



# Hydrogen halides at Comet 67P/Churyumov-Gerasimenko & what they tell us about molecular clouds

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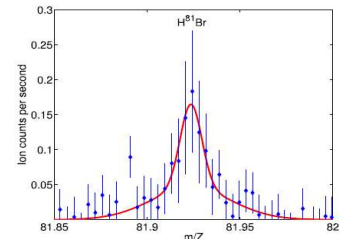
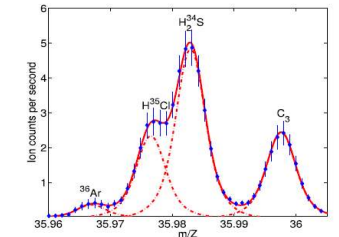
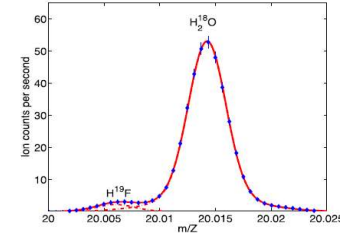
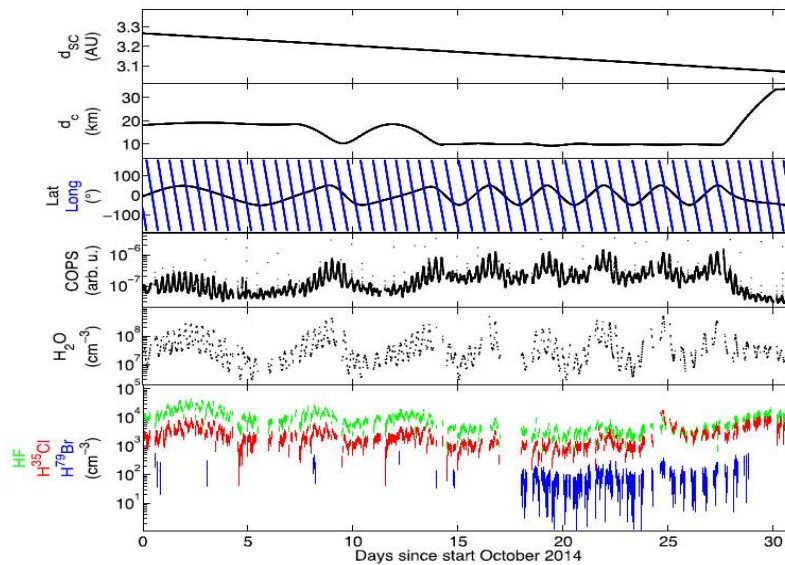
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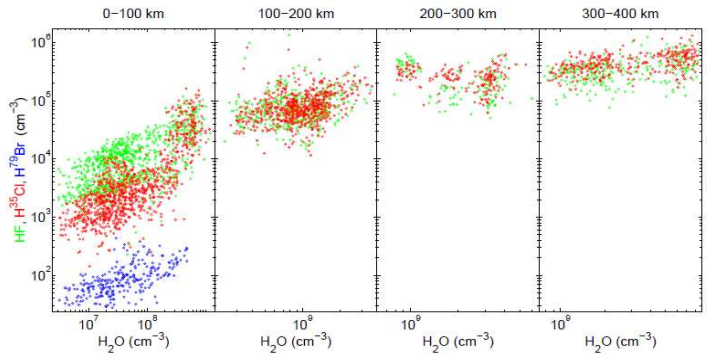


## Observations

Rosetta studied comet 67P/Churyumov-Gerasimenko up close in 2014-2016. With the ROSINA/DFMS mass spectrometer we analyzed the coma gas. We discovered halogens in the form of the **hydrogen halides** HF, HCl, and HBr.  $^{37}\text{Cl}/^{35}\text{Cl} = 0.29 \pm 0.02$  and  $^{79}\text{Br}/^{81}\text{Br} = 0.92 \pm 0.08$  are **typical solar system isotopic ratios**.



The halogen-to-oxygen ratio increases with cometocentric distance. Far from the comet the amplitude of the daily variations is less for the hydrogen halides than for water. This points to a distributed source: **hydrogen halide-enriched gas progressively sublimates from the icy dust grains** that are released by the comet nucleus.

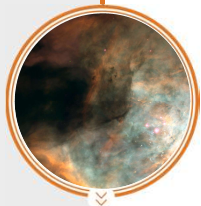


## Interpretation

- 67P contains **icy dust grains** that formed in a **molecular cloud (MC)** from which the solar system formed and were **incorporated into the comet nucleus with little modification** (beyond the snow line, no melting).
- The halides in the MC gas rapidly adsorb onto water ice deposited on grains (strong polar bonds), leading to a **depletion of halides in the MC gas phase** (as observed) and **halogen-enriched inner ice layers on the dust**.
- These enriched layers sublime relatively late in the coma since the halogen-poor ice has to sublime first, thus leading to the **distributed source**.

## References

- Dhooghe, F. et al.: Halogens as tracers of protosolar nebula material in comet 67P/Churyumov-Gerasimenko, MNRAS, 472, Issue 2, 1336, doi 10.1093/mnras/stx1911, 2017.
- De Keyser, J. et al.: Evidence for distributed gas sources of hydrogen halides in the coma of comet 67P/Churyumov-Gerasimenko, MNRAS, 469, Issue Suppl\_2, S695, doi 10.1093/mnras/stx2725, 2017.



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