



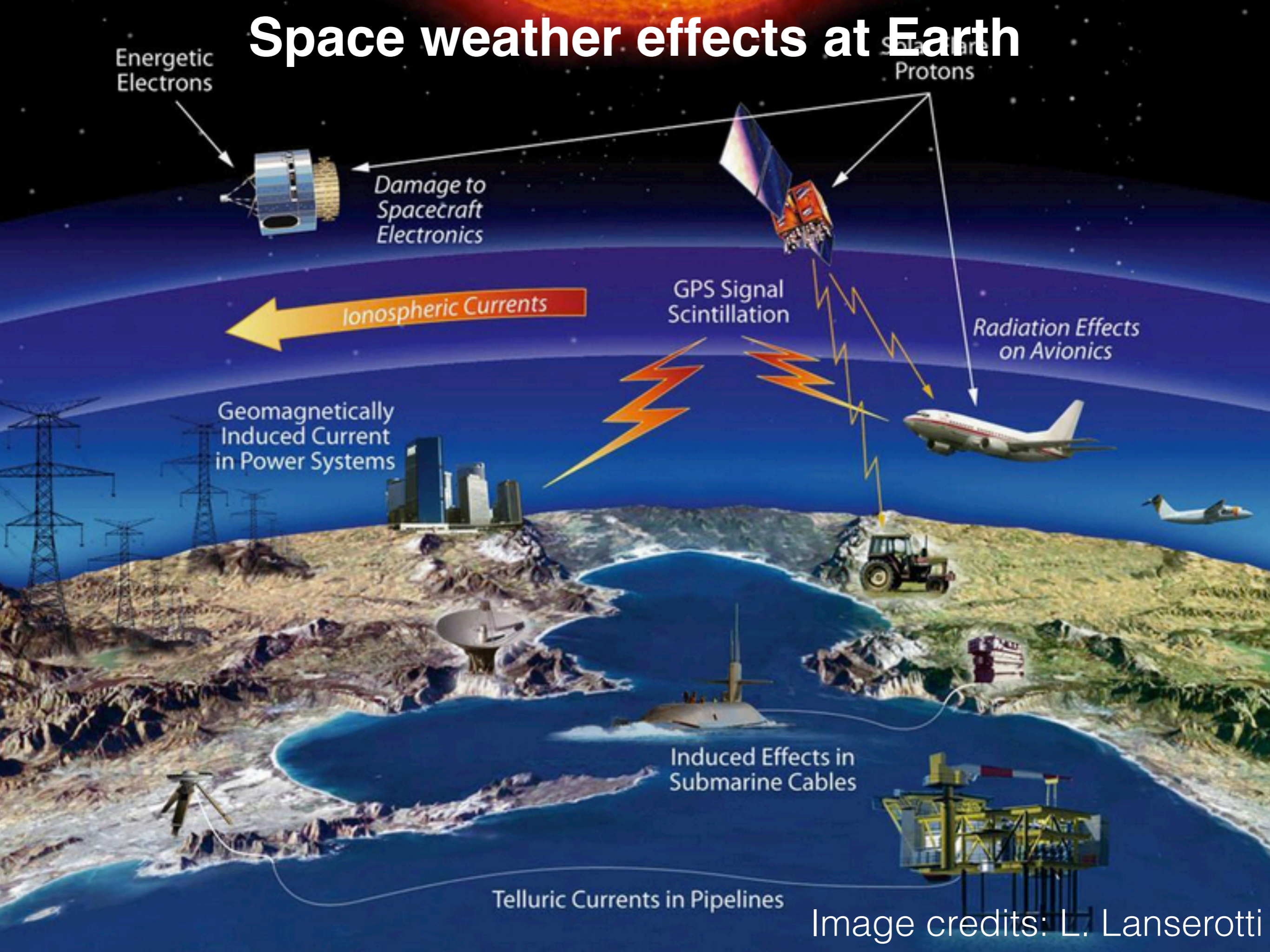
Magnetic reconnection at the terrestrial magnetopause: a simulation approach

Maria Elena Innocenti, Alex Anthony Arokiaraj,
Emanuele Cazzola, Giovanni Lapenta

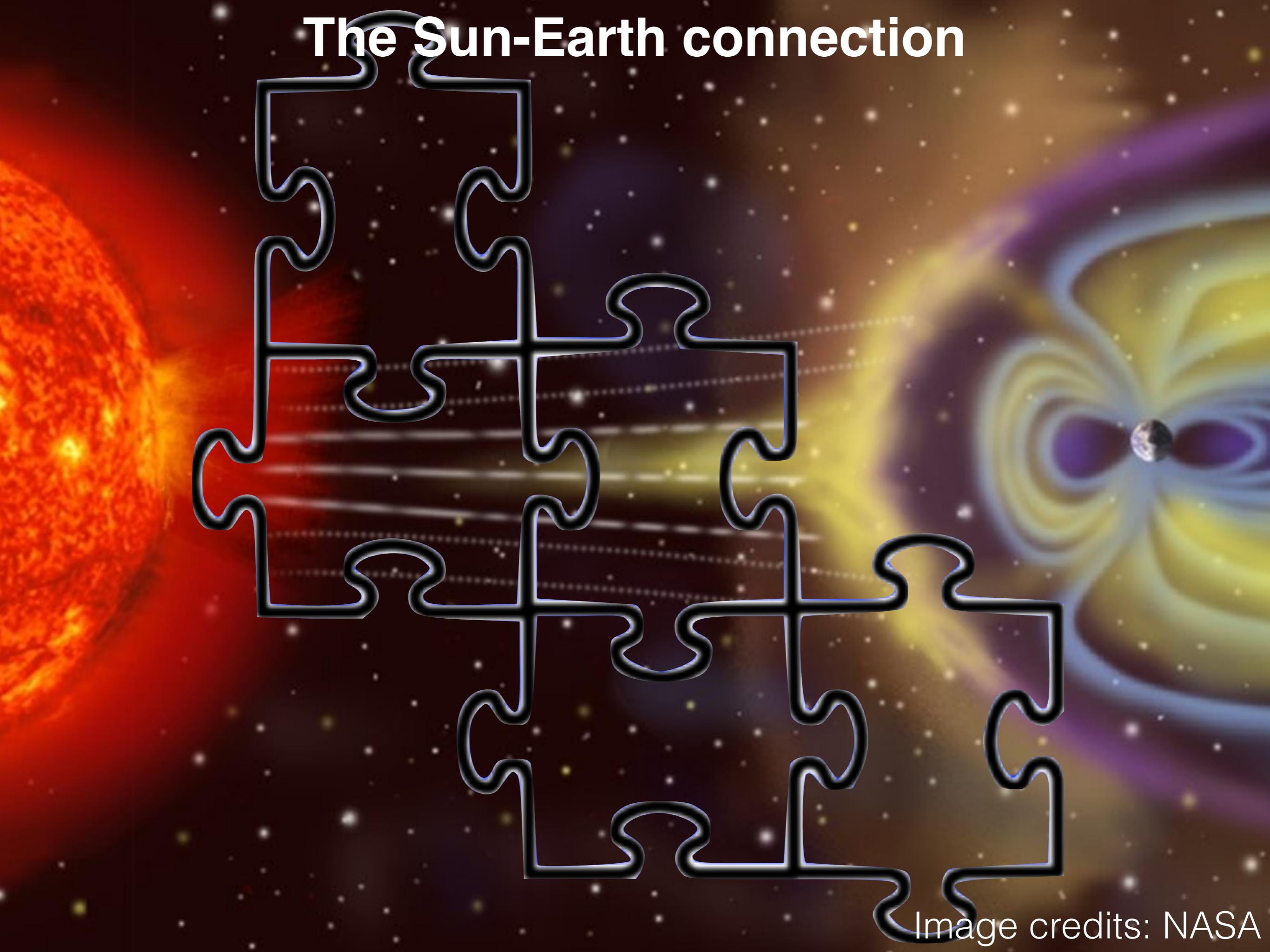
Center for mathematical Plasma Astrophysics, KULeuven, Belgium

Brussels, October 11th, 2016

Space weather effects at Earth

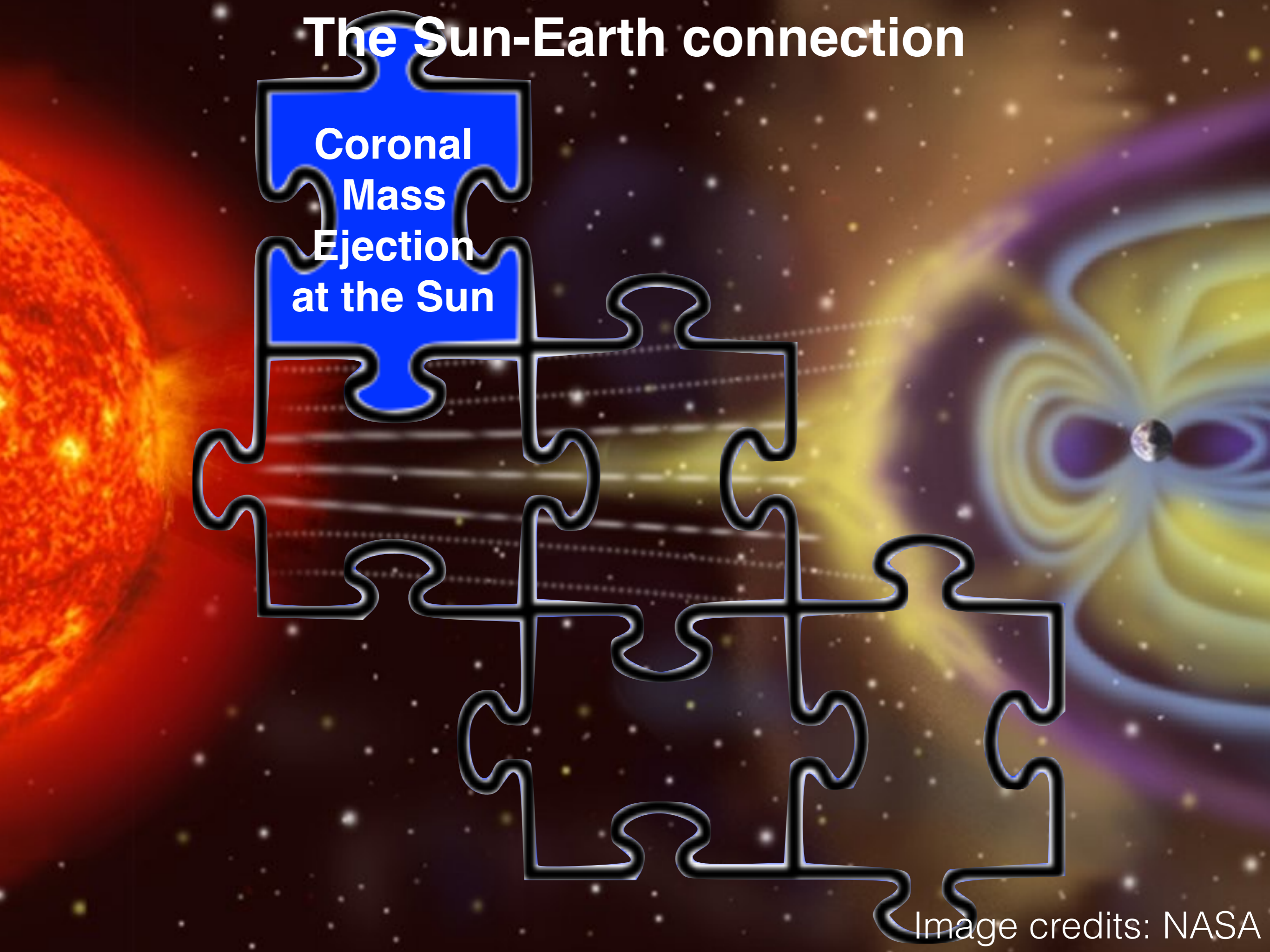


The Sun-Earth connection

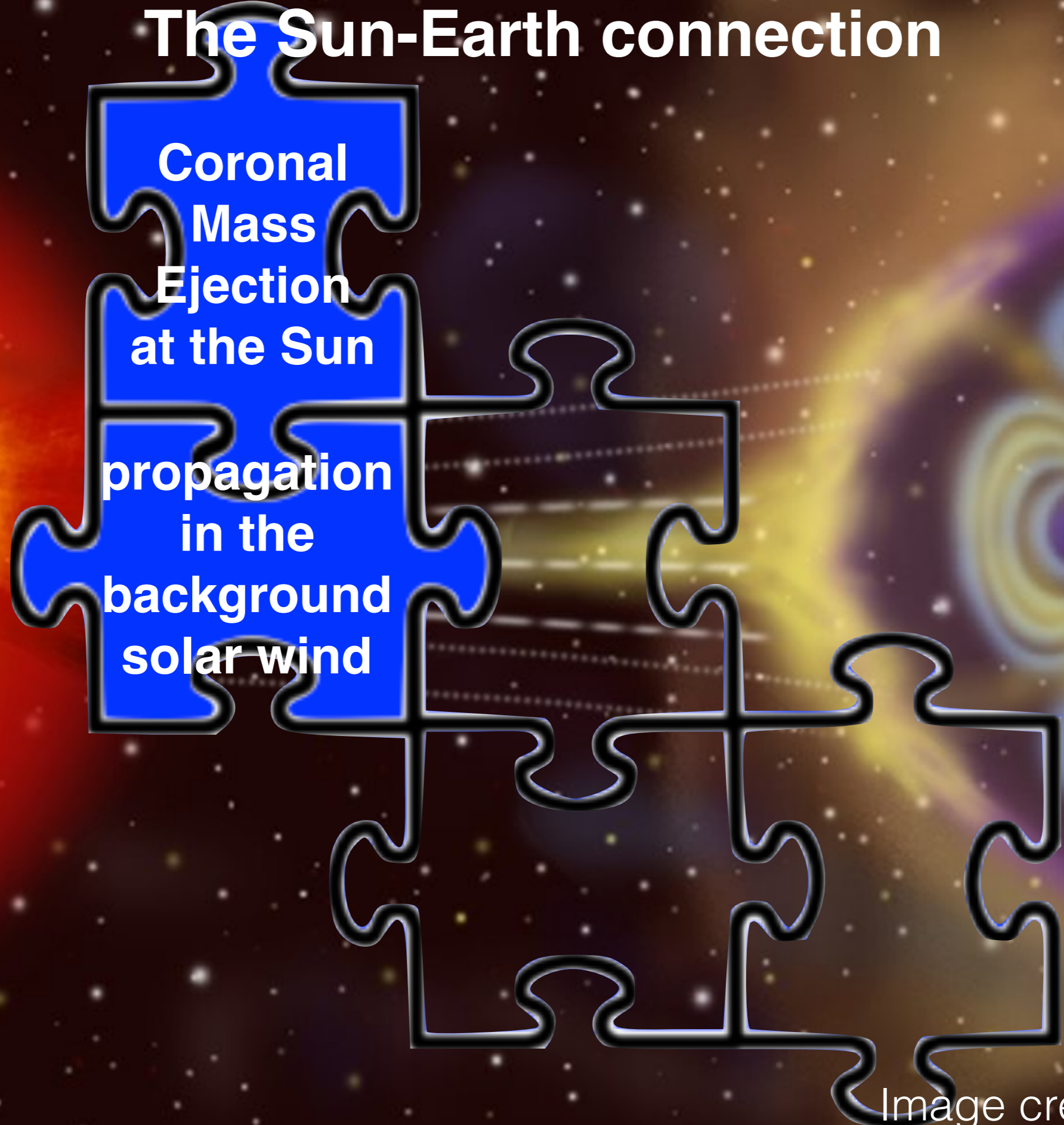


The Sun-Earth connection

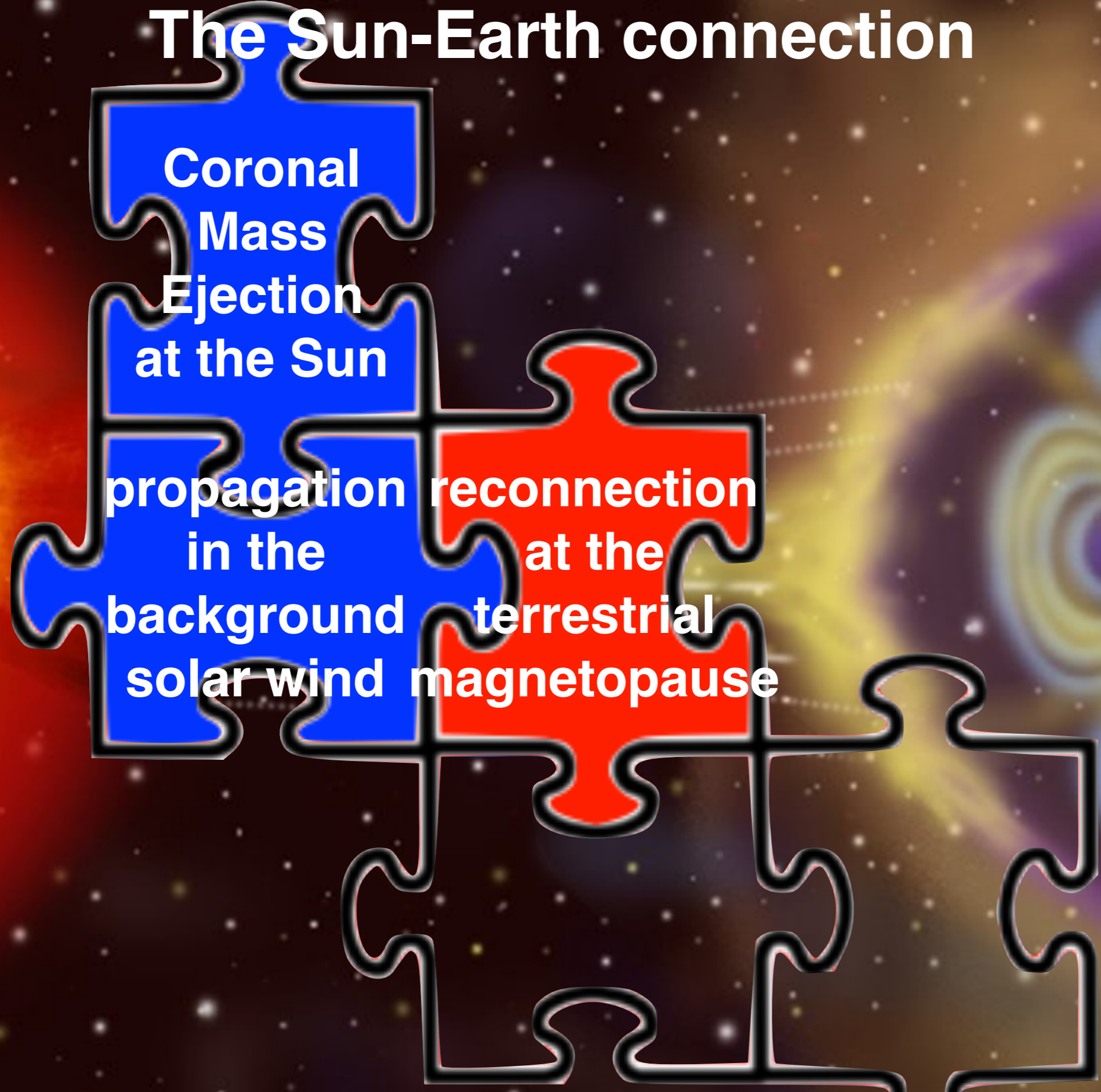
Coronal
Mass
Ejection
at the Sun



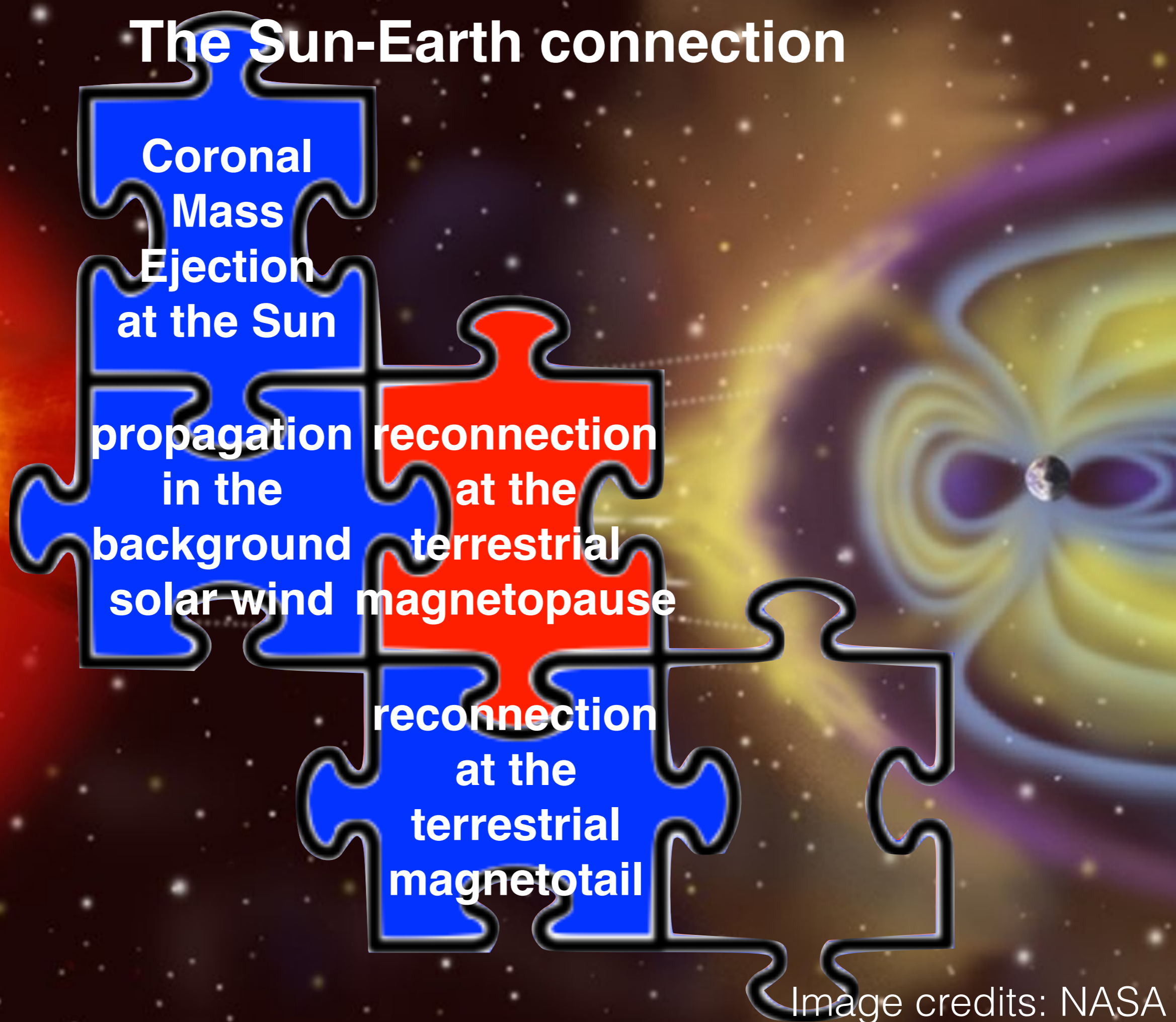
The Sun-Earth connection



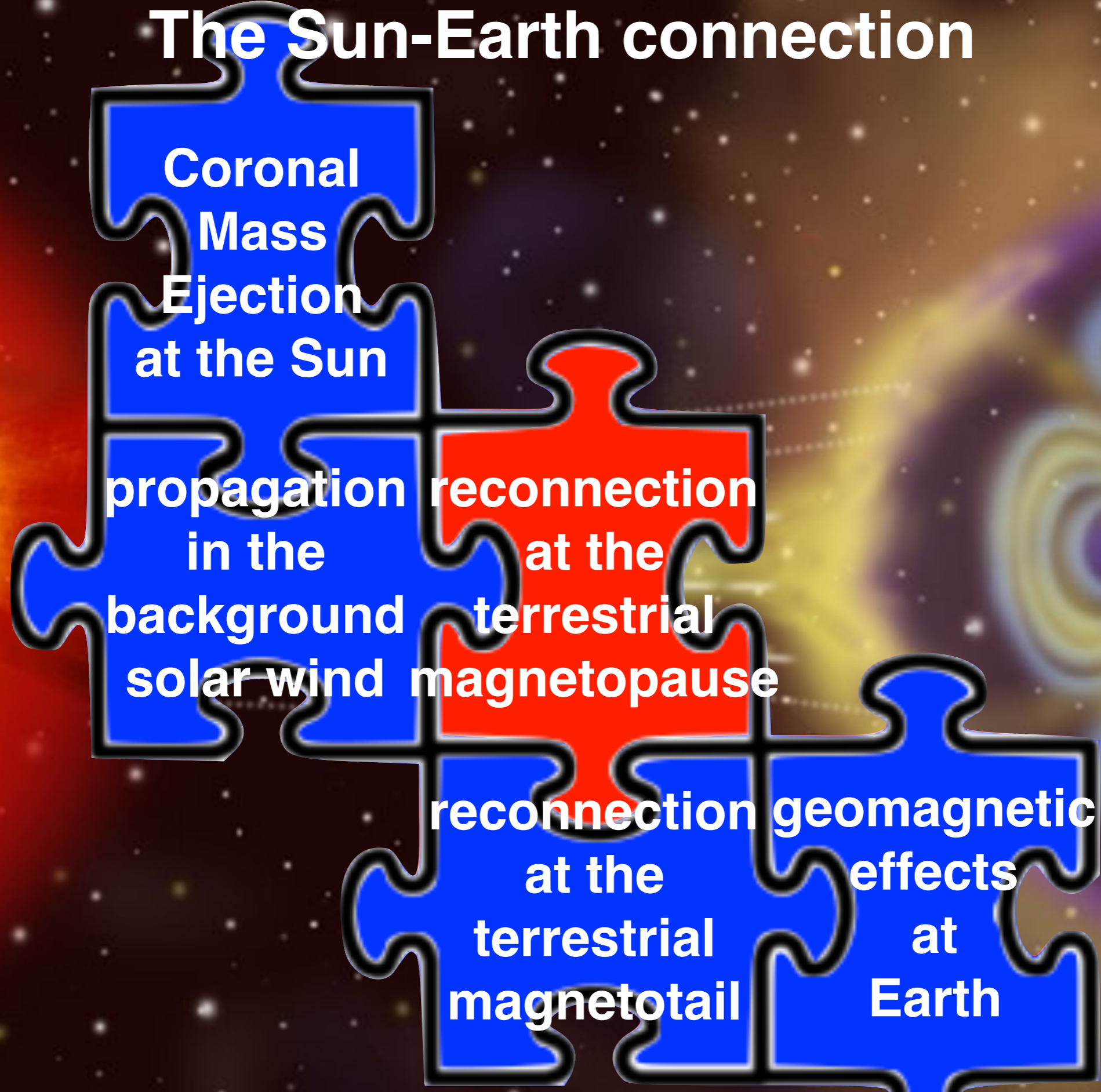
The Sun-Earth connection



The Sun-Earth connection

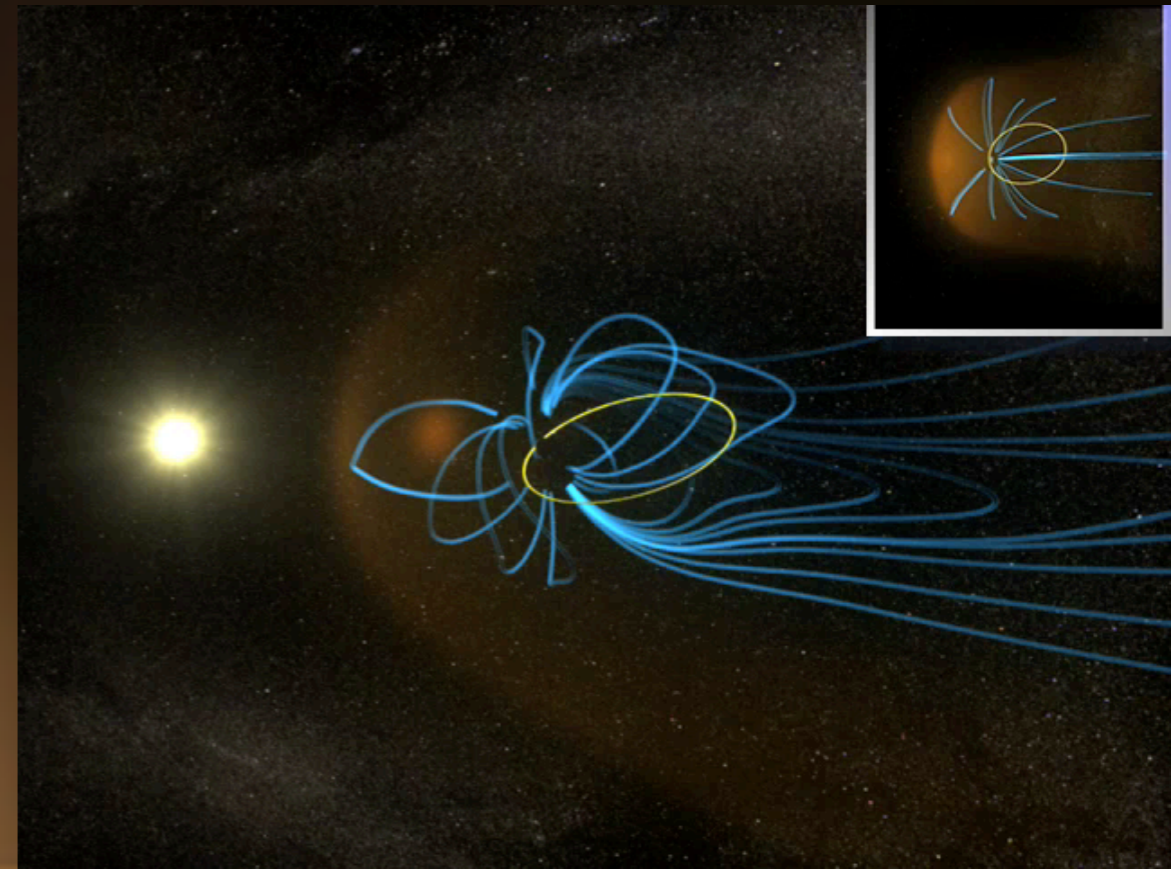


The Sun-Earth connection



The Magnetospheric Multiscale Mission (MMS)

- launched on March 13th, 2015
- four spacecrafts flying in formation with variable separation (d_i , d_e separation)
- will skim the magnetopause boundary during the first phase of operation
- aims at understanding the interplay of micro- and macro-scale in magnetopause reconnection



The terrestrial magnetopause



bow shock

magnetosheath

magnetopause

cusp

trapping region

neutral sheet

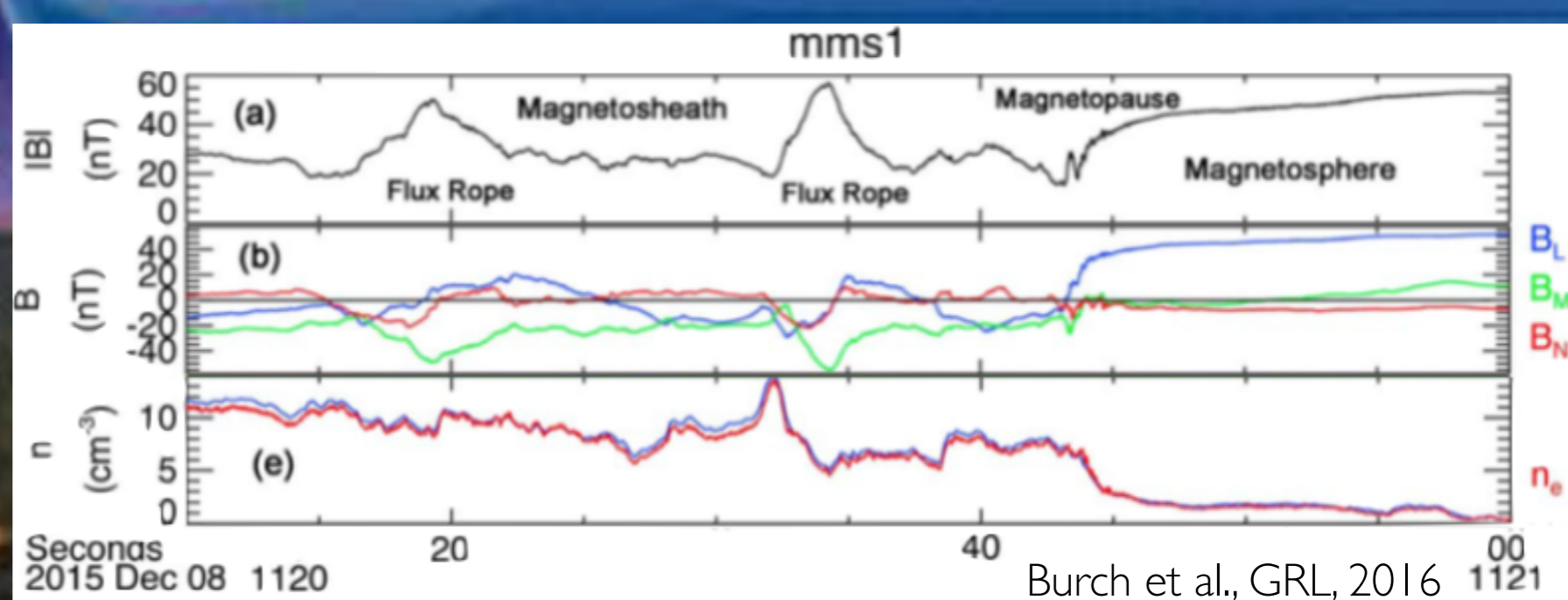


Image credits: NASA

Burch et al., GRL, 2016

The terrestrial magnetopause



bow shock

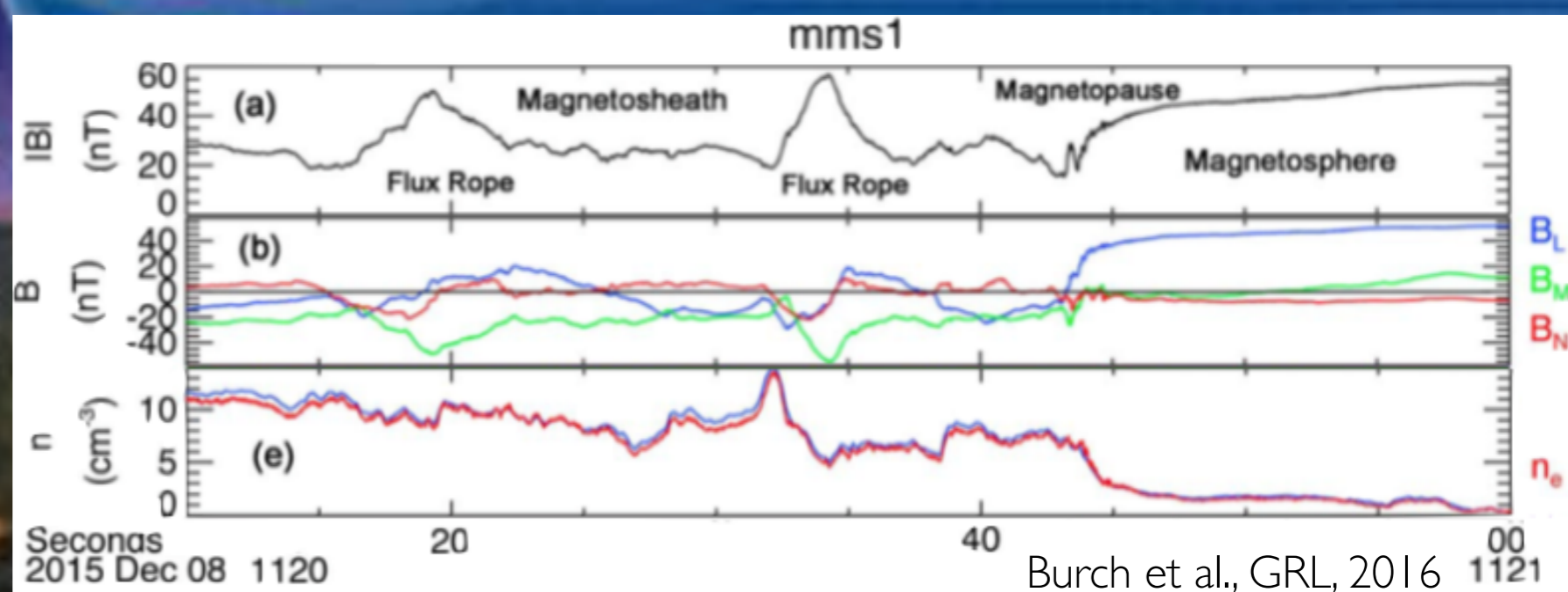
magnetosheath

magnetopause

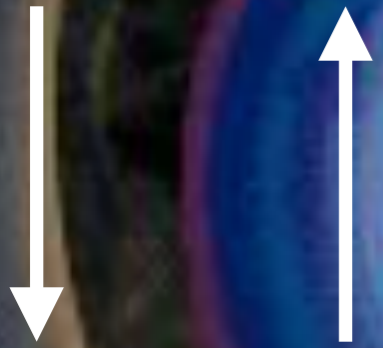
cusp

trapping region

neutral sheet



The terrestrial magnetopause



bow shock

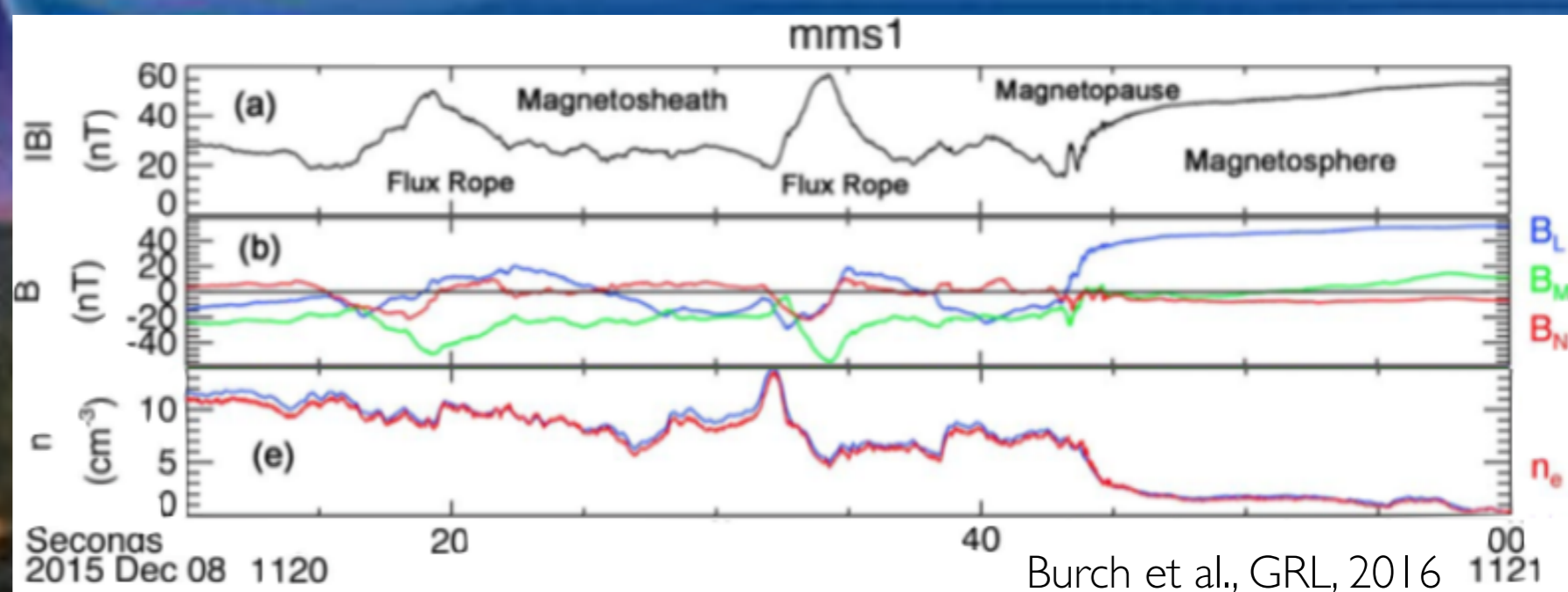
magnetosheath

magnetopause

cusp

trapping region

neutral sheet



The terrestrial magnetopause

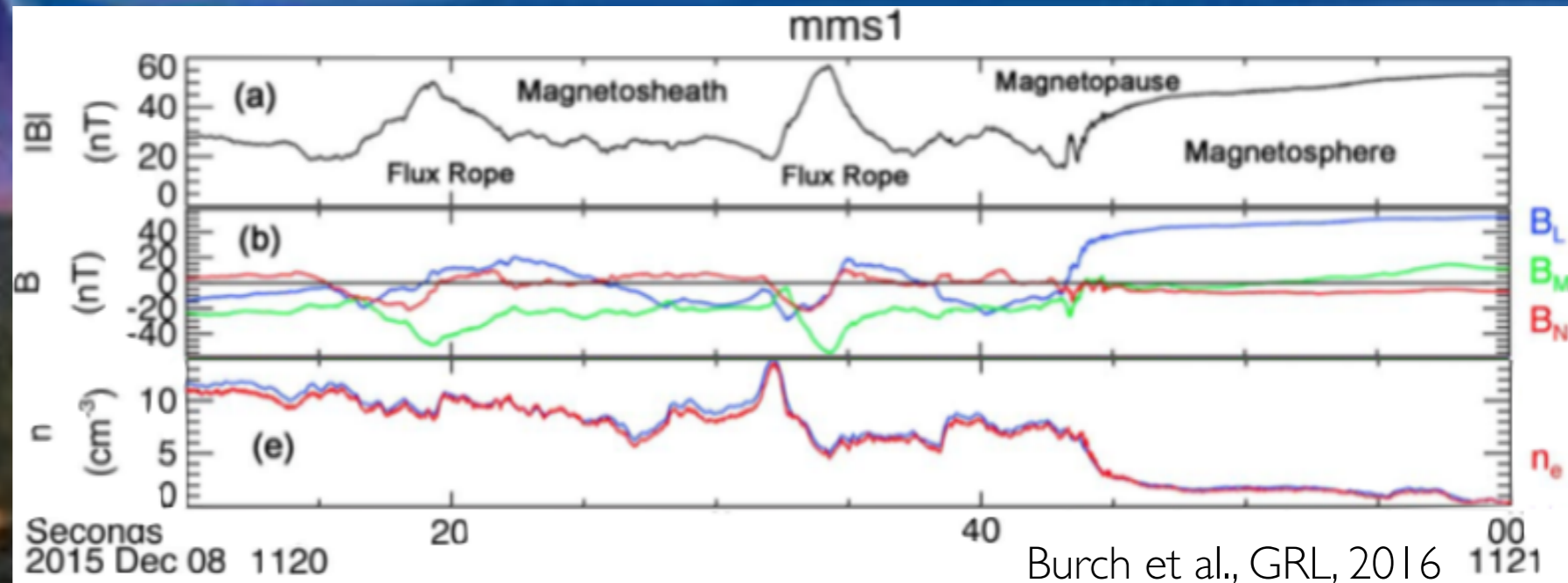
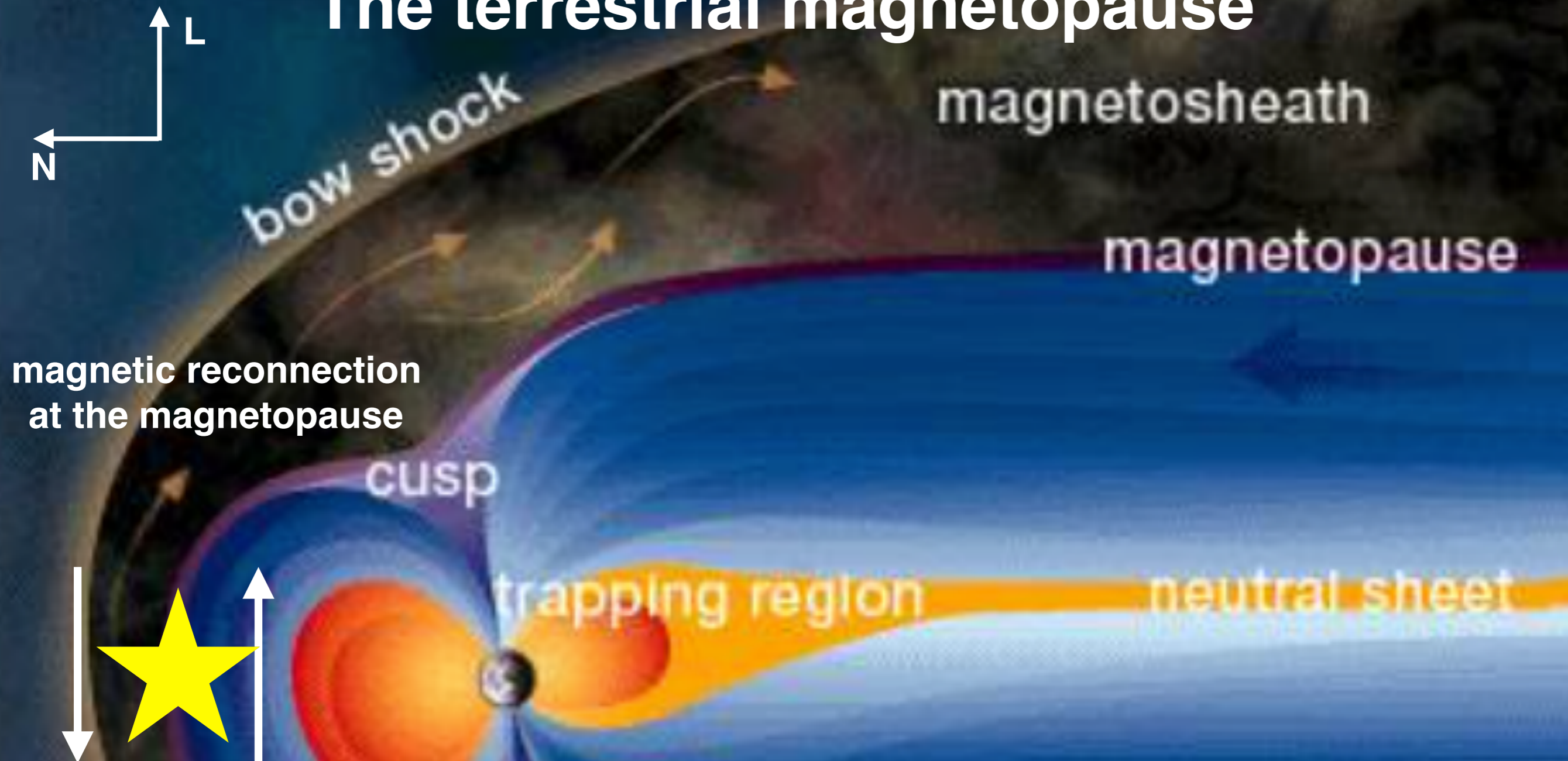
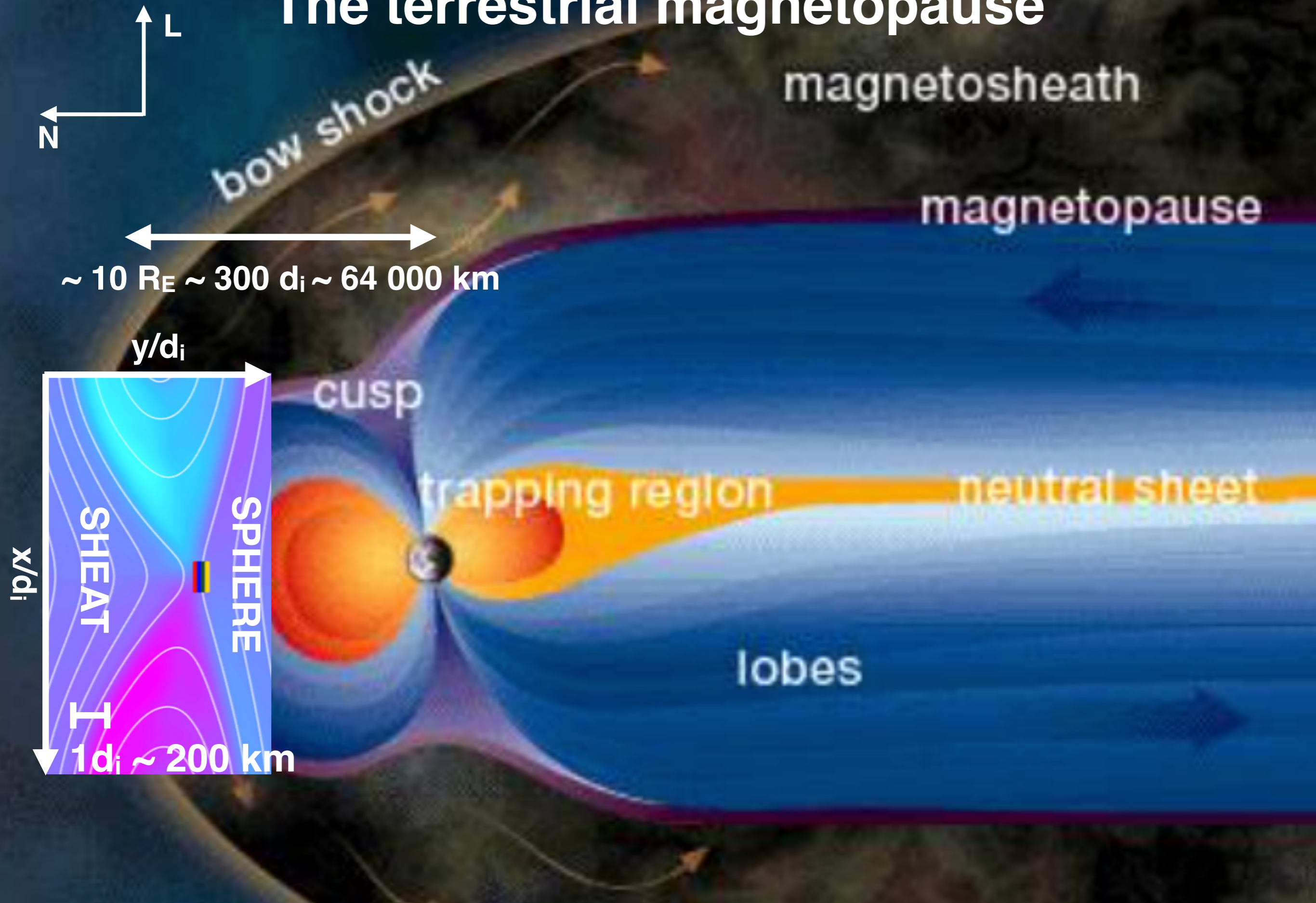


Image credits: NASA

Burch et al., GRL, 2016 1121

The terrestrial magnetopause



The terrestrial magnetopause



bow shock

magnetosheath

magnetopause

$\sim 10 R_E \sim 300$

SPHERE

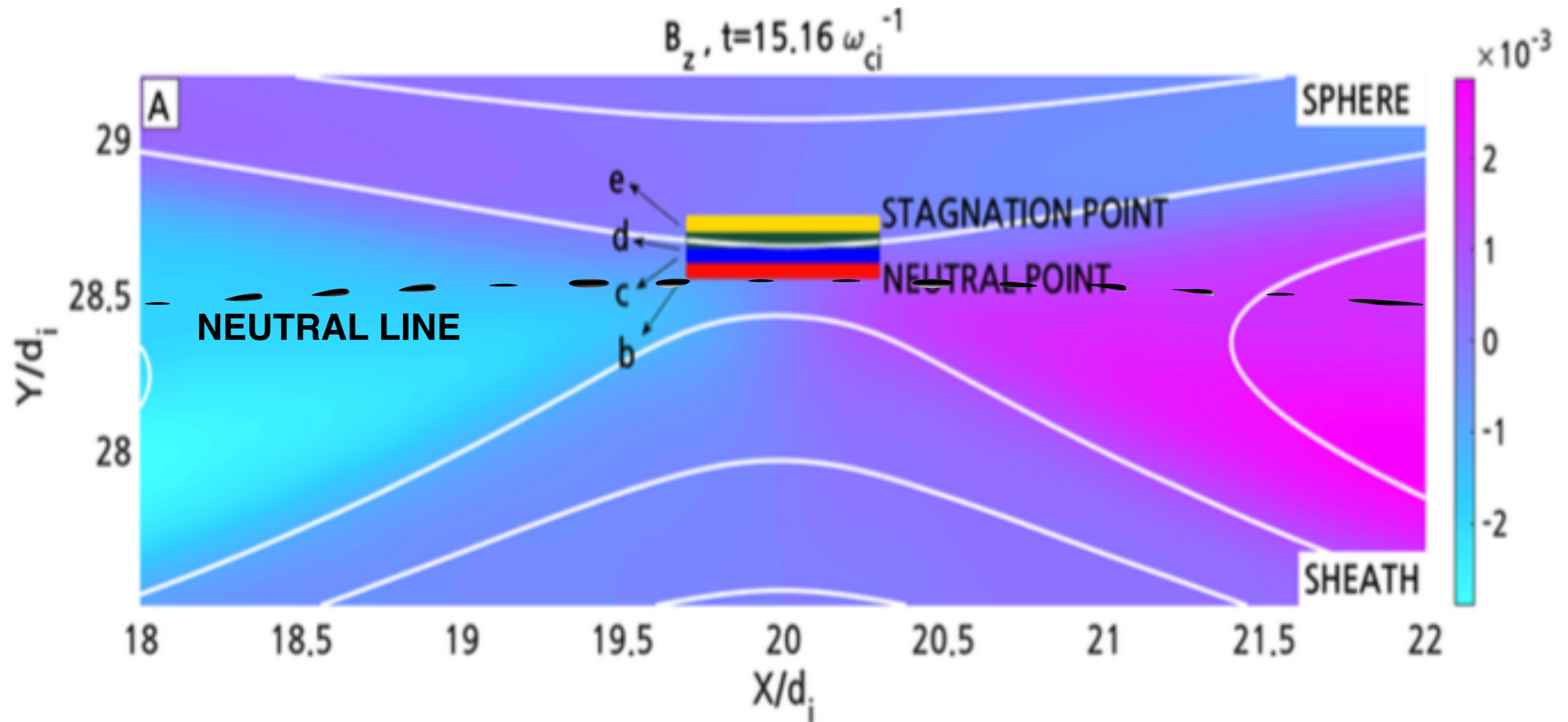
y/d_i

SHEAT

$1 d_i \sim 200 \text{ km}$

x/d_i

Focus on the Electron Diffusion Region



Identifiable by:

- inversion of the longitudinal magnetic field component, B_L (B_x)
- increased out of plane current, J_M (J_z)
- area of increased dissipation, $\mathbf{J} \cdot \mathbf{E} > 0$

Important because:

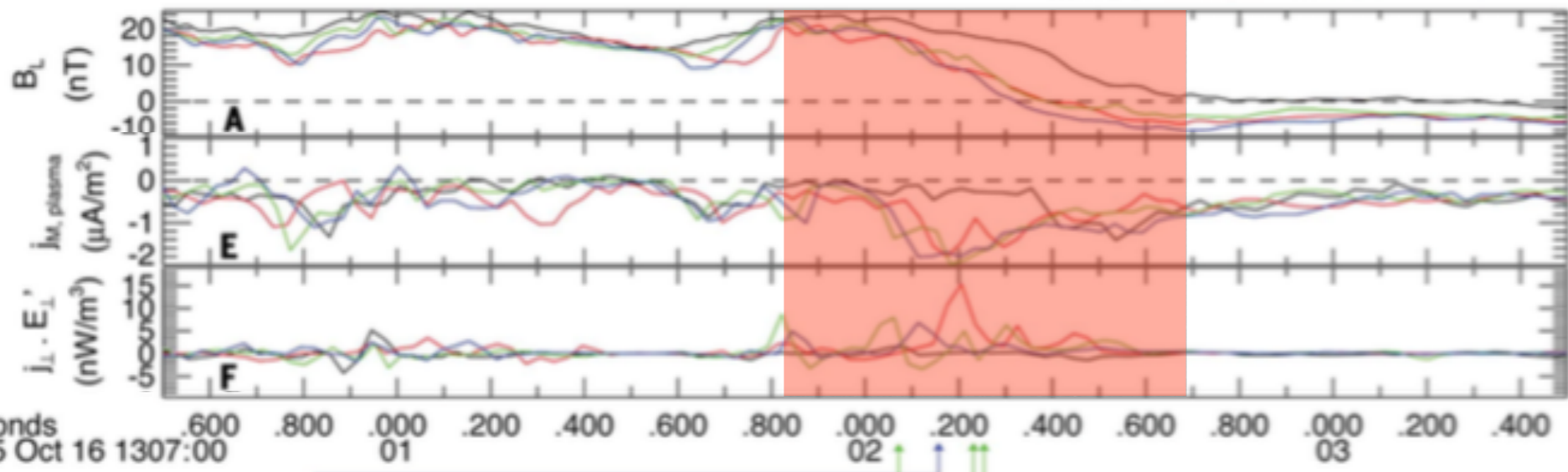
- driver of the reconnection process
- preferential channel of entrance for solar wind particles into the magnetospheric system

MMS encounters with Electron Diffusion and observation of the electron crescent

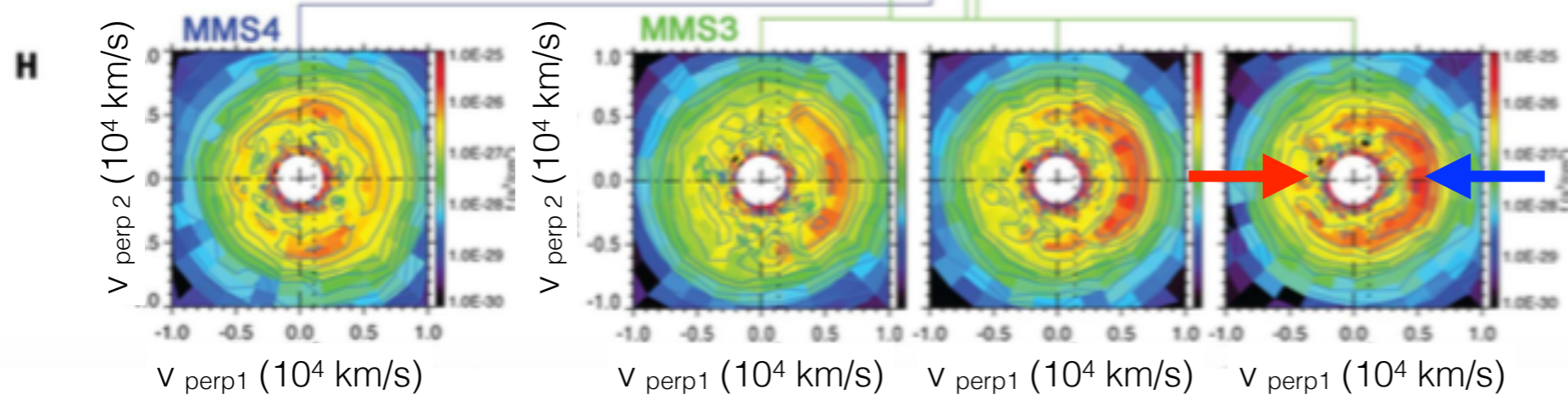
Burch et al., Science, 2016

SPHERE TO SHEAT
STAGNATION TO NEUTRAL POINT

MMS1 MMS2 MMS3 MMS4



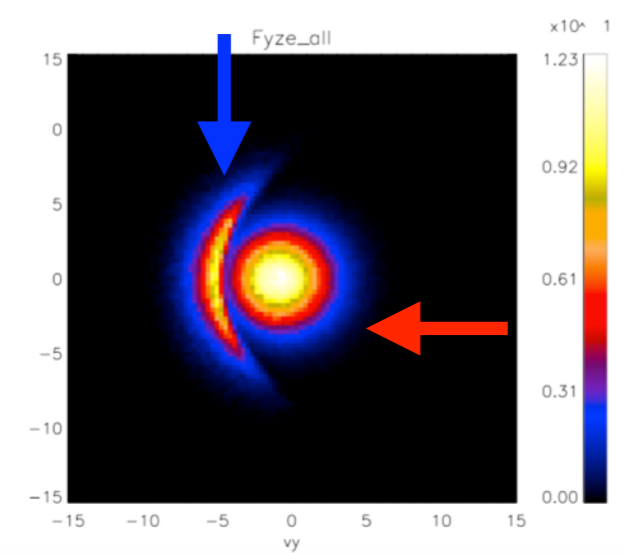
- inversion of B_L
- J_M peak
- J.E peak



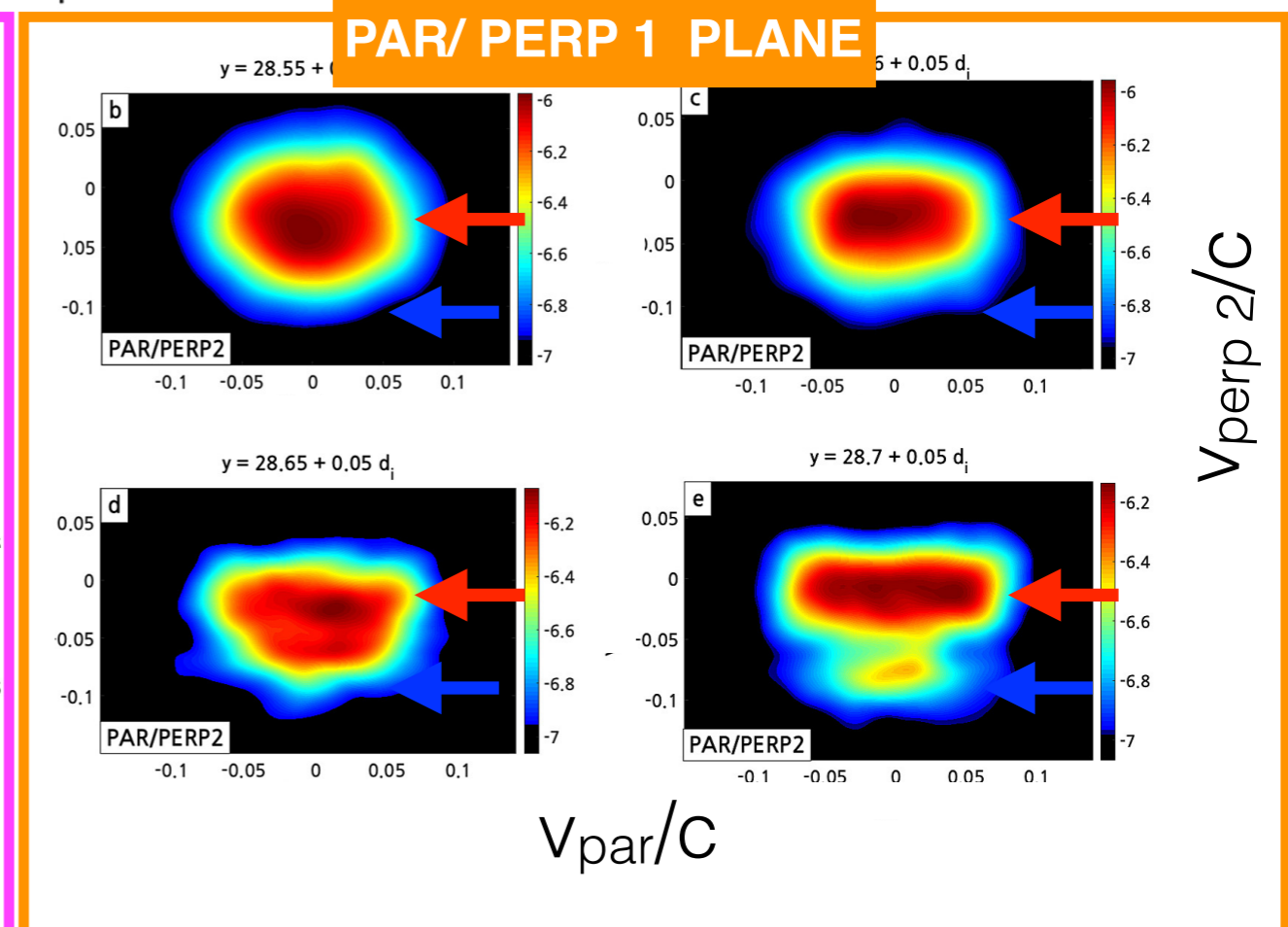
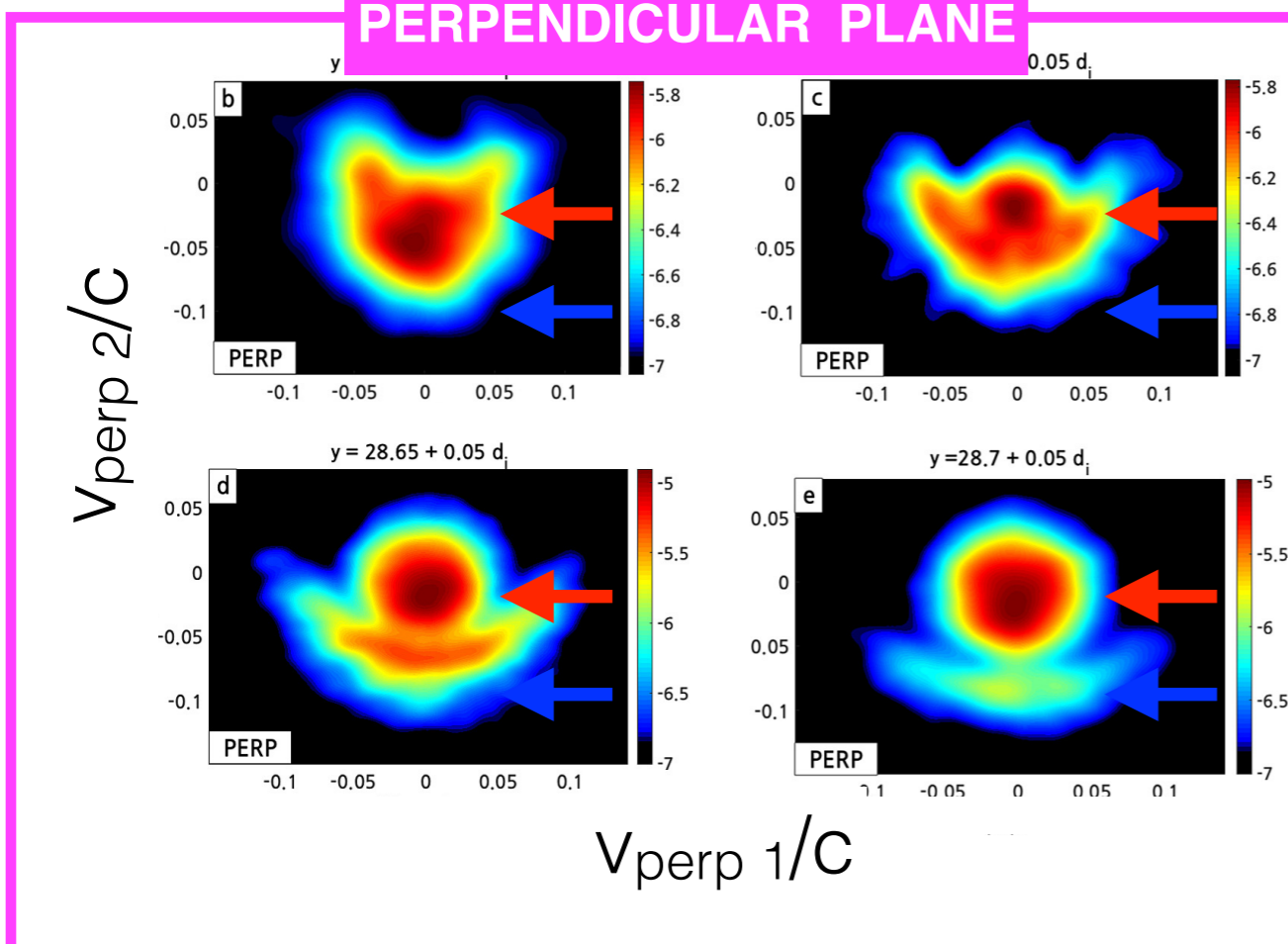
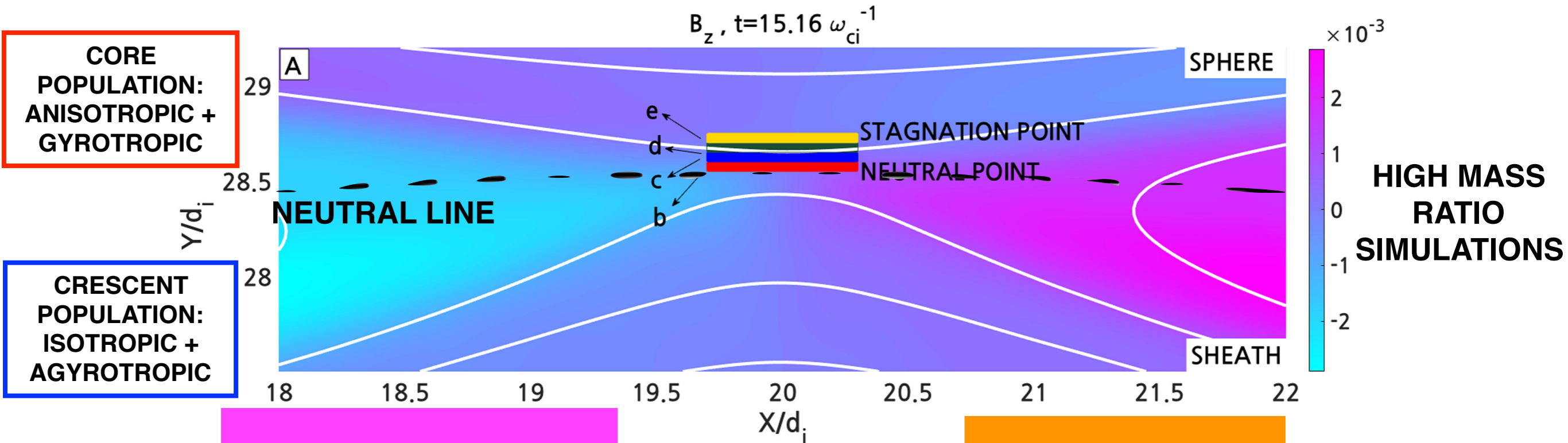
observation of a core + crescent structure in the perpendicular velocity space, as expected from Hesse et al, 2014, 16

CORE POPULATION: ANISOTROPIC + GYROTROPIC

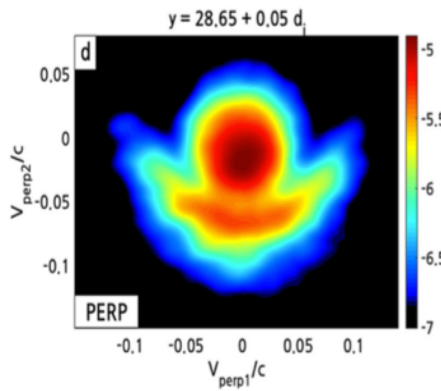
CRESCENT POPULATION: ISOTROPIC + AGYROTROPIC



Fully kinetic simulation of magnetopause reconnection



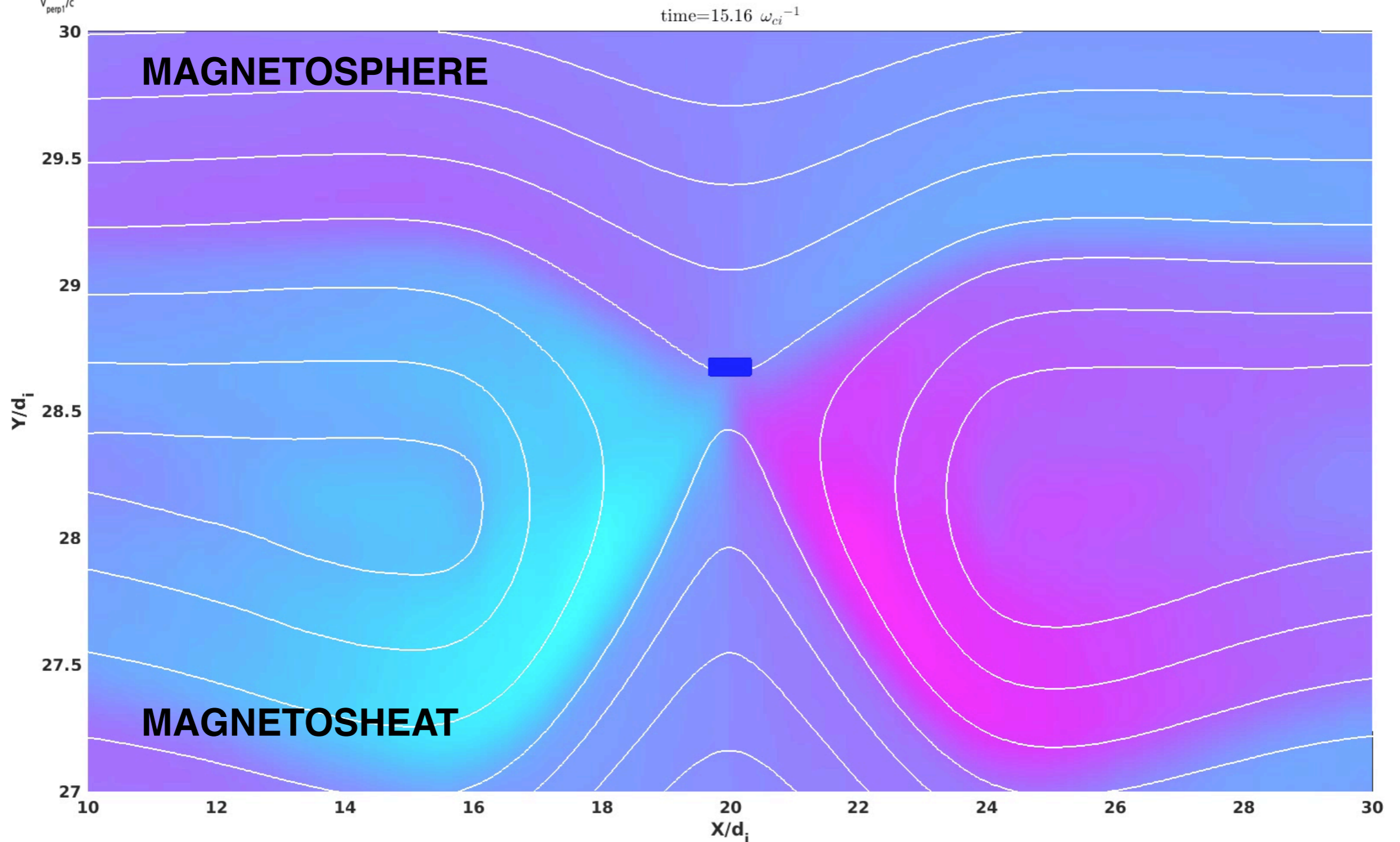
Origin of the core and crescent electron population



**CORE
POPULATION:
ANISOTROPIC +
GYROTROPIC**

**CRESCENT
POPULATION:
ISOTROPIC +
AGYROTROPIC**

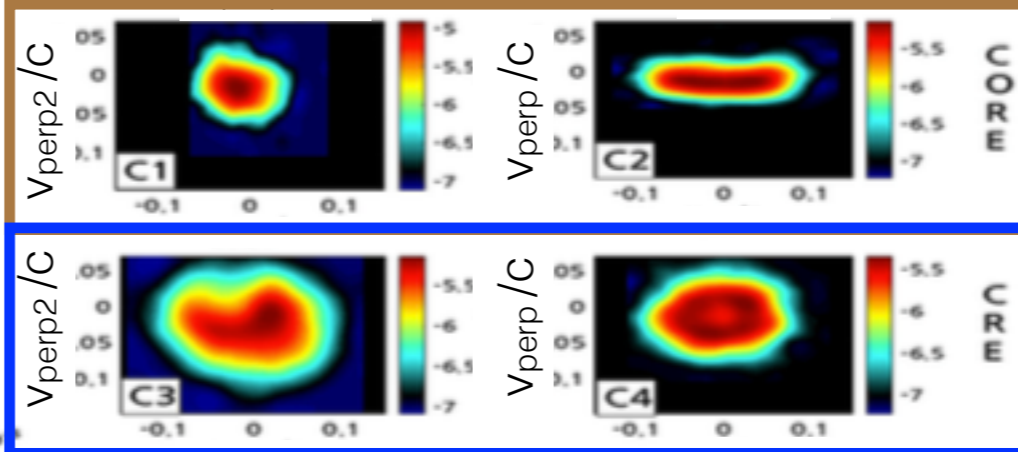
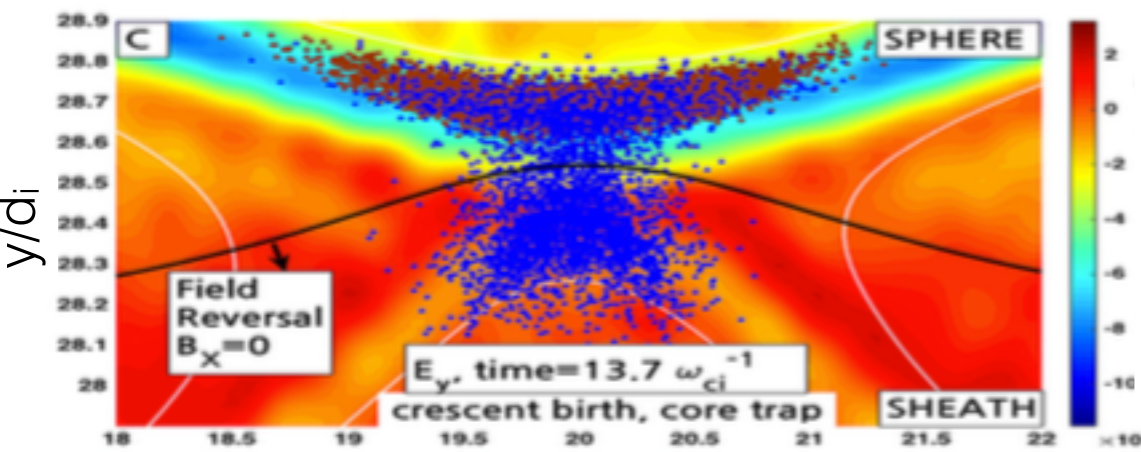
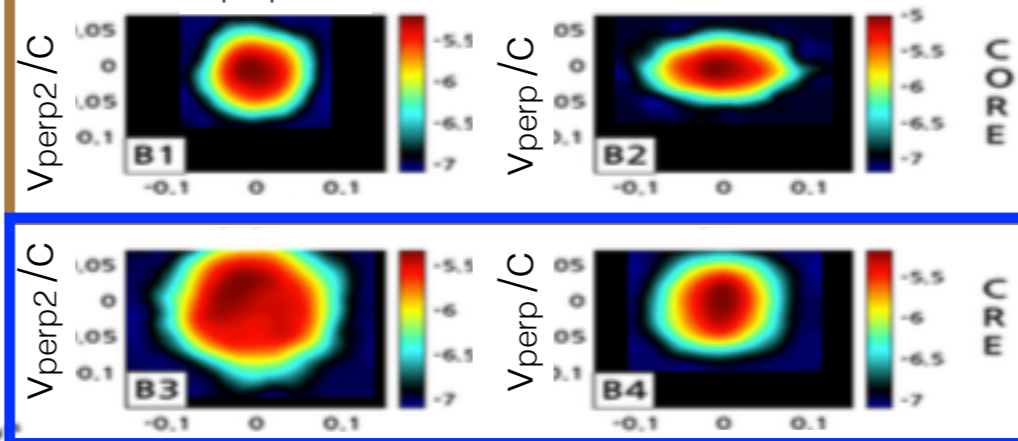
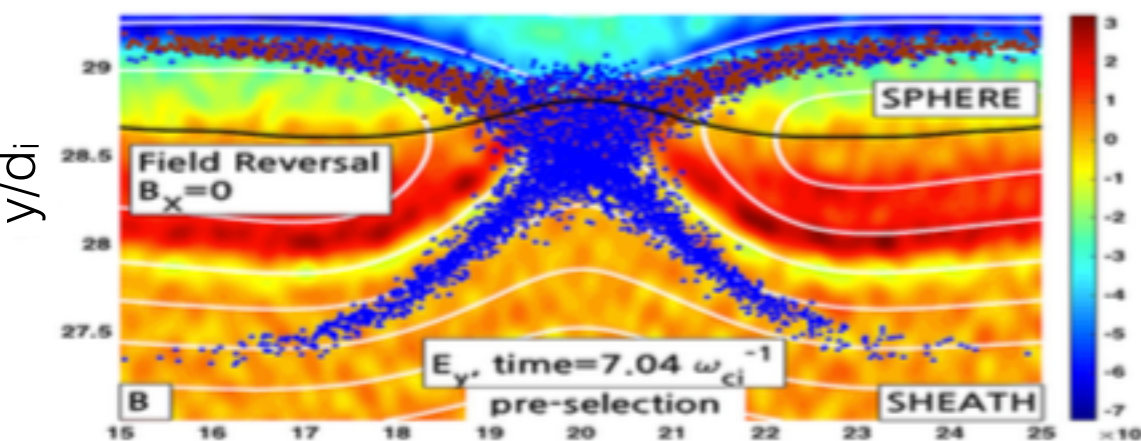
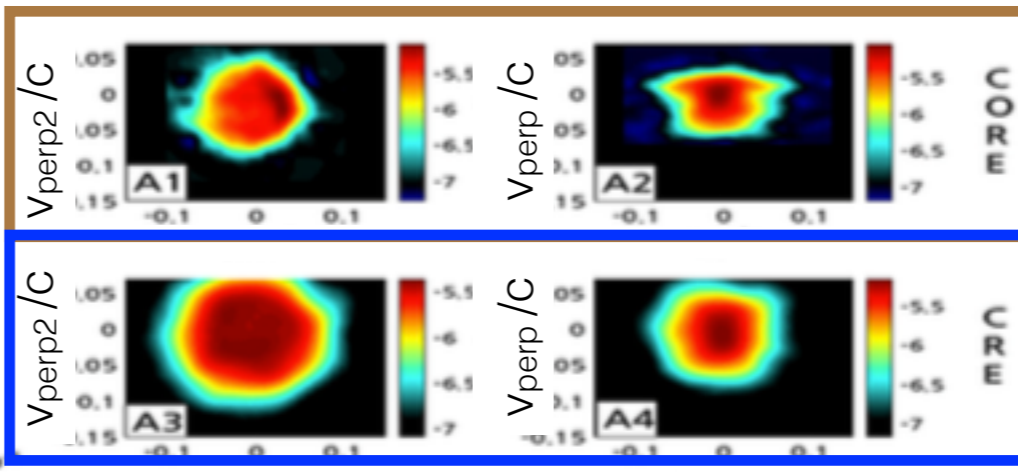
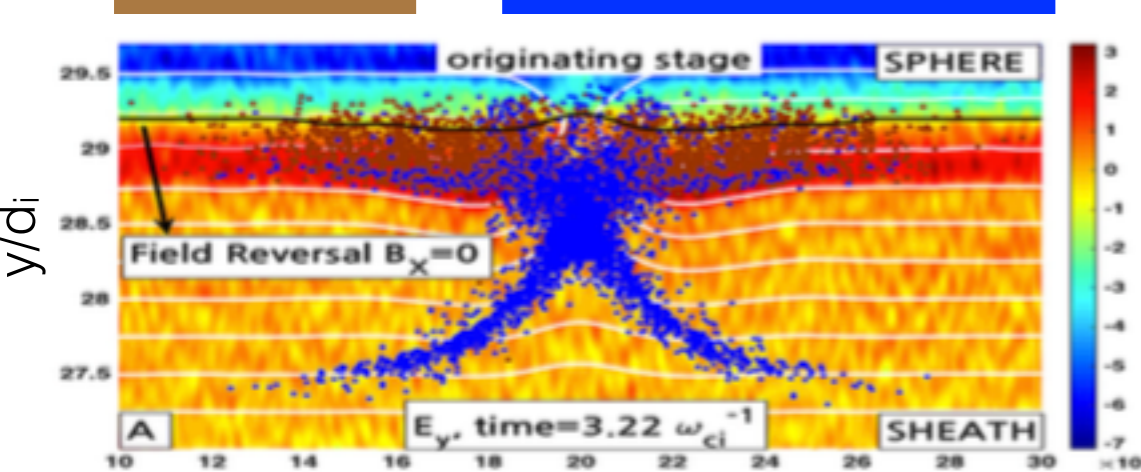
backtracing of core and
crescent electrons



Origin of the core and crescent electron population

CORE

CRESCENT



ORIGINATING STAGE

- crescent particles originate from deeper into the magnetosheat
- core particles cross the exhaust from the sheath to the upper separatrix
- crescent particles do not cross the exhaust

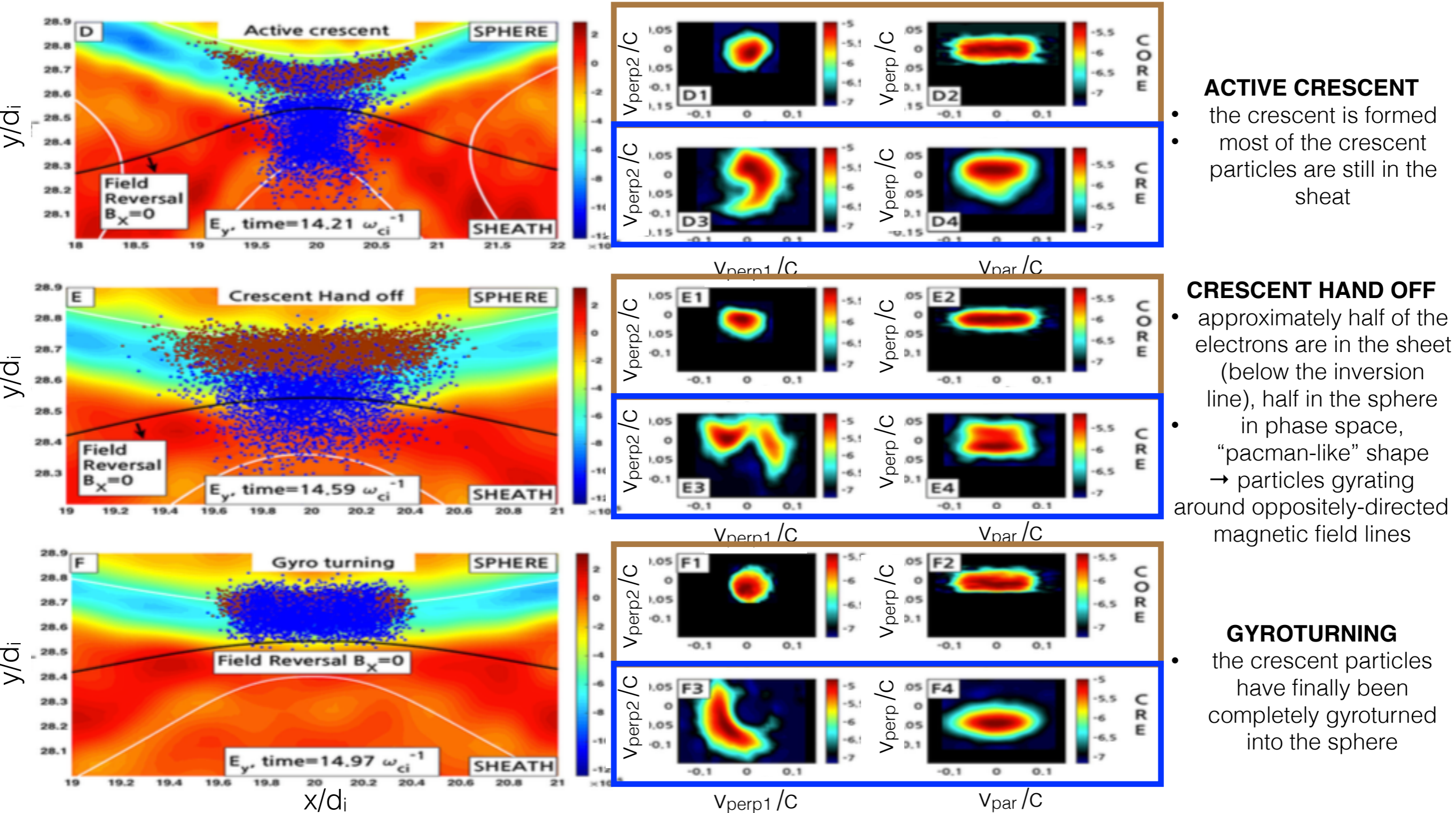
PRESELECTION STAGE

- crescent and core particles are separated in space
- differences in phase space start being evident

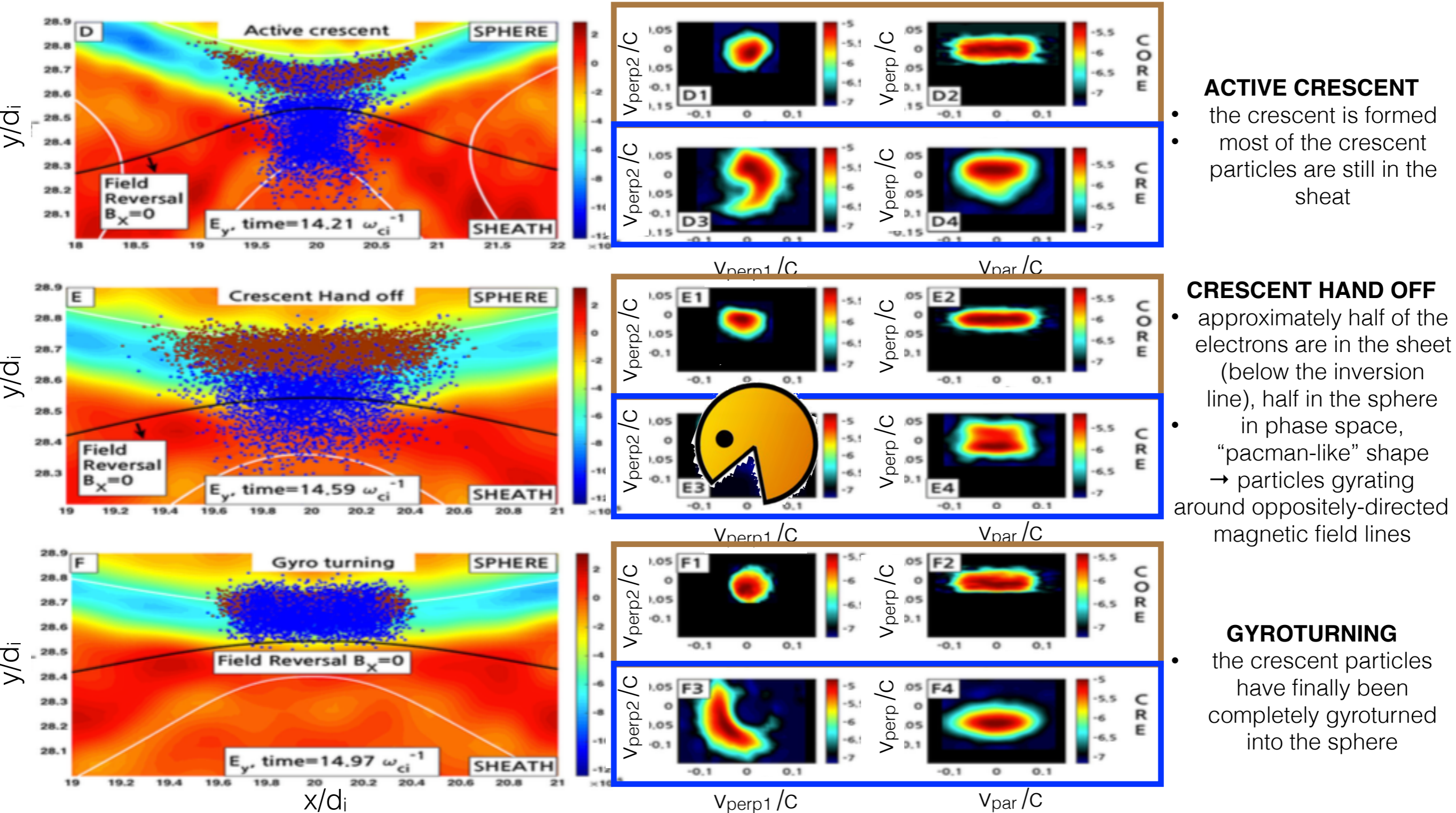
CRESCENT BIRTH & CORE TRAP

- core particles are segregated at the shoulder of the ambipolar field
- parallel acceleration
- crescent particles around the EDR
- preliminary signature of the crescent

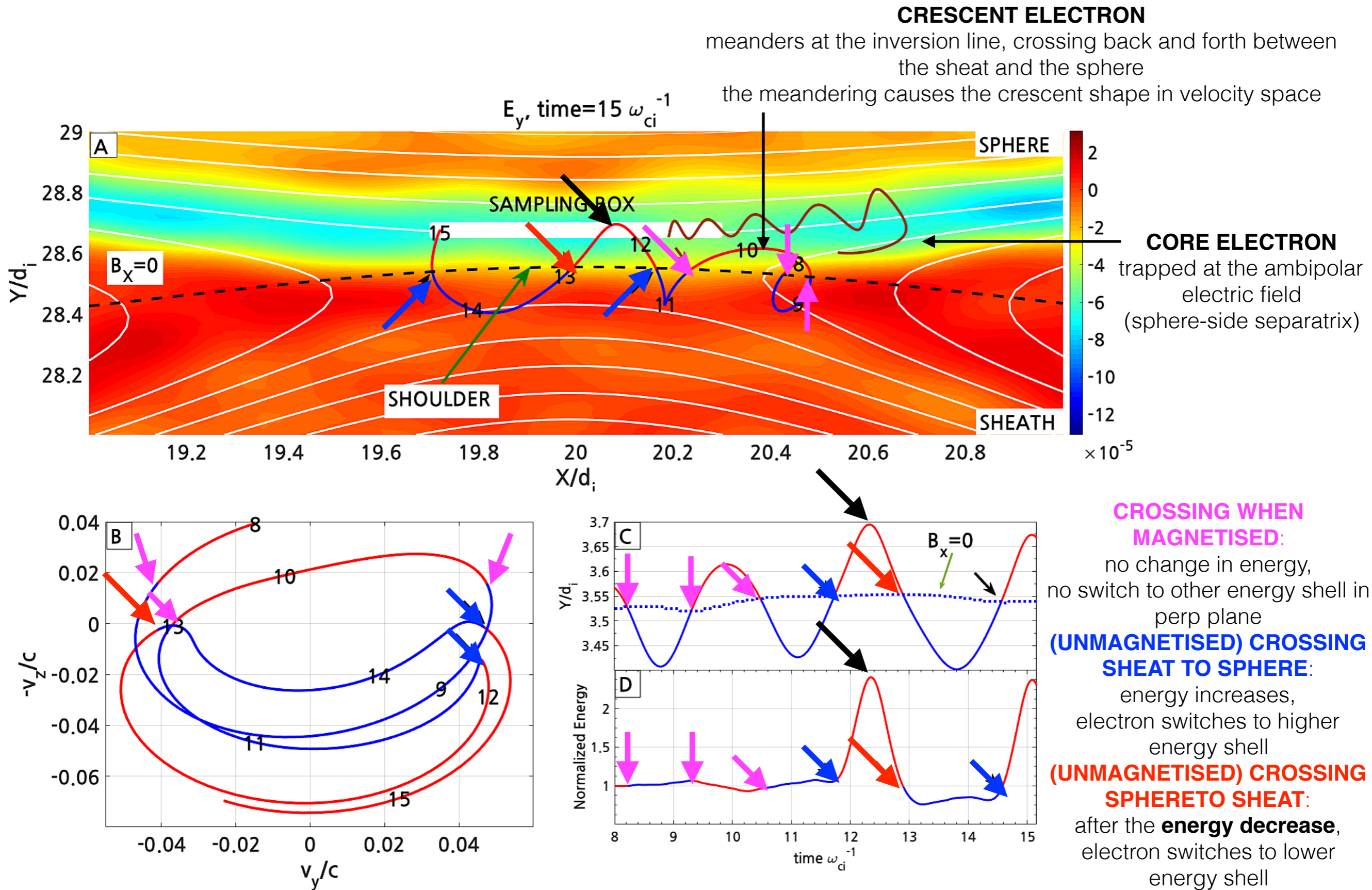
Origin of the core and crescent electron population



Origin of the core and crescent electron population



Trace of single electrons



Conclusions

Reconnection at the terrestrial magnetopause is a fundamental piece of the Sun-Earth connection puzzle

The MMS mission has been recently launched to investigate magnetopause reconnection at the ion and electron scales

The MMS can, for the first time, investigate velocity space with electron-scale resolution
→ observations and simulations are finally on an equal footing as regards electrons

We investigate the origin of the crescent distribution observed in the perpendicular velocity space with fully kinetic simulations

Our ion to electron mass ratio is higher than comparable simulations ($m_r = 256$ vs $m_r = 25$)

We identify different path of access to the Electron Diffusion Region for core and crescent electrons

Core electrons cross the exhaust and are eventually trapped at the sphere-side separatrix, where they get accelerated in the parallel direction by the ambipolar field

Crescent electrons meander back and forth between the sheath and the sphere; crossing the neutral line when unmagnetised translated to switching to an higher (sheath to sphere) or lower (sphere to sheath) energy shell
→ crescent formation

REFERENCE: Arokiaraj, Innocenti, Cazzola, Lapenta, “*On the electron mixing of the crescent and core populations in reconnection at the Earth’s magnetopause*”, in preparation



THANK YOU FOR YOUR ATTENTION