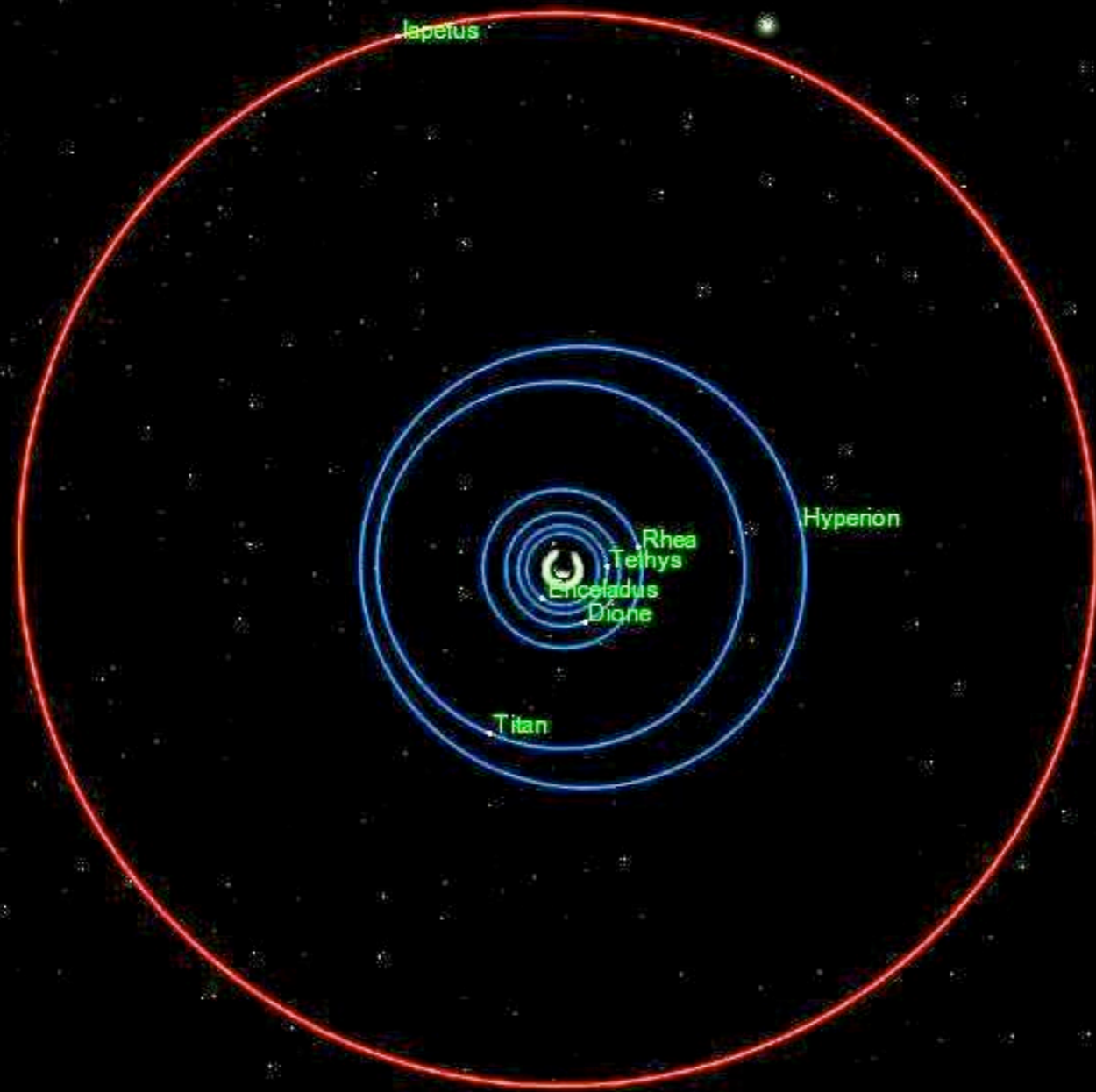


Iapetus' ridge,
the remnant of a global contraction event

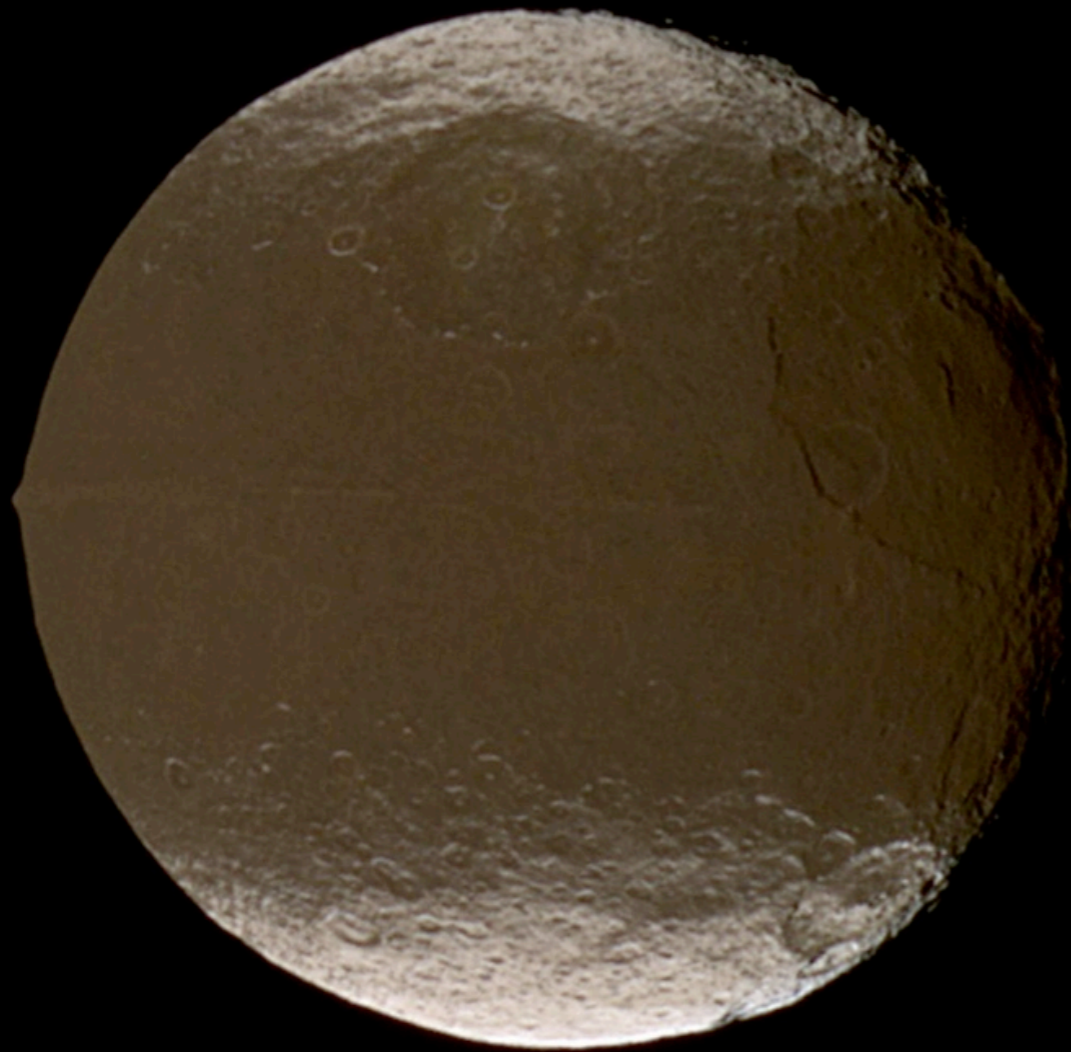


Mikael Beuthe
Royal Observatory of Belgium



Mystery I: Brightness dichotomy

Leading side: dark
Cassini 1705



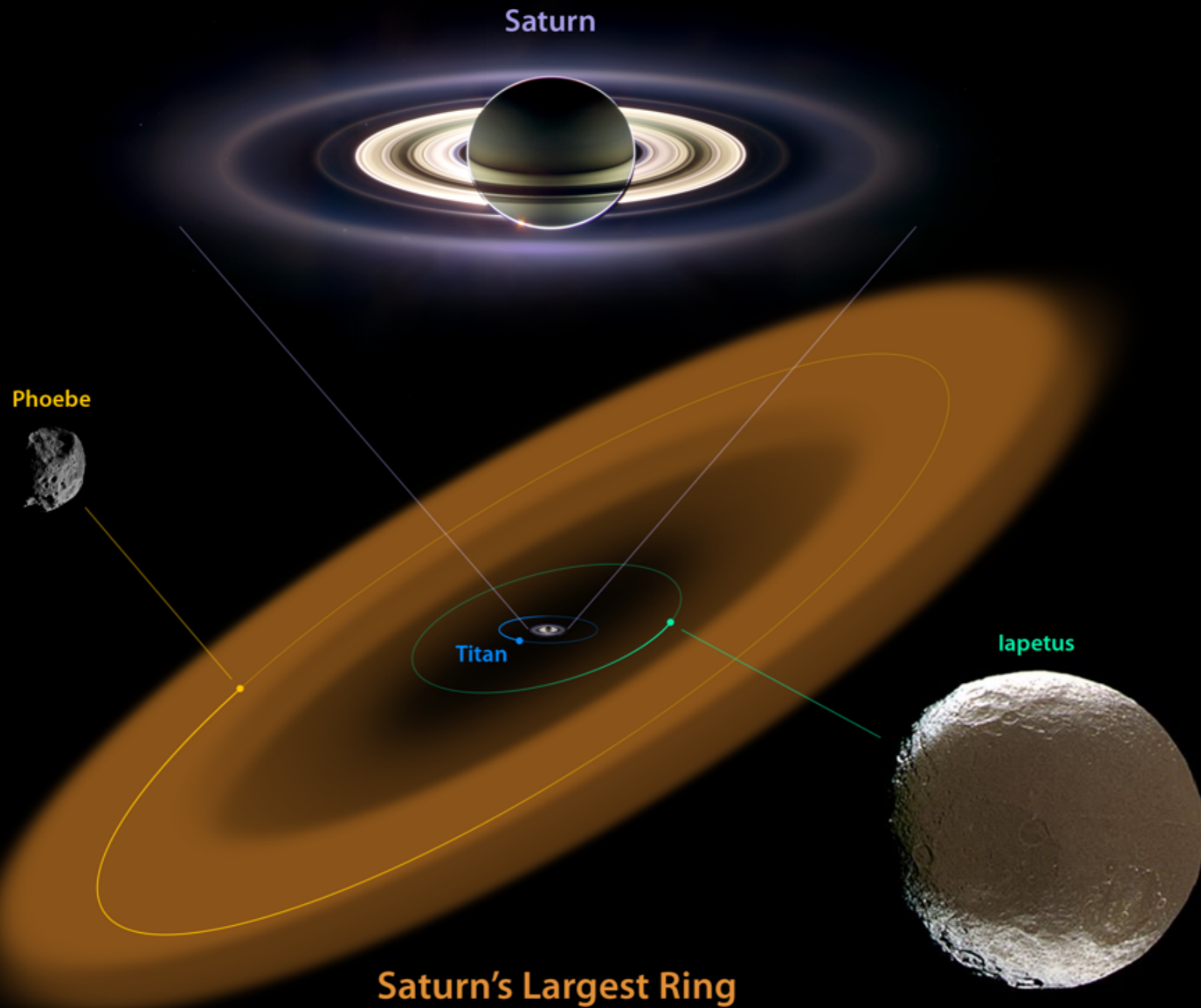
Cassini flyby 2004

Trailing side: bright
Cassini 1671



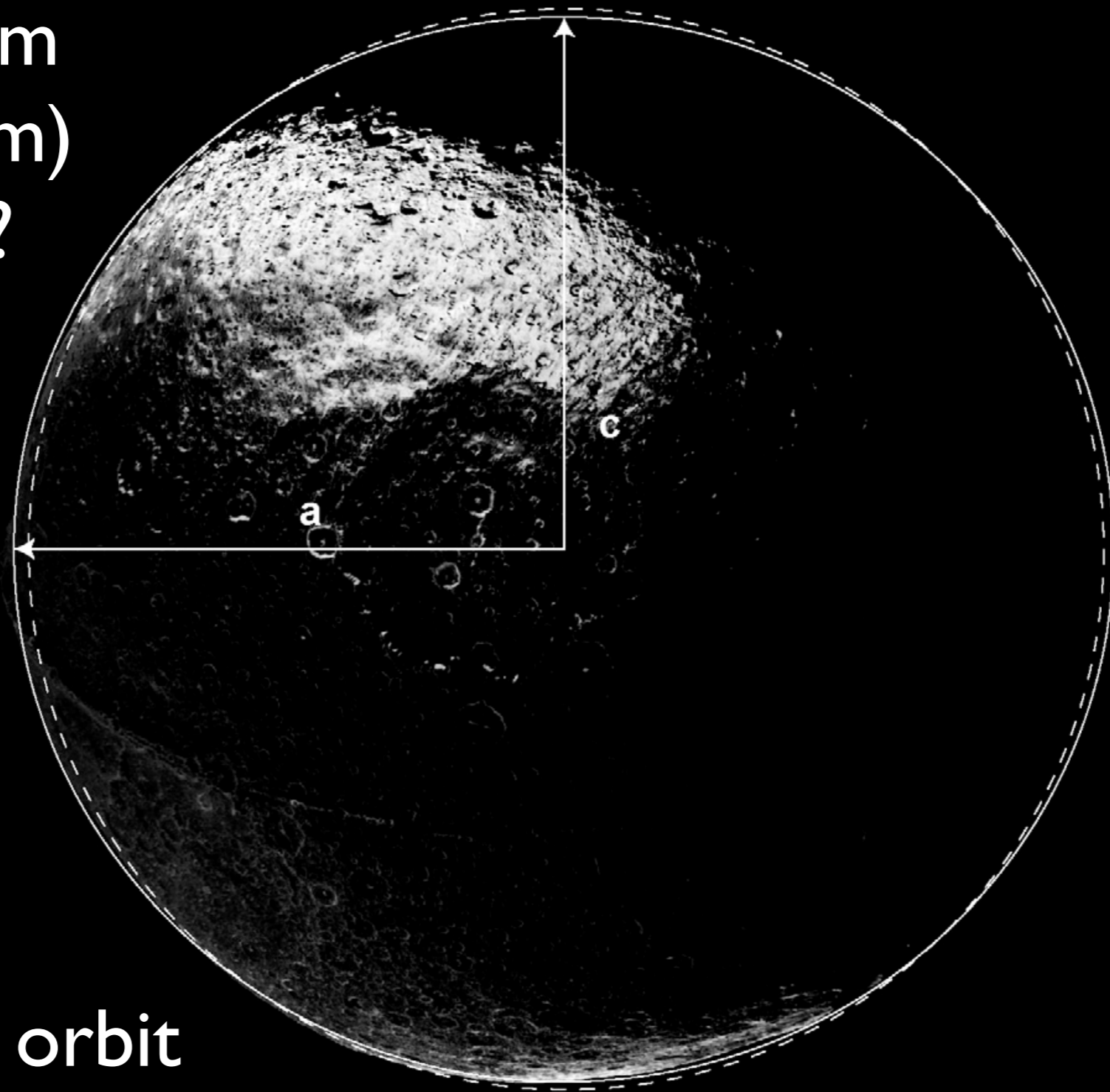
Cassini flyby 2007

2009: Saturn's colossal ring



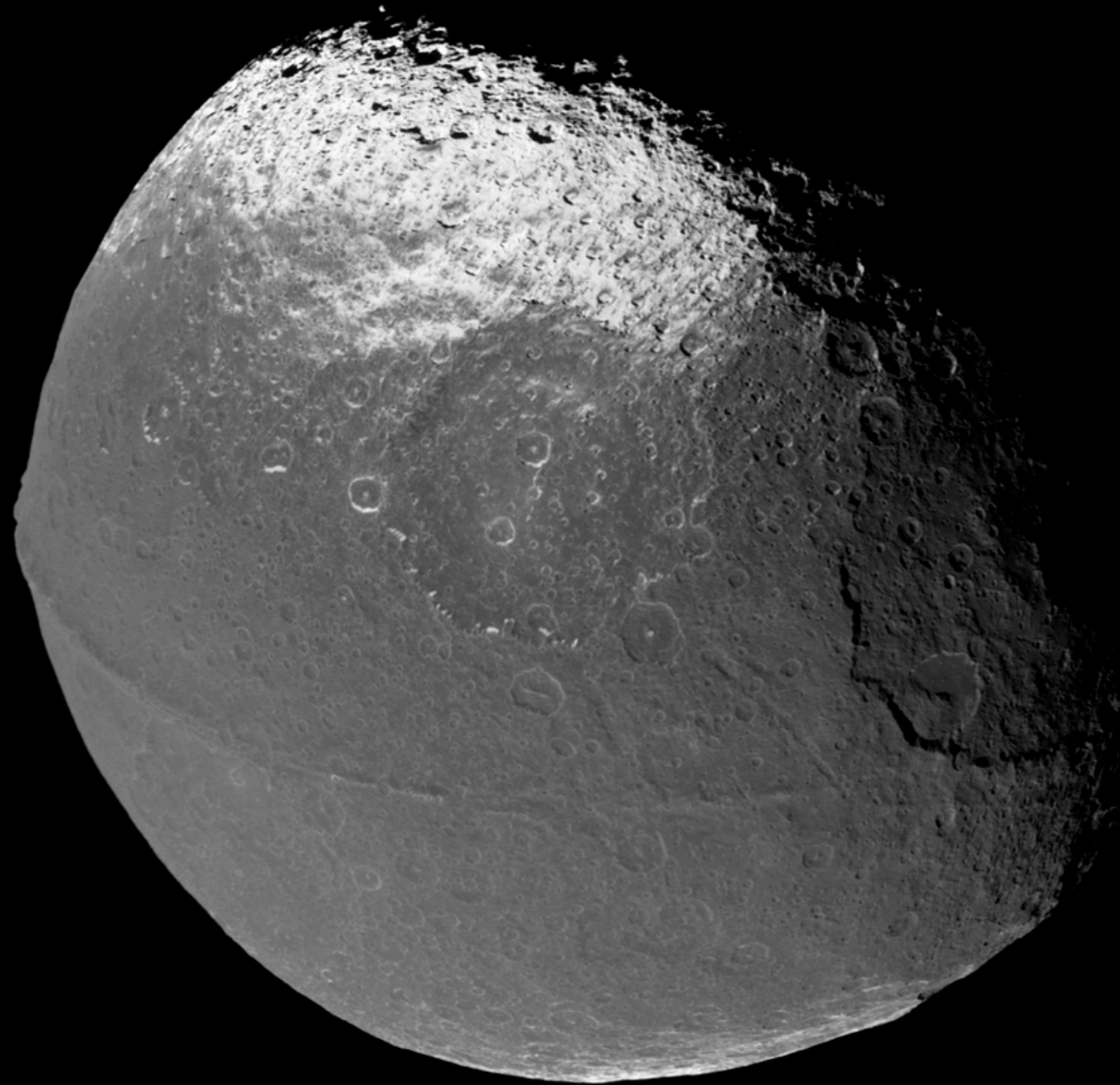
Mystery II: Anomalous flattening

$c - a = 35 \text{ km}$
(Earth: 21 km)
 $\sim T = 16 \text{ h} ?$

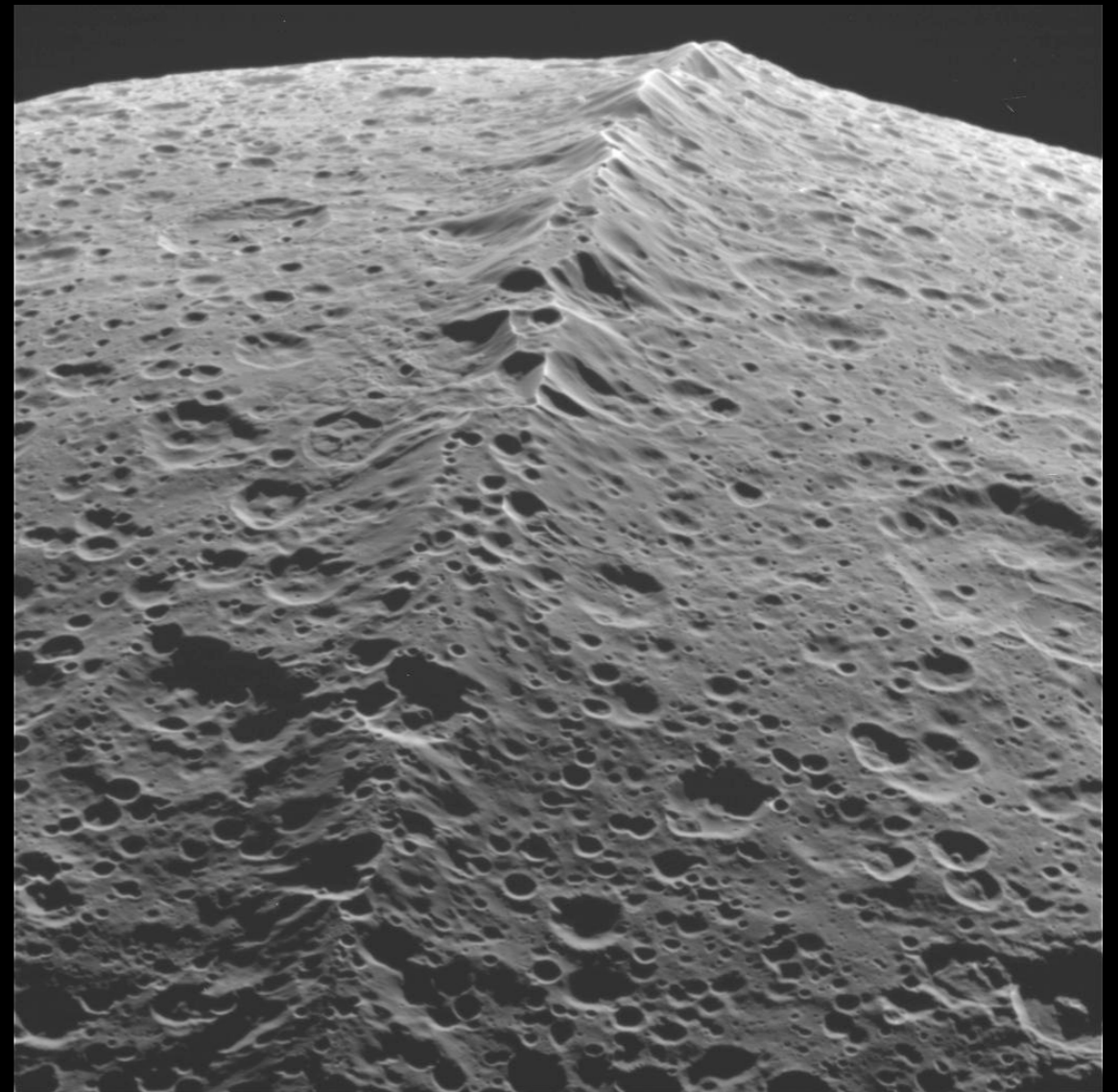


Synchronous orbit
 $T = 79 \text{ days}$

Mystery III: Equatorial ridge on Iapetus



Cassini flyby 2004



Cassini flyby 2007

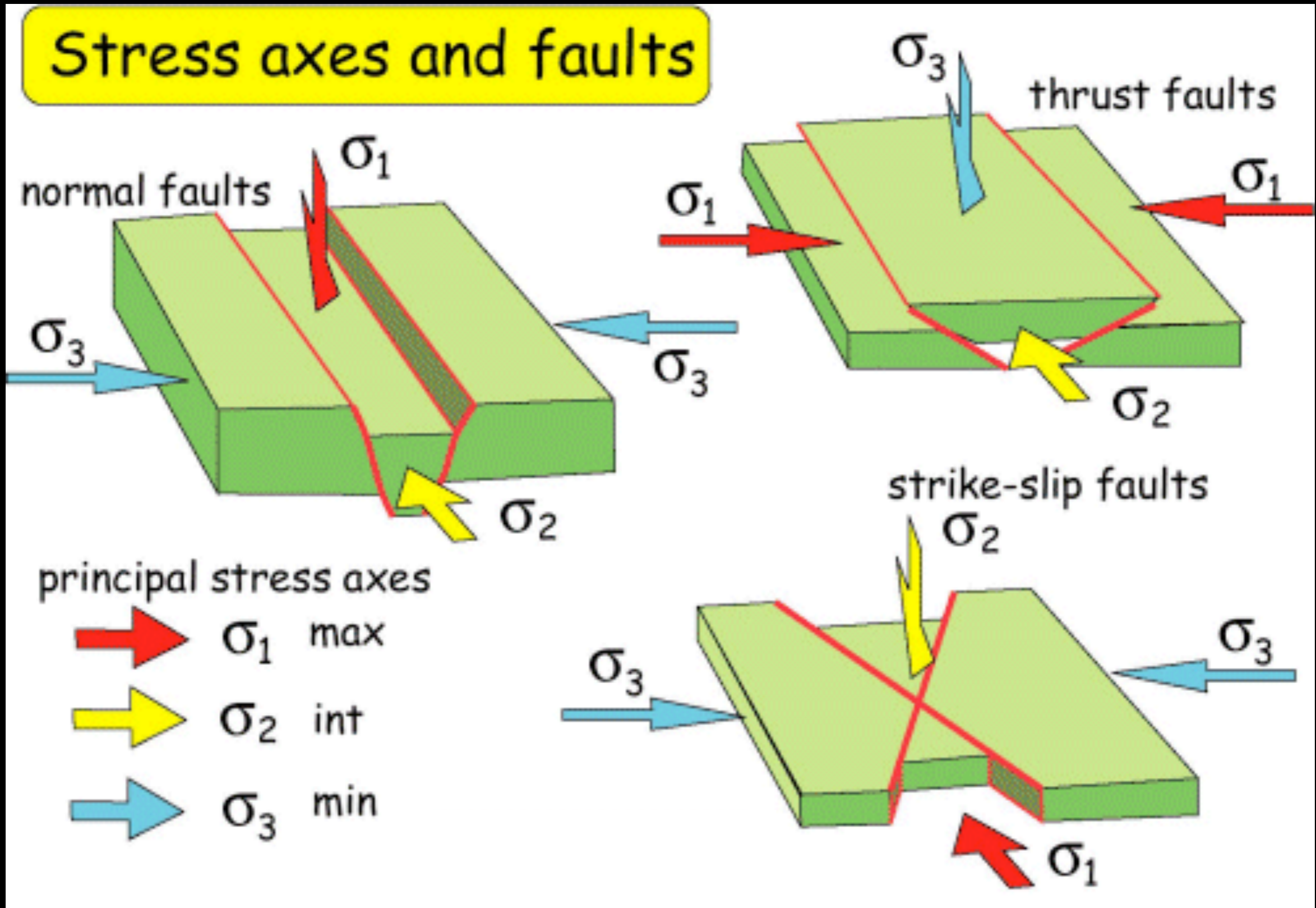
Modeling global tectonics

Global deformation of lithosphere → axial pattern ?

	Lithospheric thickness	Deformation	Pattern
Old model	constant	despinning	wrong
My model	thinner at equator	contraction, expansion	right

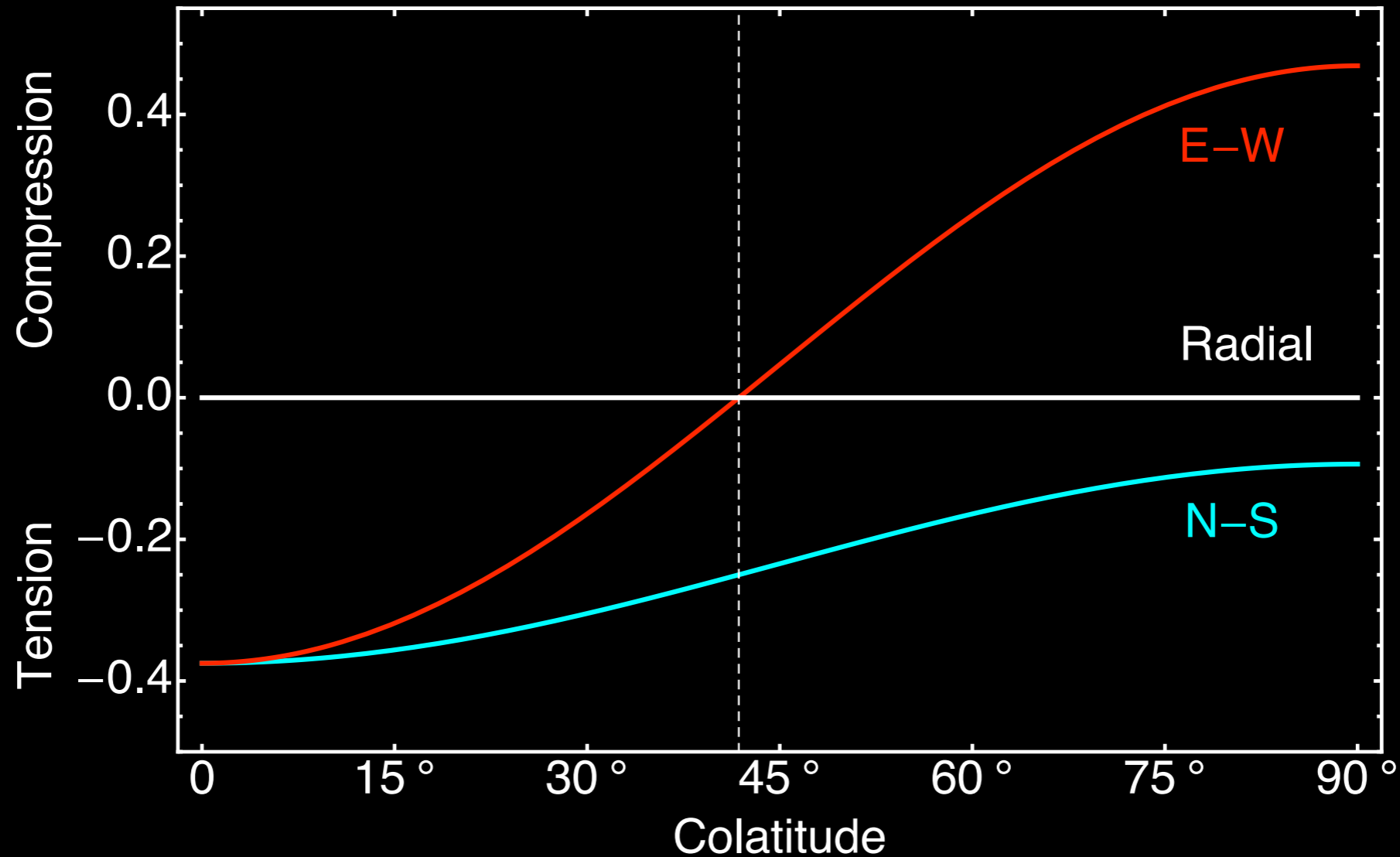
Model: Thin elastic shell with variable thickness

Stress and faulting



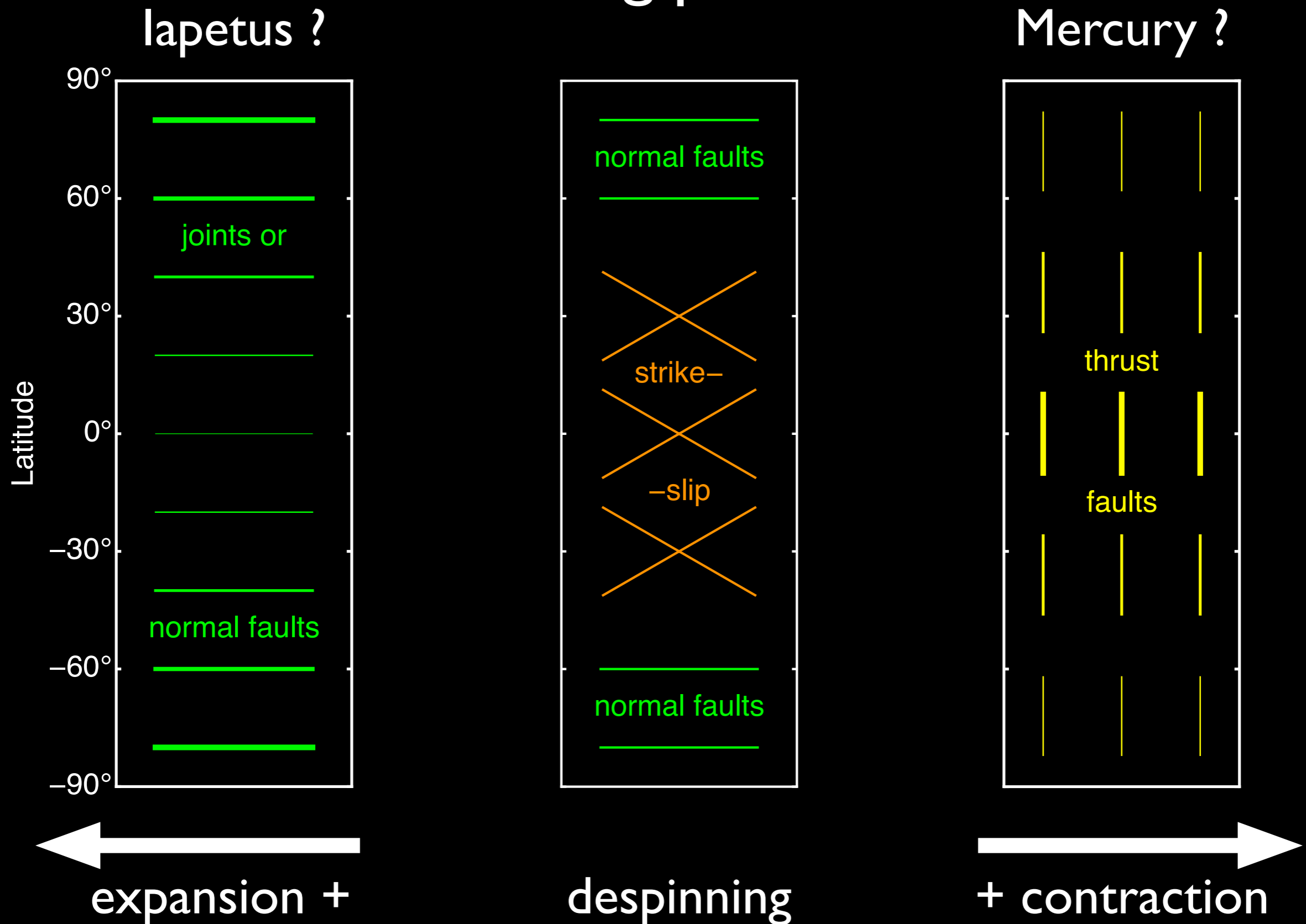
Lithosphere of constant thickness: despinning

Lithospheric stress



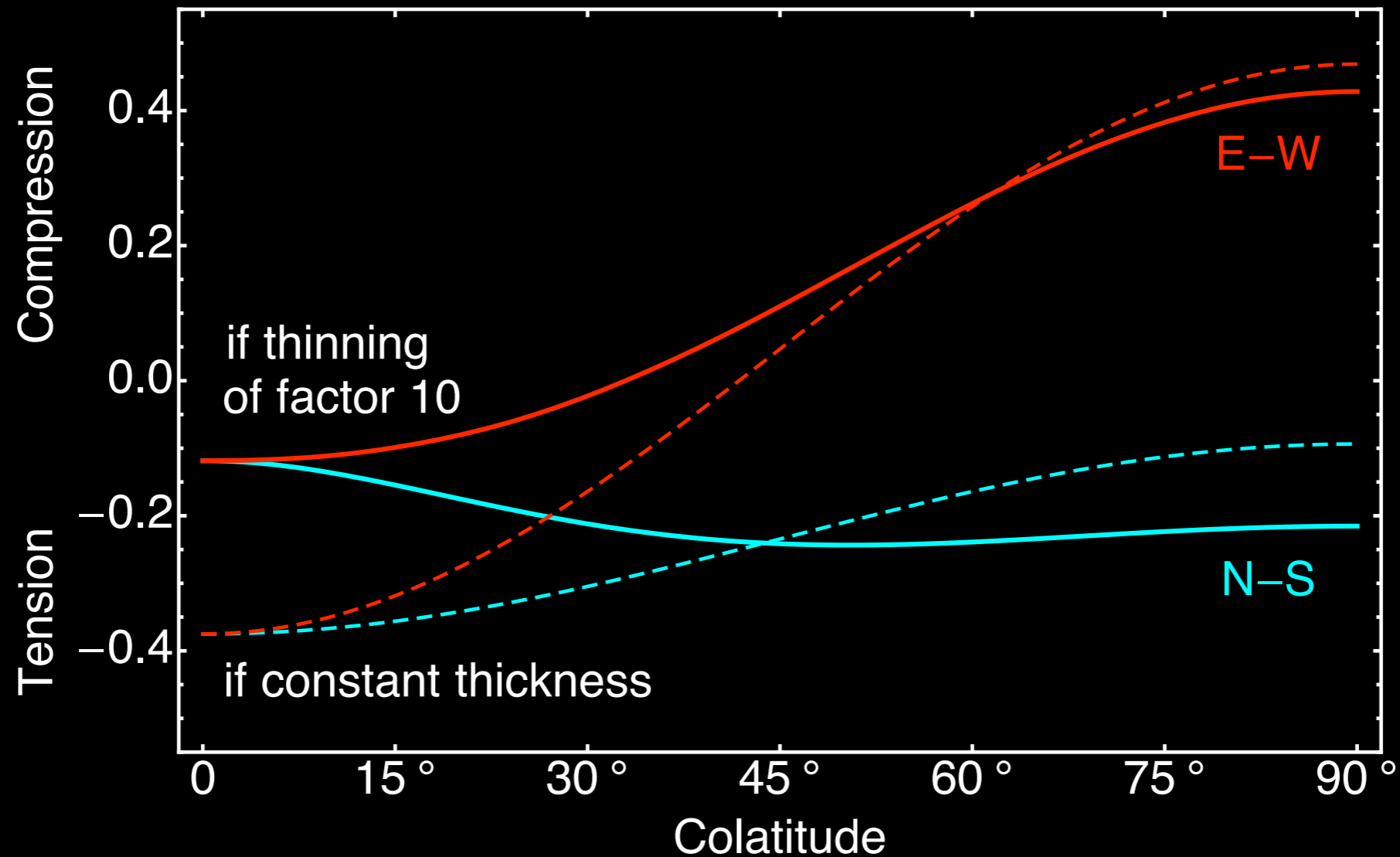
- E-W stress more compressive than N-S stress
- maximum compression at the equator
- maximum extension at the poles

Lithosphere of constant thickness: faulting patterns



Lithosphere thinner at equator: despinning

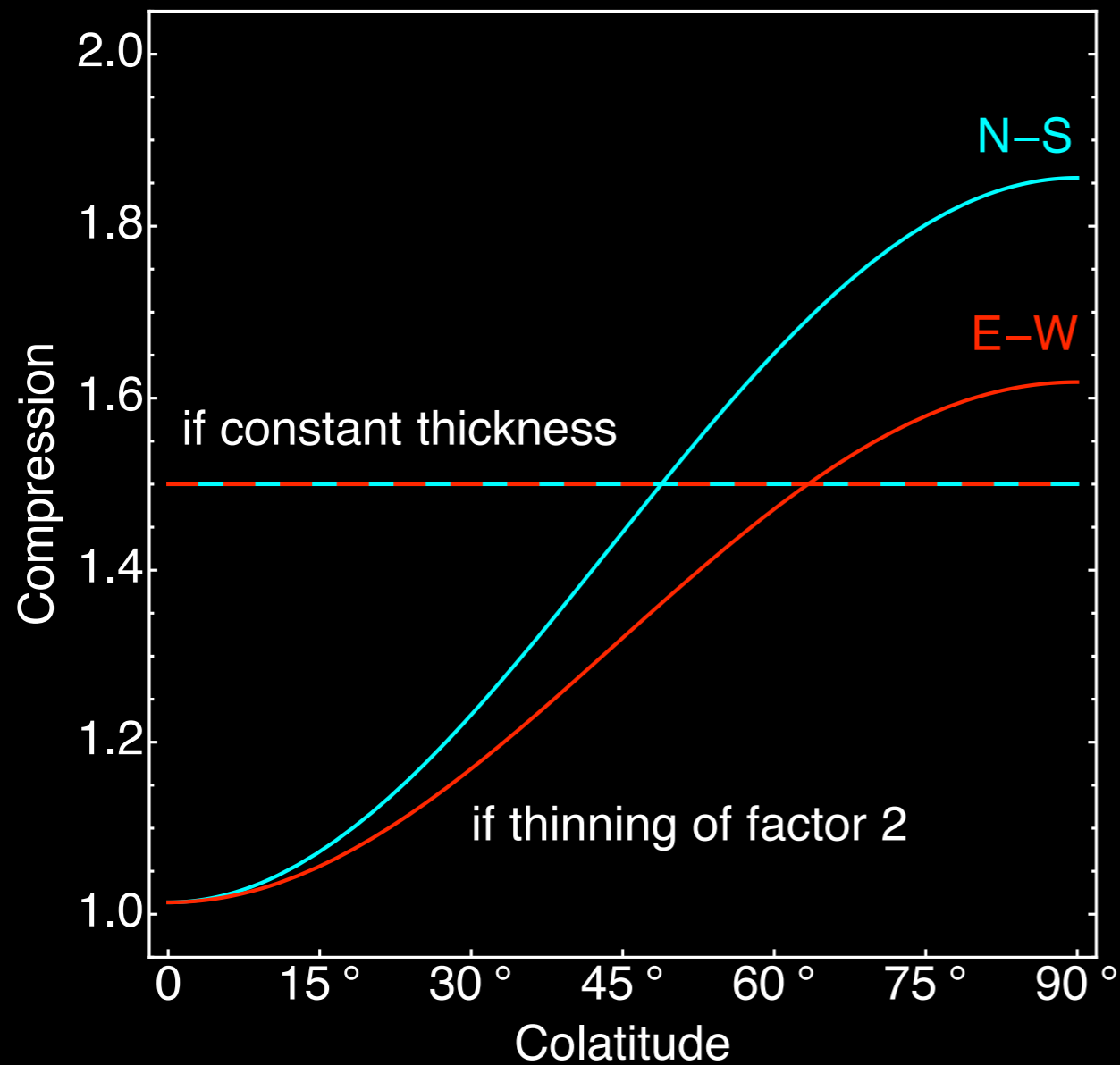
Lithospheric stress



- E-W stress still more compressive than N-S stress
- faulting pattern weakly affected

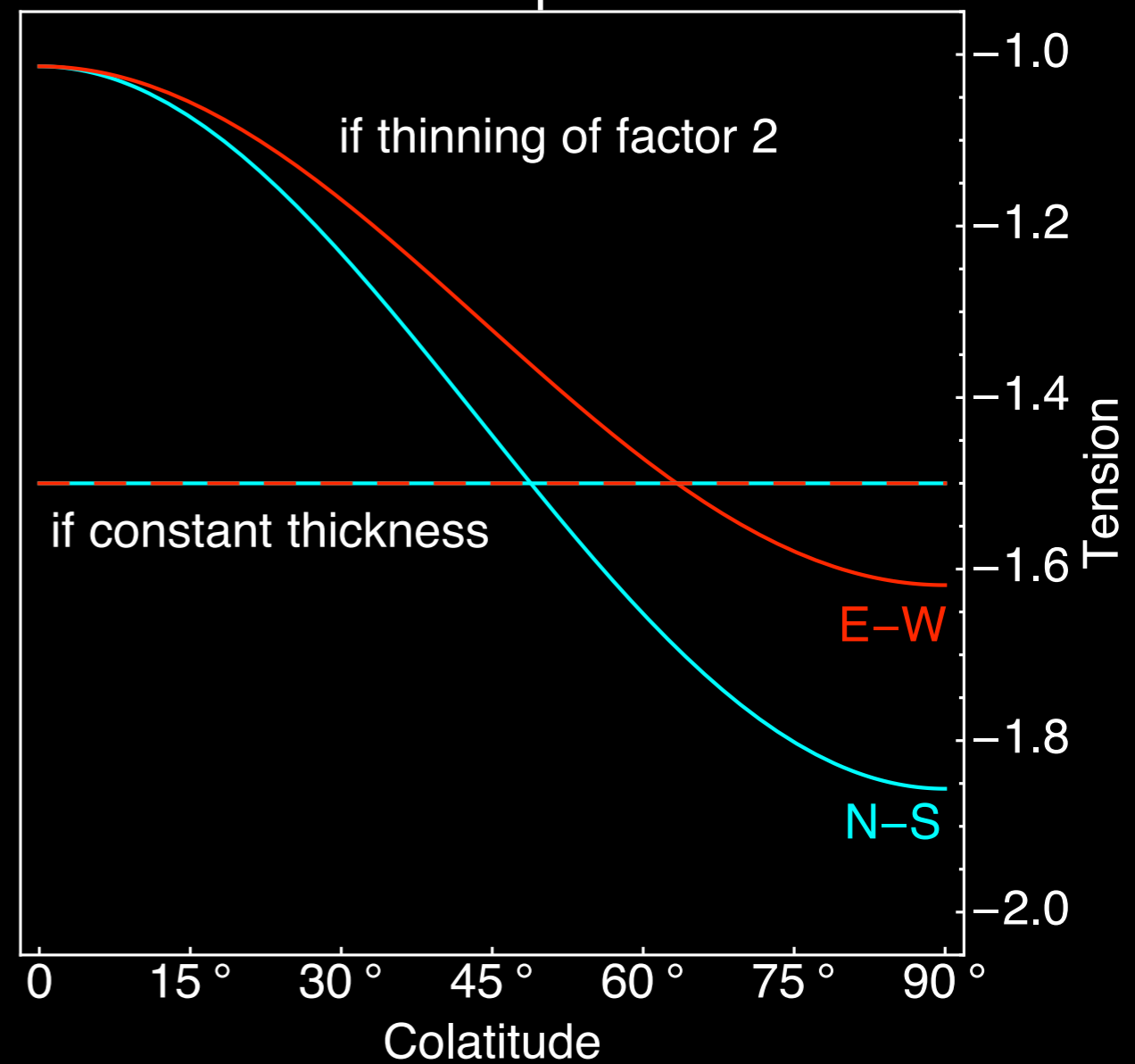
Lithosphere thinner at equator: contraction or expansion

Stress if contraction



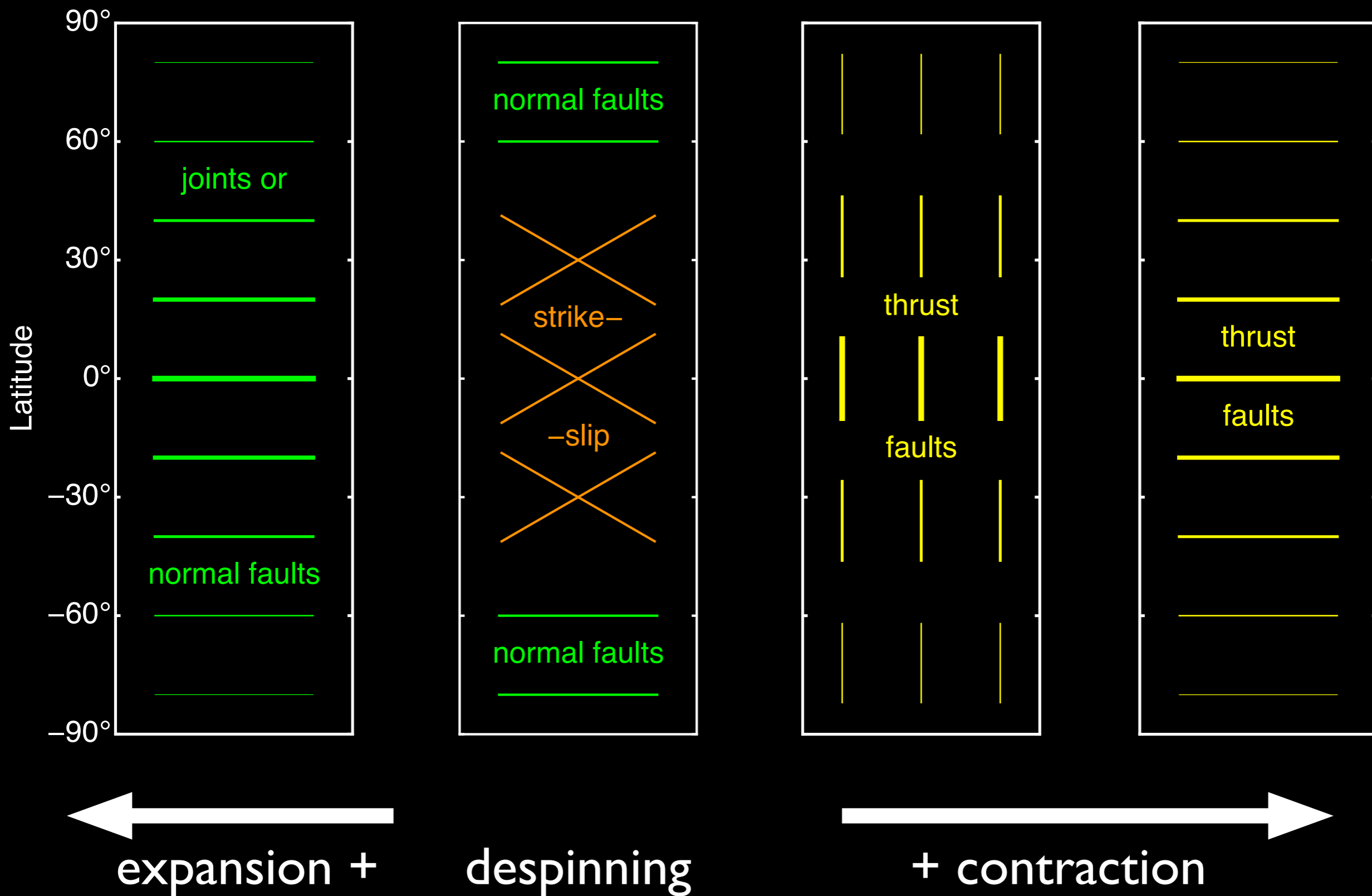
- N-S stress $>$ E-W stress
- maximum compression at equator

Stress if expansion



- N-S stress $<$ E-W stress
- maximum extension at equator

Lithosphere thinner at equator: faulting patterns



Formation of Iapetus' ridge

Scenarios:

A. Contraction, despinning later ($T_0 \sim 16h$)

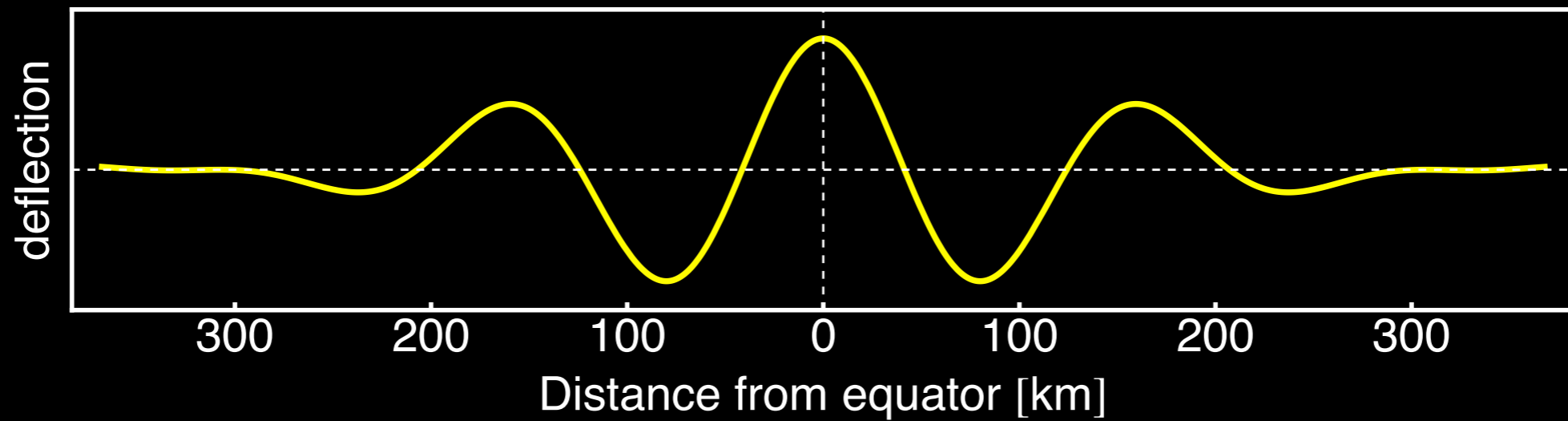
B. Contraction during despinning ($T_0 \sim 16h$):

E-W thrust faults if $\Delta R > 13 \text{ km}$

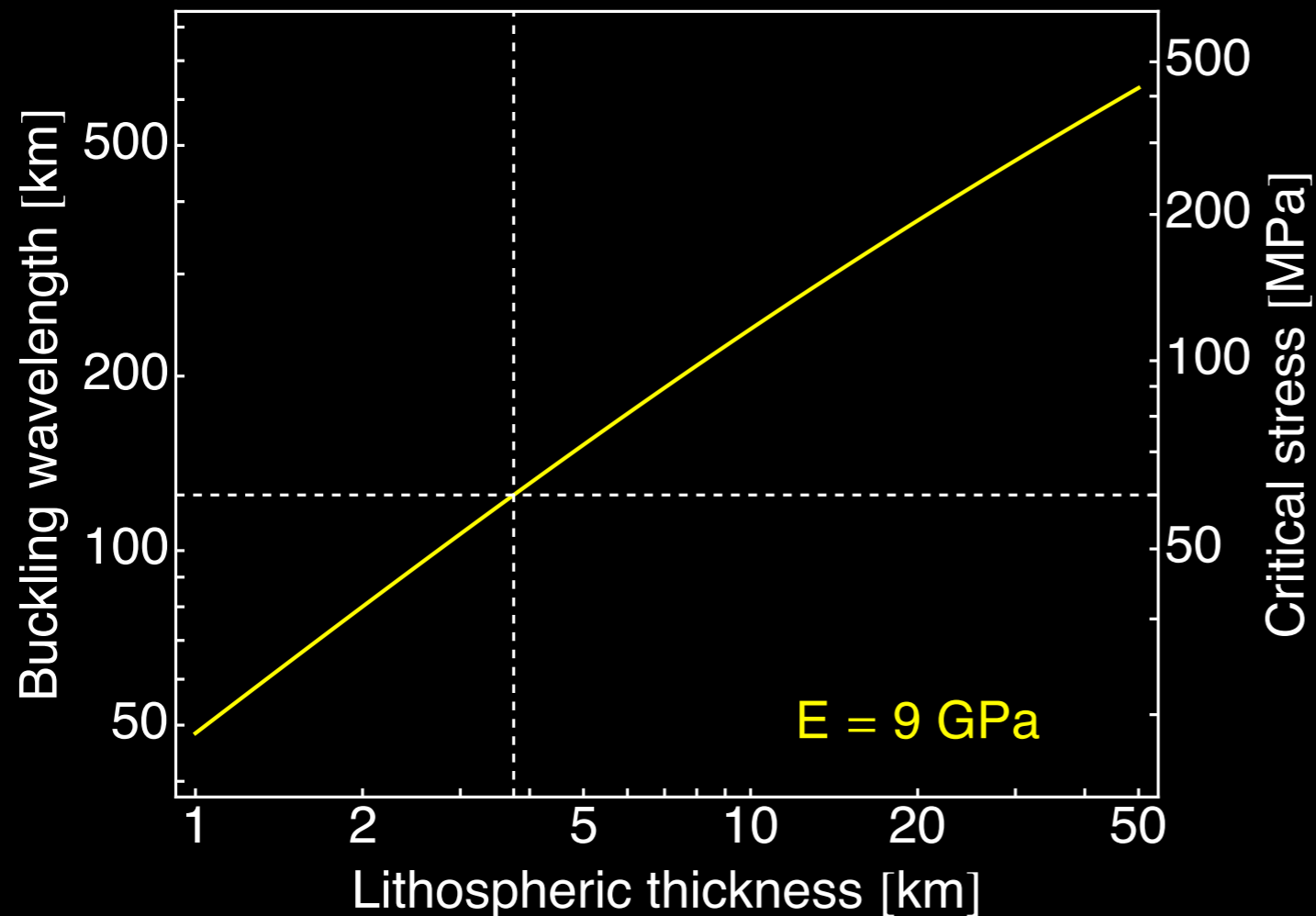
C. Expansion and despinning:

E-W joint \Rightarrow dike

Elastic buckling with variable thickness



Critical wavelength and stress

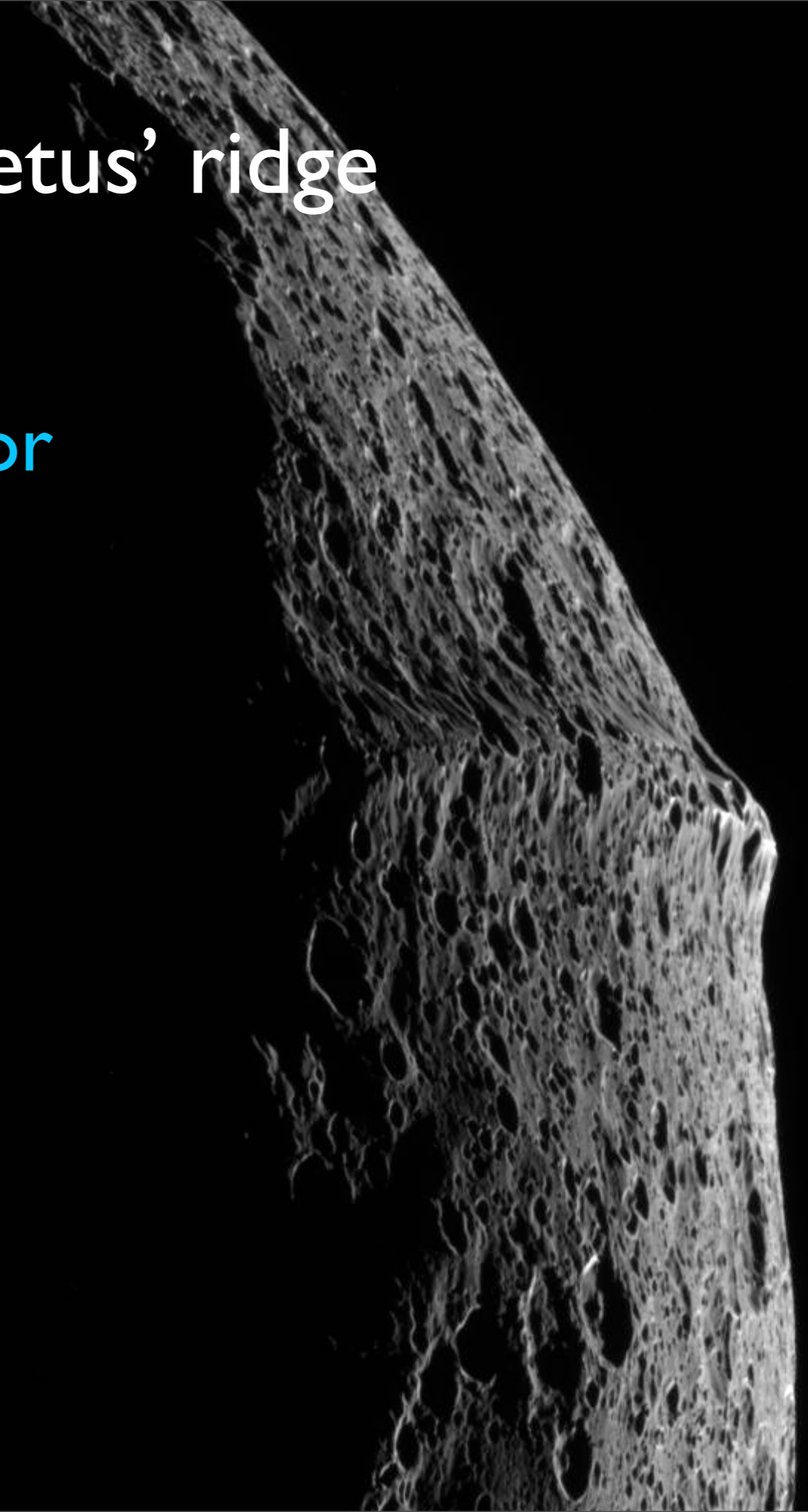


Yield stress ice ~
a few MPa

if $E \div 1000$
then
stress $\div 10$
thickness $\times 10$

Conclusions on Iapetus' ridge

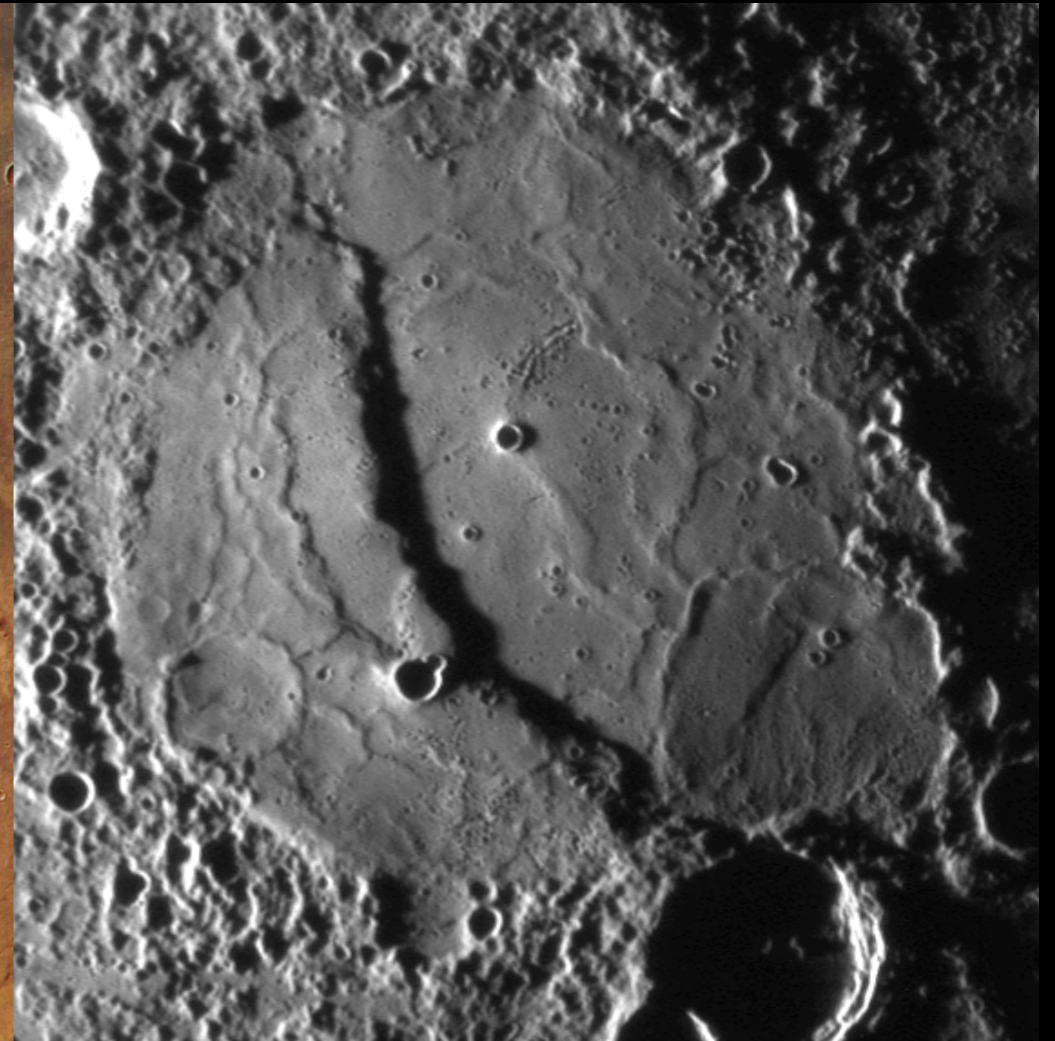
1. lithosphere thinner at equator
+ contraction/expansion
2. no despinning tectonics:
initial period $\sim 16\text{h}$?
3. no elastic buckling:
critical stress too high



Other planets

Mars: wrinkle ridges

Mercury: lobate scarps



References:

Beuthe M., 2008. Thin elastic shells with variable thickness for lithospheric flexure of one-plate planets, Geophys. J. Int. 172, 817.

Beuthe M., 2010. East-west faults due to planetary contraction, Icarus (in press).

Left: Hesperia Planum imaged by HRSC on Mars Express (ESA/DLR/FU Berlin)

Right: MESSENGER, 2nd flyby (NASA/Johns Hopkins University Applied Physics Laboratory/Arizona State University/Carnegie Institution of Washington)

Acknowledgments

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