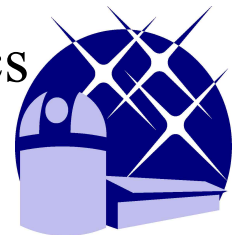


# ***Mercator Telescope @ the Roque***



P. Papics



Contact Group Meeting 25/05/2010

# ***HERMES***

## **HERMES: Intro**

**Consortium**

**Design Requirements**

## **HERMES: The Project**

**History of the project**

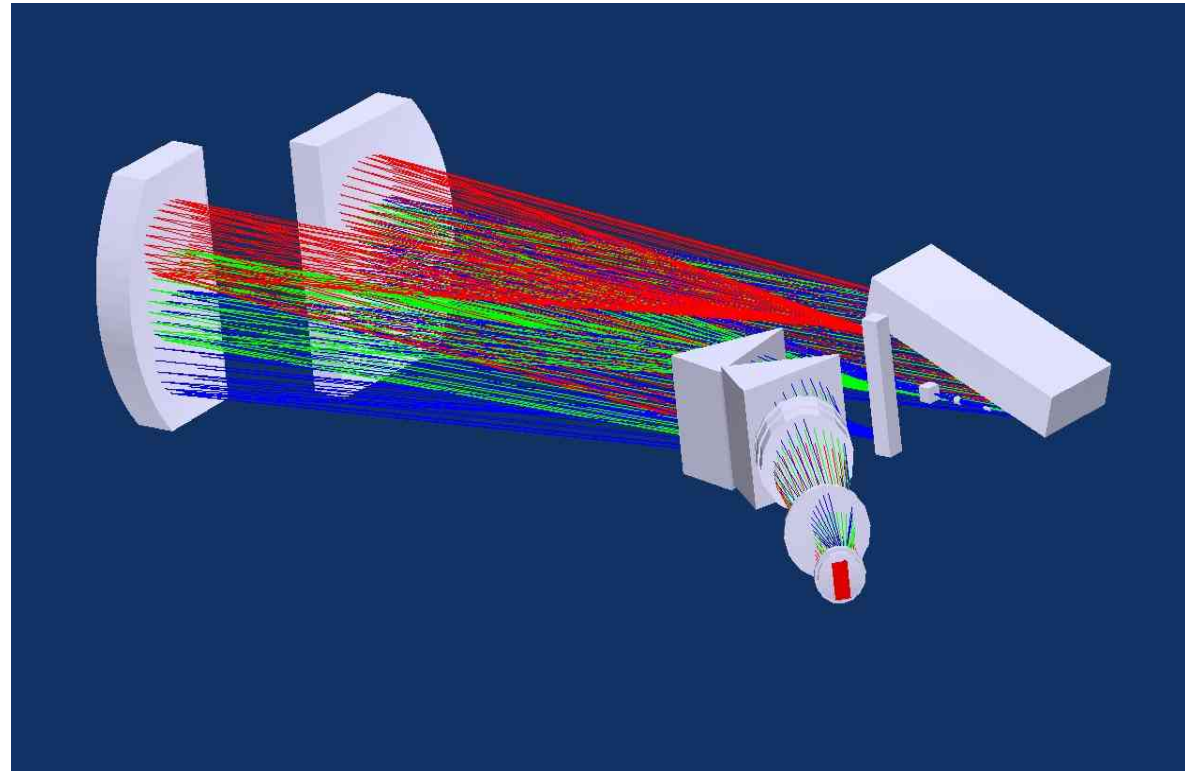
**Instrument:**

- **Design Spectro**
- **Adaptor, Fibre Link etc.**
- **Software**

**Characteristics**

**Running Mercator+HERMES**

**Some illustrative results**

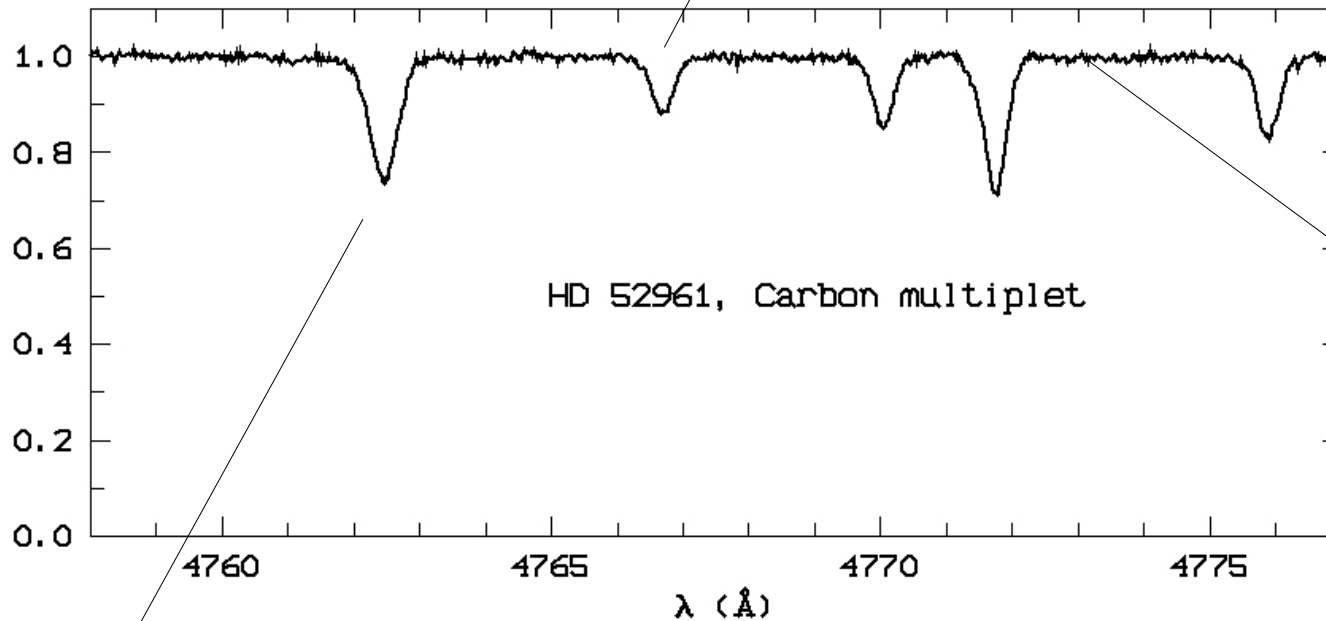


# Why a high-resolution spectrograph ?

Strength of line =  $f(T_{\text{eff}}, \log(g), \text{chemical composition})$

Wide spectral coverage !

Good sampling required

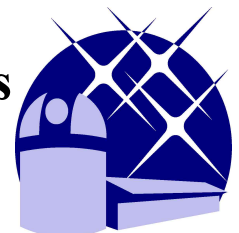


S/N: small telescope requires efficient instrument

Radial velocity :  $(\lambda - \lambda_0) / \lambda_0 = \text{velocity}/c$

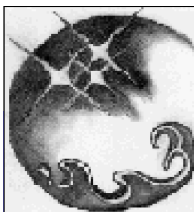
high resolution: 85 000 corresponds to 3.5 km/s

High stability to allow 10 m/s accuracy. Wide spectral domain helps



# ***HERMES-Consortium: kick-off in Leuven 19/01/2005***

**Proj. Engineer: G. Raskin**  
**PI : H. Van Winckel**  
**Co-i: C. Waelkens**  
**W. Pessemier, S. Prins**  
**J. Padilla, F. Merges**  
**J. Morren, P. Degroote**  
**K. Smolders**



IvS-KUL



ROB

**Co-i: H. Hensberge**  
**Y. Fremat**  
**L. Dumortier**  
**J.P. De Cuyper**

**Co-i: A. Jorissen**  
**S. Van Eck**  
**C. Siopis**



IAA-ULB

# **HERMES**

**Co-i: H. Lehmann**  
**J. Winkler**  
**U. Laux**



Landessternwarte Tautenburg

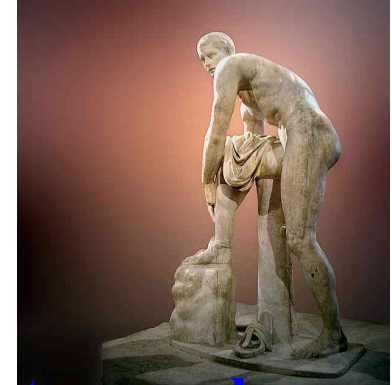


Observatoire de Genève

**R. Dubosson**  
**B. Michaud**



# ***HERMES: Niche in the telescope-market***



**Time series in radial velocity and in individual spectral lines,  
high S/N spectra of fainter stars: robust, efficient,  
high-resolution, spectrograph on Mercator**

=

**HERMES: High-Efficiency and Resolution Mercator Echelle Spectrograph**

**Pooling of HERMES observations with priority driven scheduling.**

**Priorities are given by consortium board + external TAC**

**All HERMES consortium nights are mainly (80%) scheduled from pool.**

**HERMES consortium time: 50% of remaining telescope time.**



## ***HERMES: Science drivers: stellar astrophysics***

**KUL :** Asteroseismology (B-stars); late phases of (binary) stellar evolution; AGB Nucleosynthesis (single and binary stars; s-process)

**ULB:** Late phases of (binary) stellar evolution; Nucleosynthesis (RGB; AGB, s-r-process). Interacting binaries; Doppler tomography

**ROB:** Asteroseismology; Wind variability in B-stars (single and binaries)

**TLS (Tautenburg):** Asteroseismology, extra-solar planets

**Geneva :** Asteroseismology (solar-like oscillations), extra-solar planets



# ***HERMES: User Requirements***

**HERMES is not a single goal instrument. Some compromise between efficiency and stability is needed. Solution: 2 different Science Fibres**

**Project requirements include Spectrograph requirements and Operational requirements.**

**Spectrograph requirements (goals):**

**Range : 380 - 880 in single shot**

**Fibre Fed**

**Resolution (500nm): min 40 000 – max 90 000**

**Efficiency: 25% (maximal efficiency)**

**Stability: 1 m/s (short term)**

**Stability: 5 m/s (long term)**

**Cycling time: 30-60 sec.**

**DRS**



# ***HERMES: Operational Requirements***

## **Operational requirements:**

Automatic: including pointing, centering, guiding

On-line full data reduction pipeline

Efficient Semi-automatic Scheduling from pooled programs

Archiving System

## **Overall requirements:**

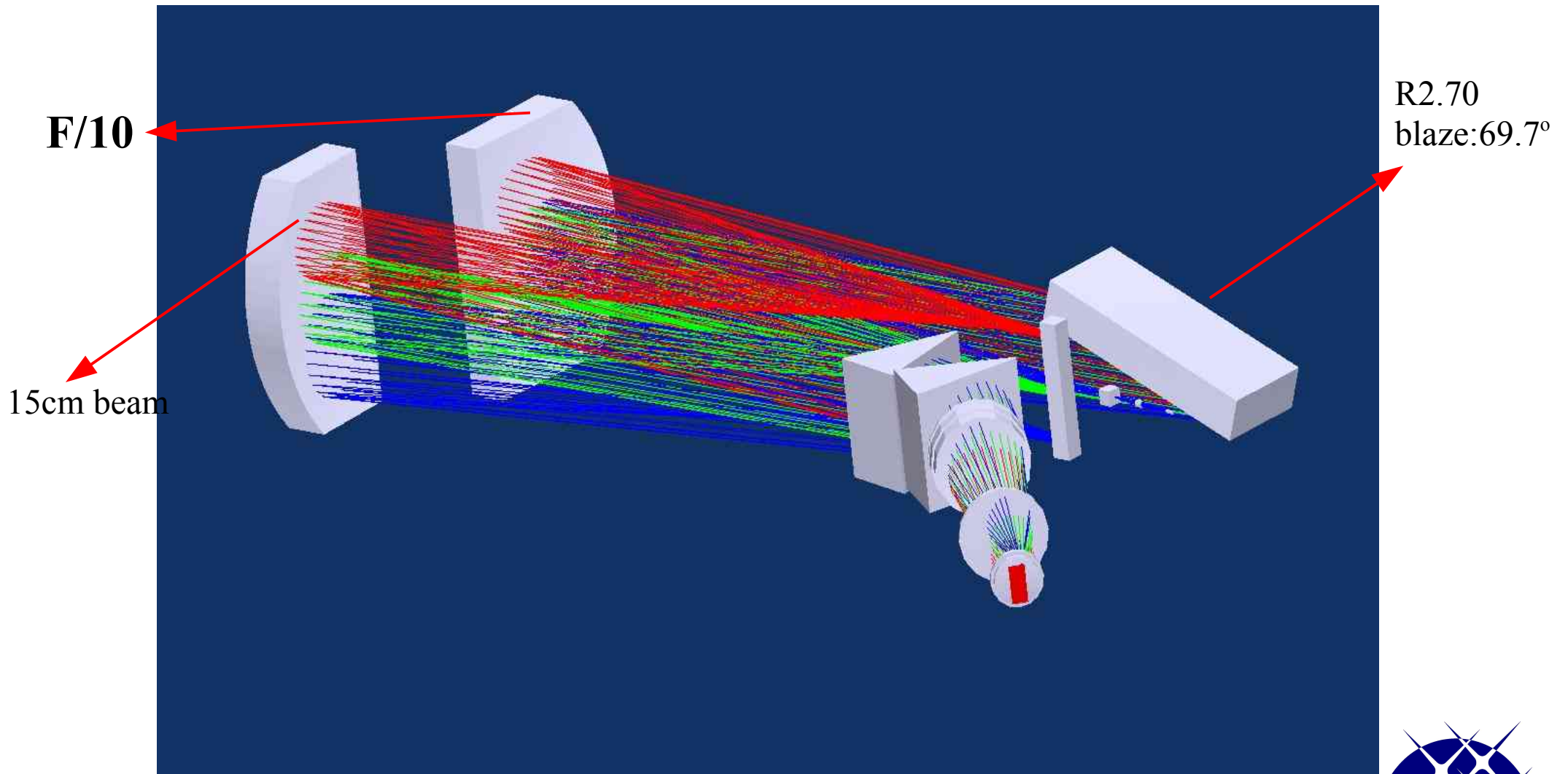
Robust design, sober in maintenance efforts and costs

Within Budget (980 kEuro)

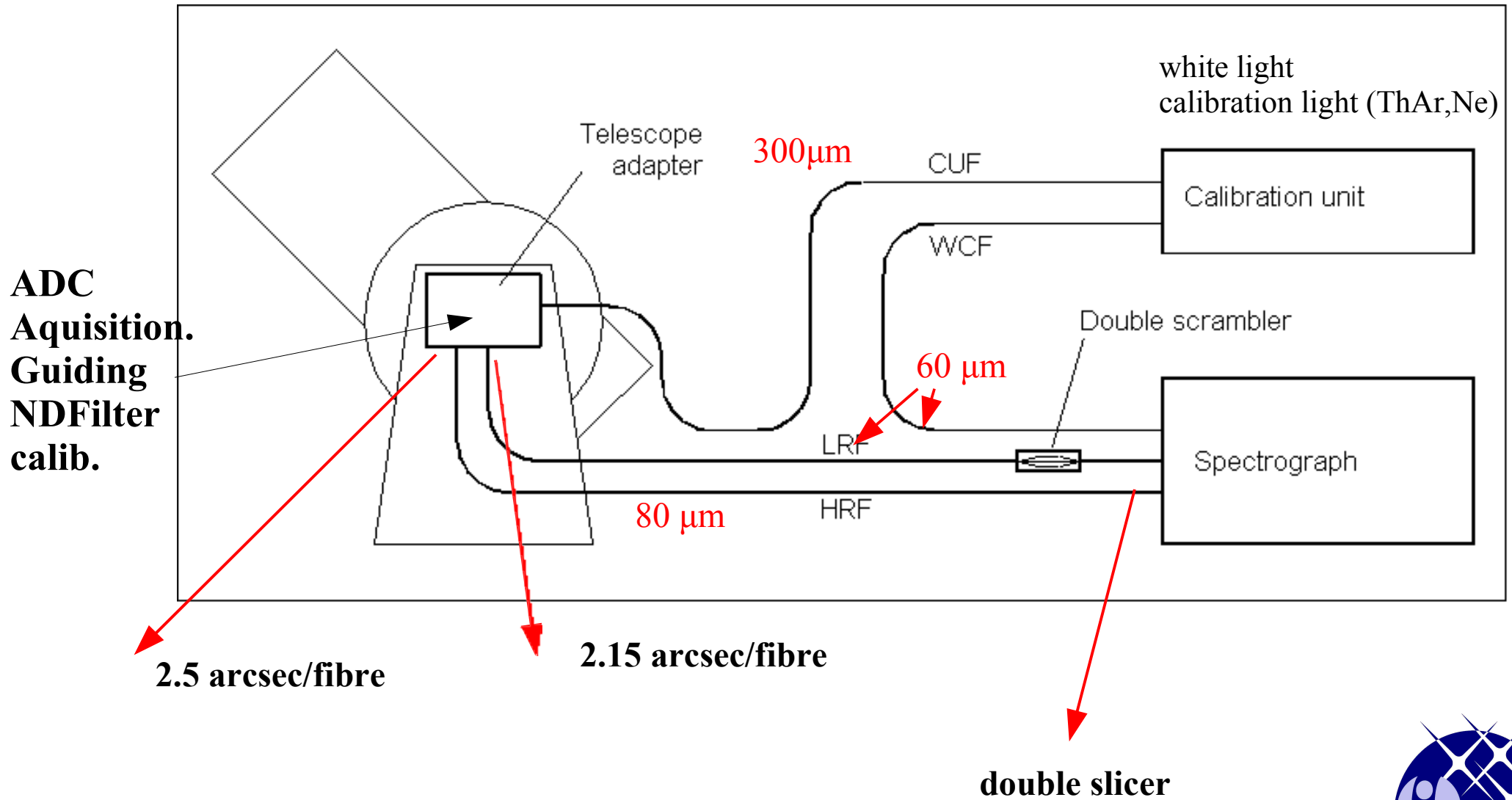




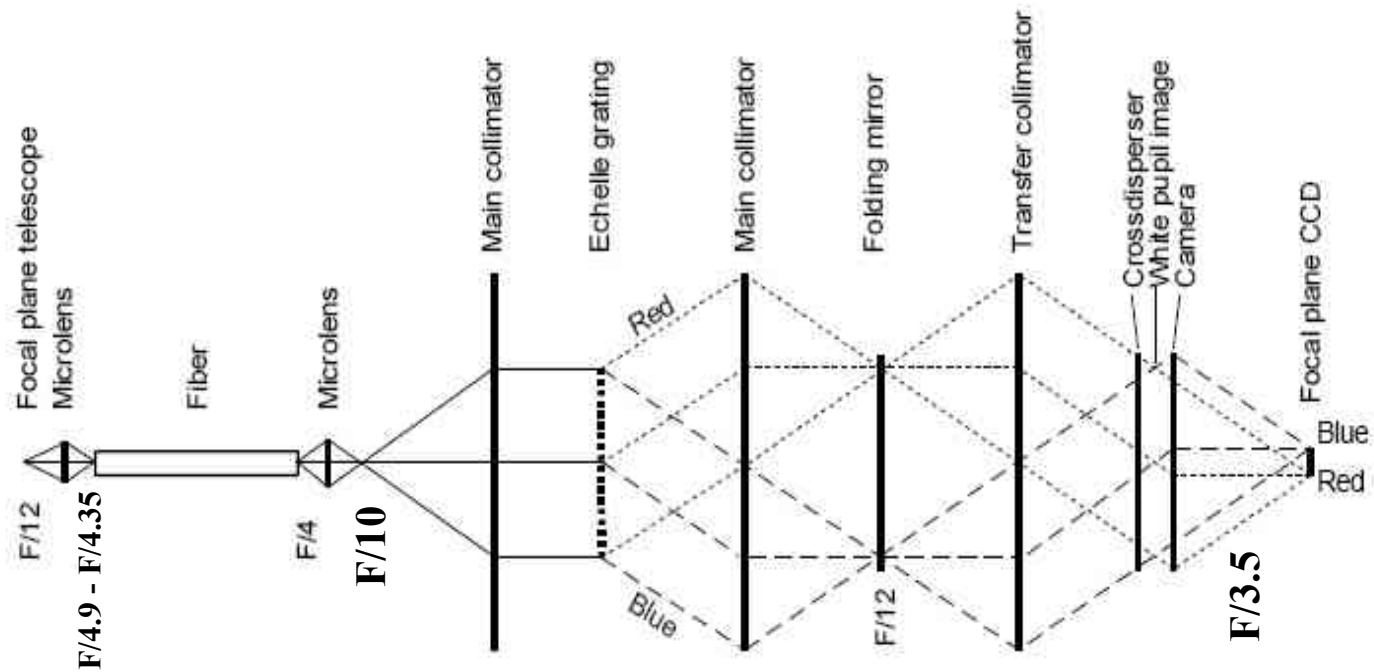
# *Hermes: white pupil Design*



# Hermes Design



# HERMES design



Overview is given following the optical path



# FOD

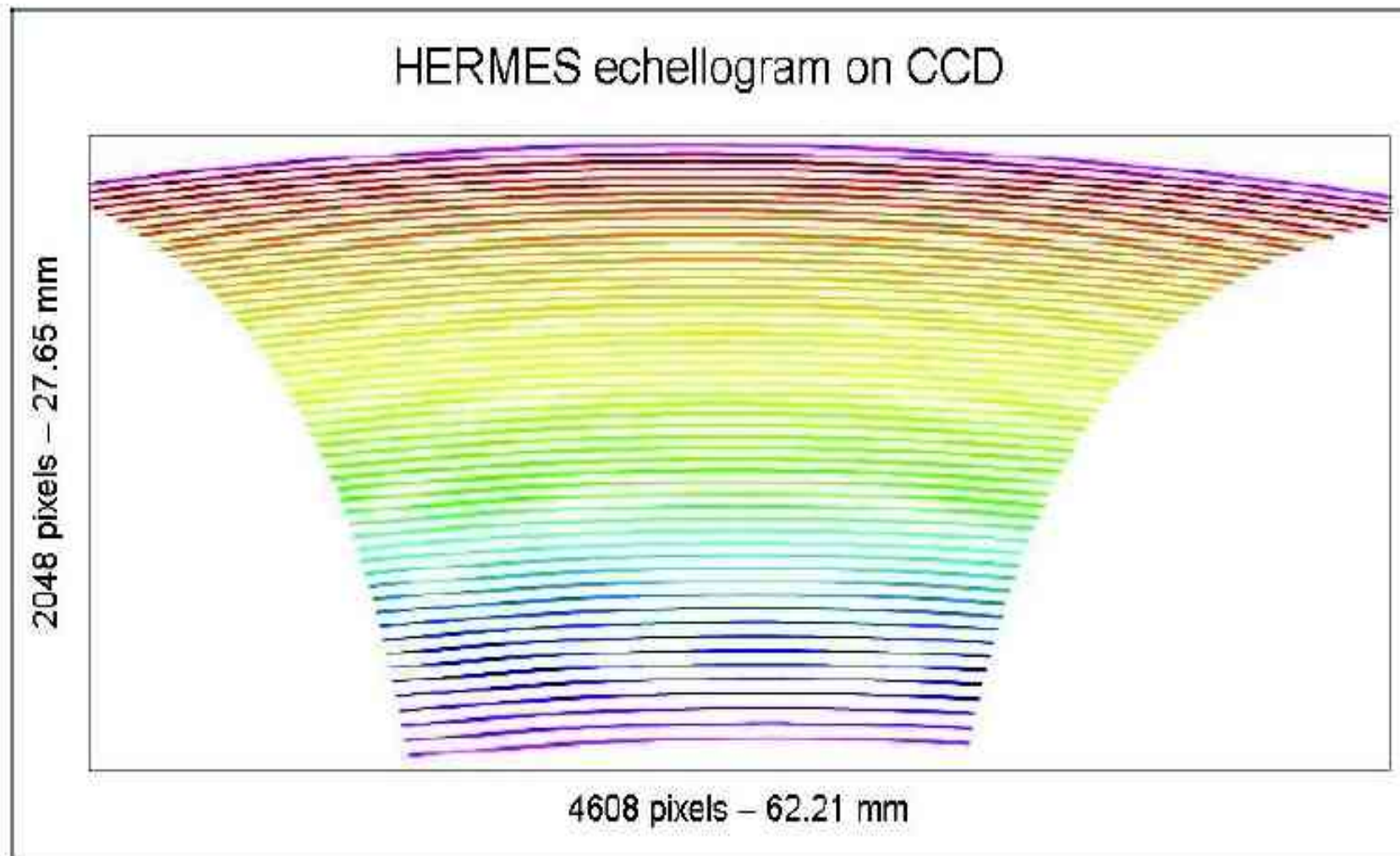


Fig. 17. Layout of the 55 orders on the CCD (color scale is compressed for visibility).



# **Project History: milestones**

- \* **External Project Reviewers: European Experts in Instrumentation  
Dr. Francesco Pepe (Geneva), Prof. Dr. Otmar Stahl (Heidelberg),  
Dr. Gerardo Avila (ESO), Prof. Dr. Ramon Garcia Lopez (IAC)**

**Final Optical Design Review: 12/05/2005, Leuven**

**Preliminary Design Review: 27/01/2006, Leuven**

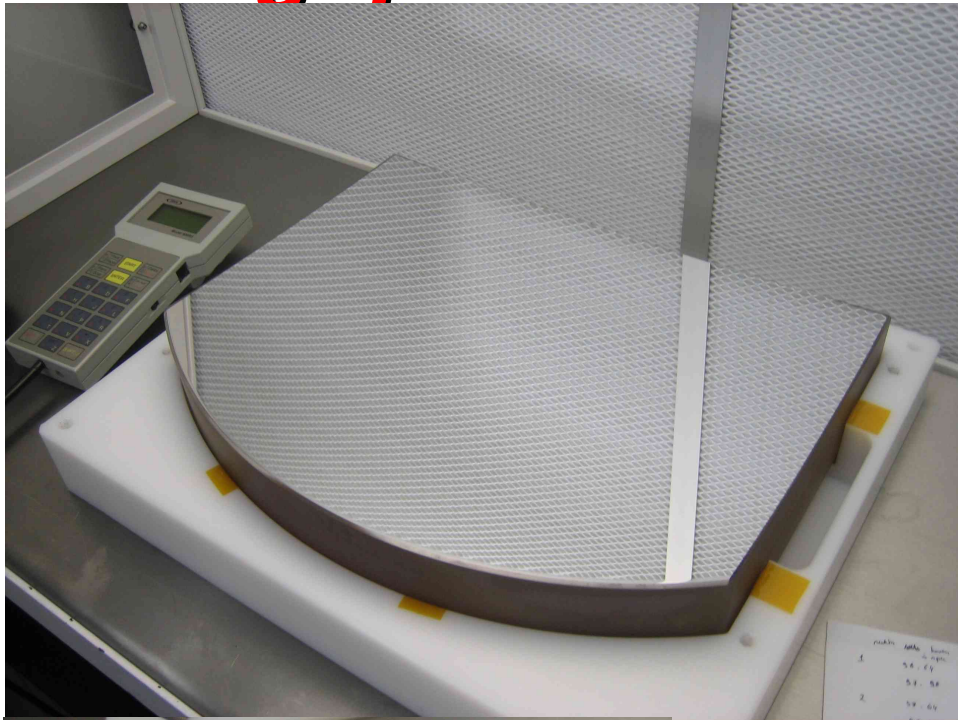
- \* **Web-based document system and version control system for software  
(adapted from IvS investment in PACS project)**
- \* **Feb 2006 – April 2009 optimization design, ordering +quality control  
optics, final detailed drawings mechanics, procurement (workshops  
department KUL, Tautenburg and Geneva), integration on site and  
software, software, software**



# ***Spectrograph Optics***



# Coating: *protected UV-enhanced silver*

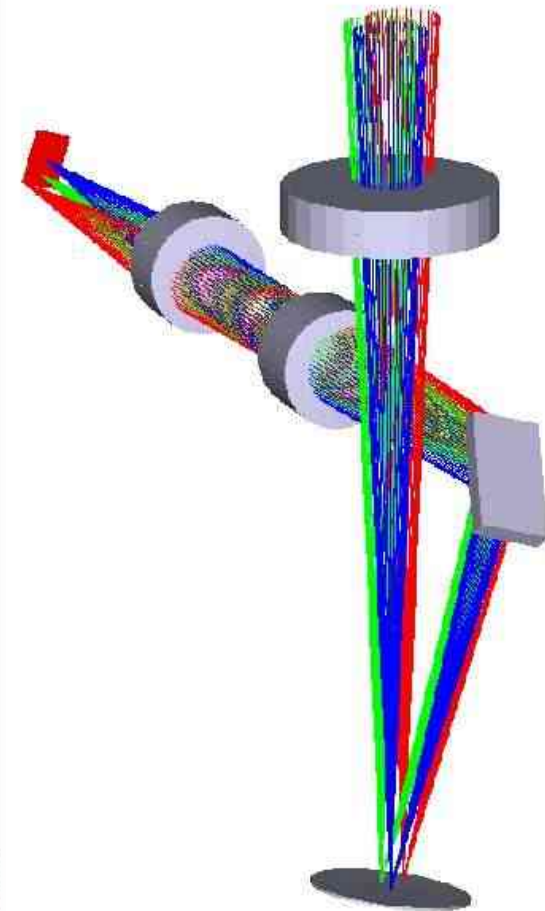
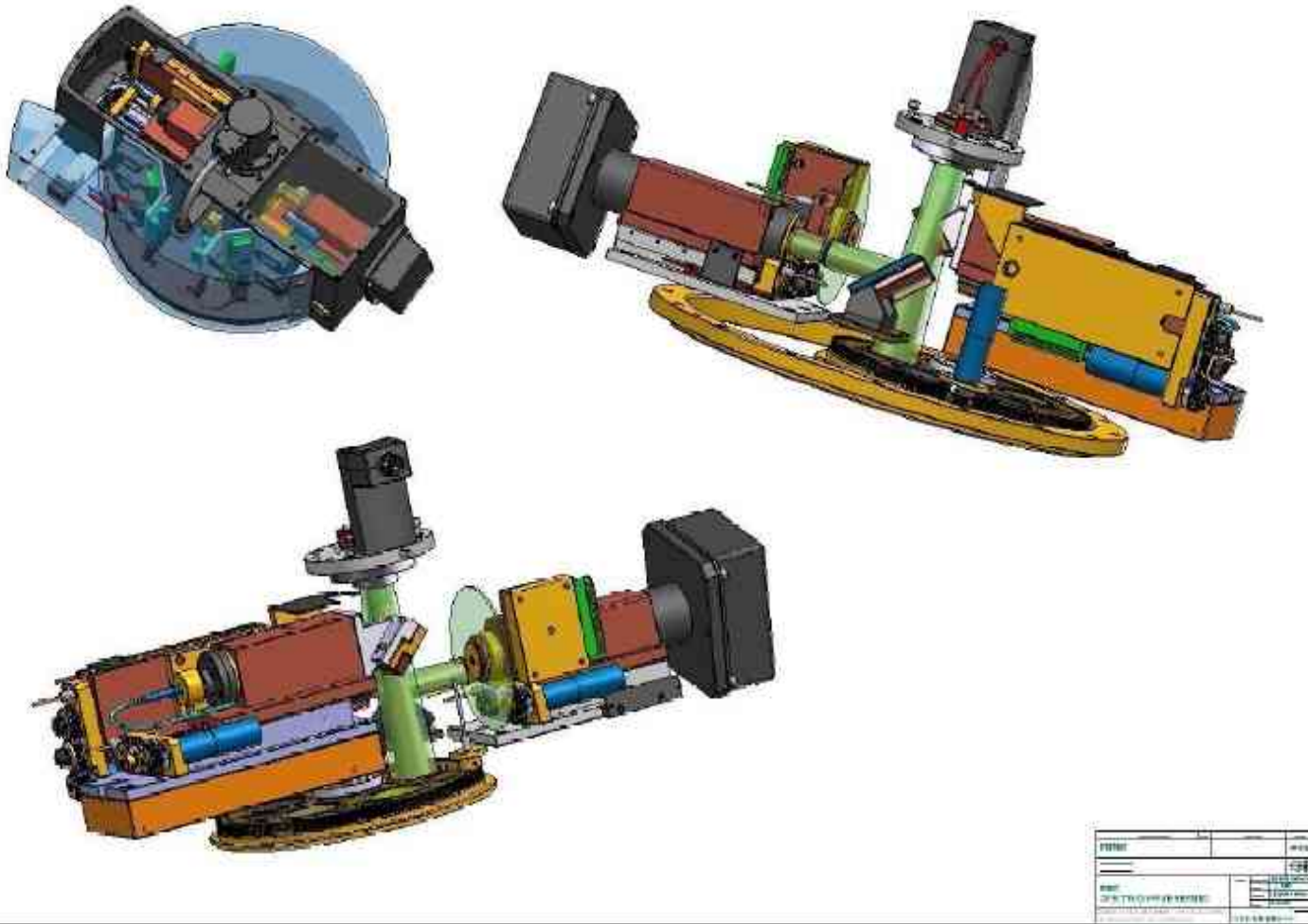


received :  
10/9/2008



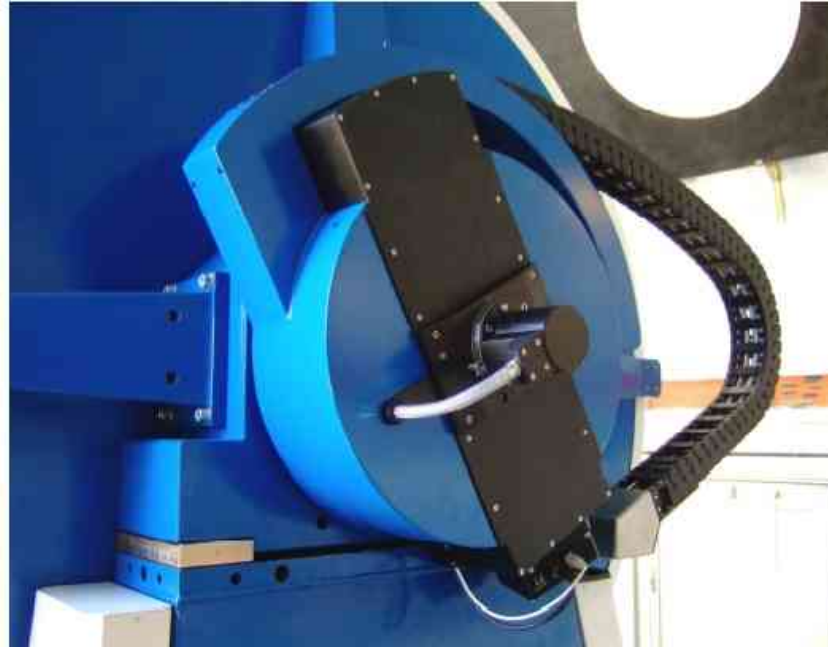
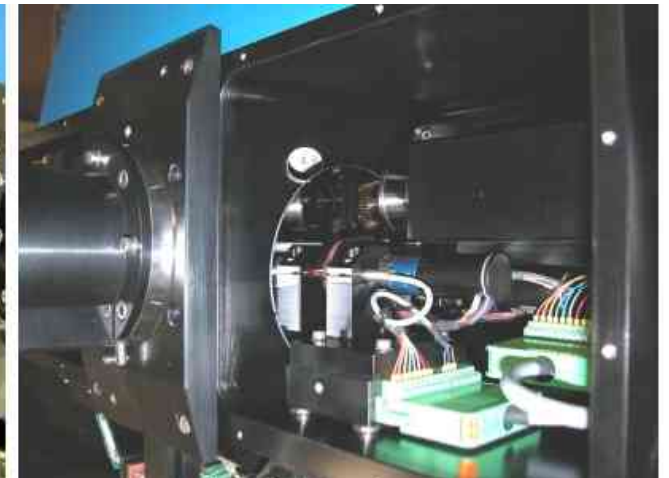
# 4 discrete ADC positions (380-900nm)

zenith distance	length without ADC	with ADC
31	1.12''	0.61''
60	2.71''	0.33''





# *Adaptor*



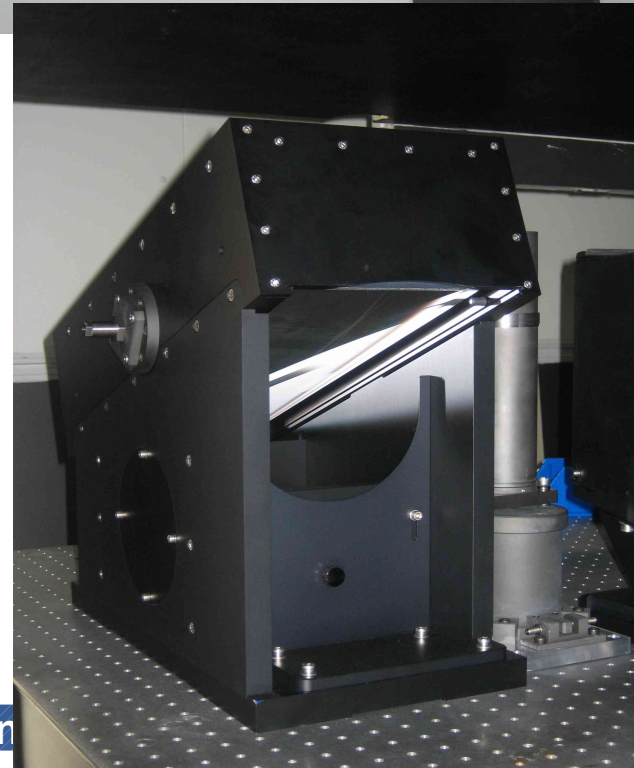
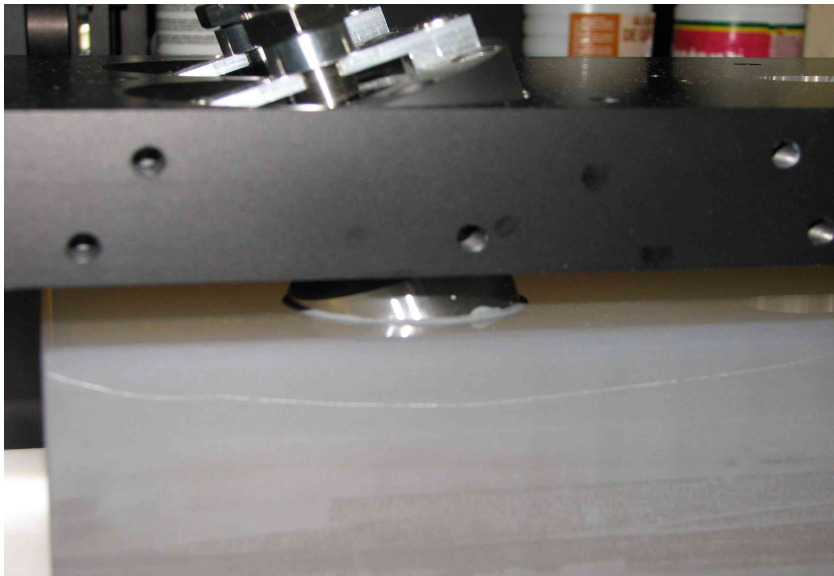
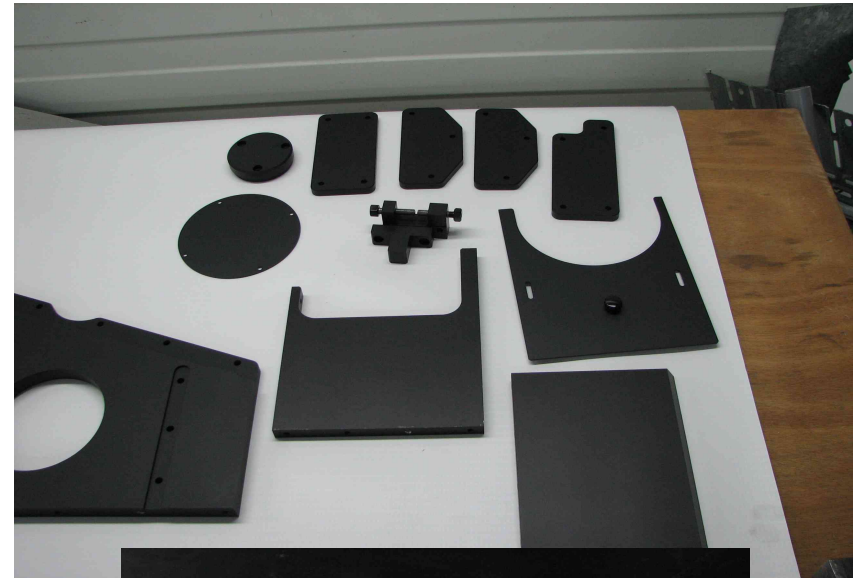
8 July 2008



***Mounts: 320 kg of mechanics were shipped in August 2008***



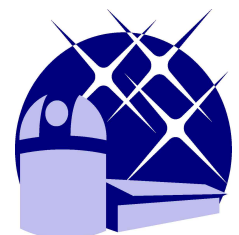
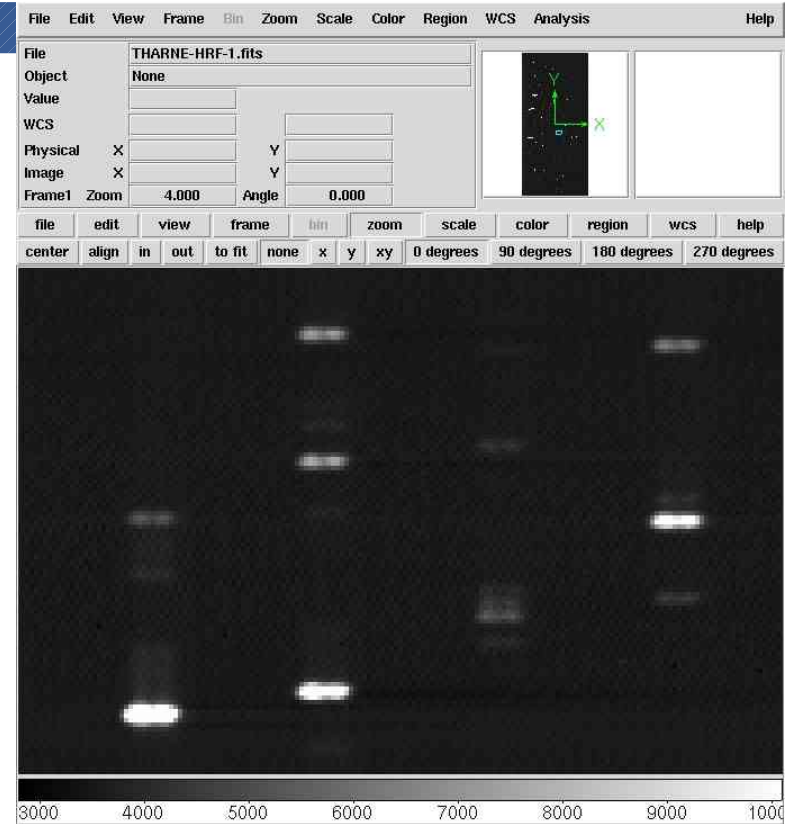
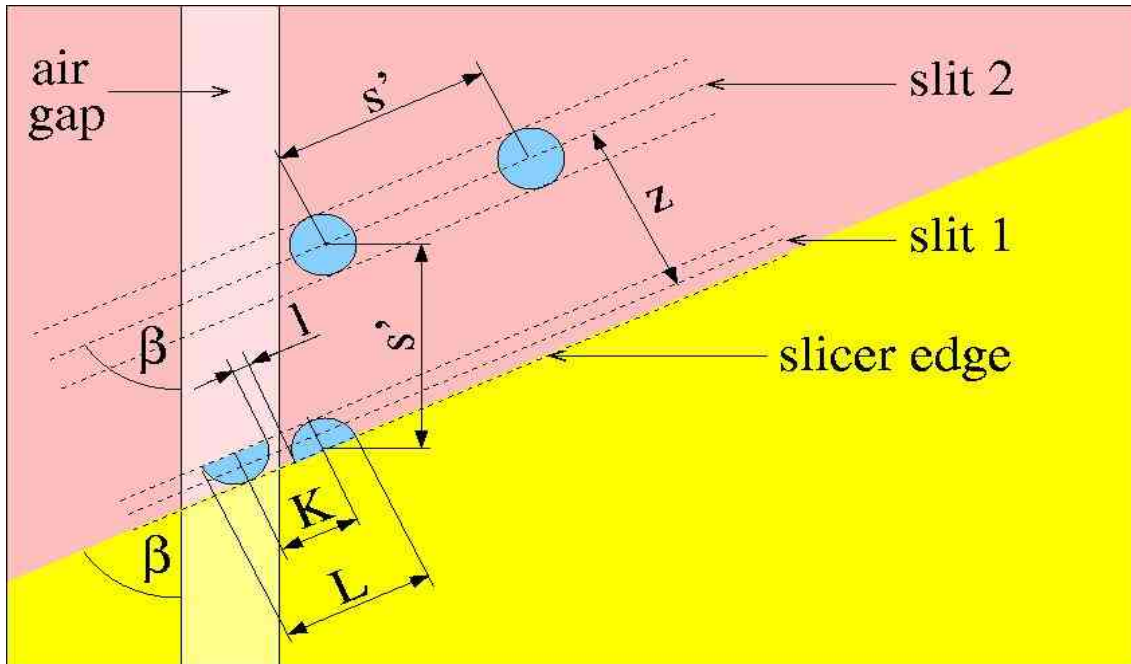
# Grating



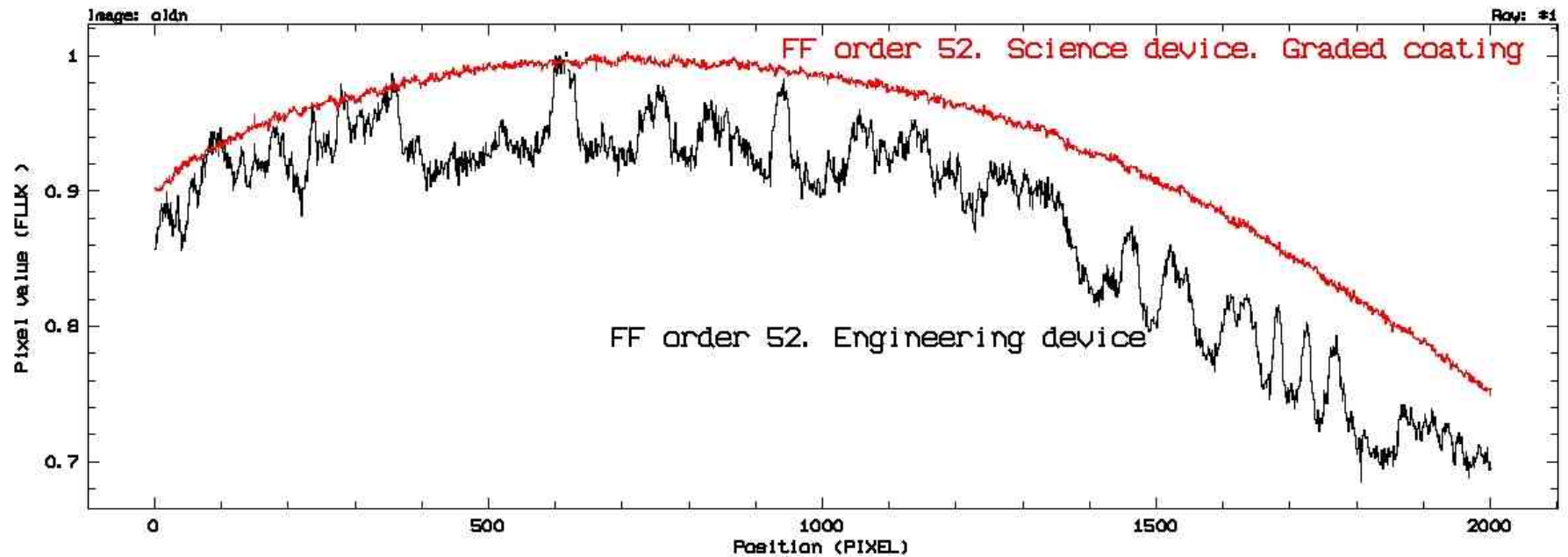
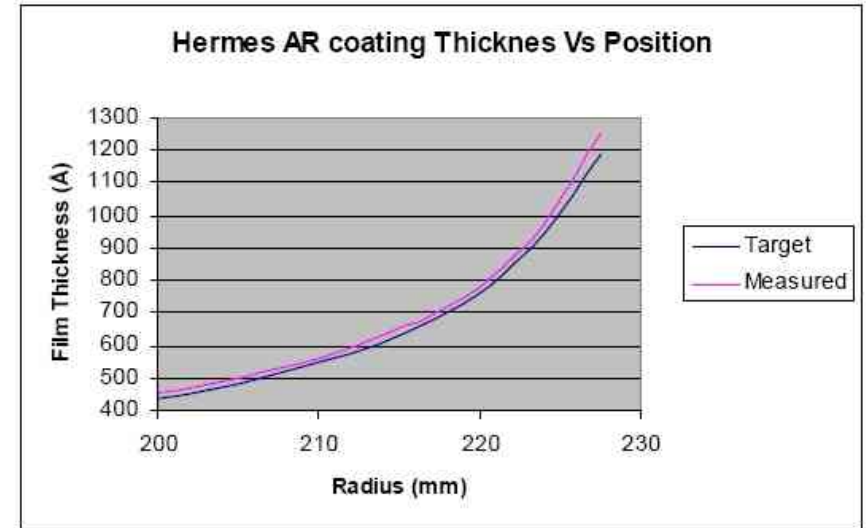
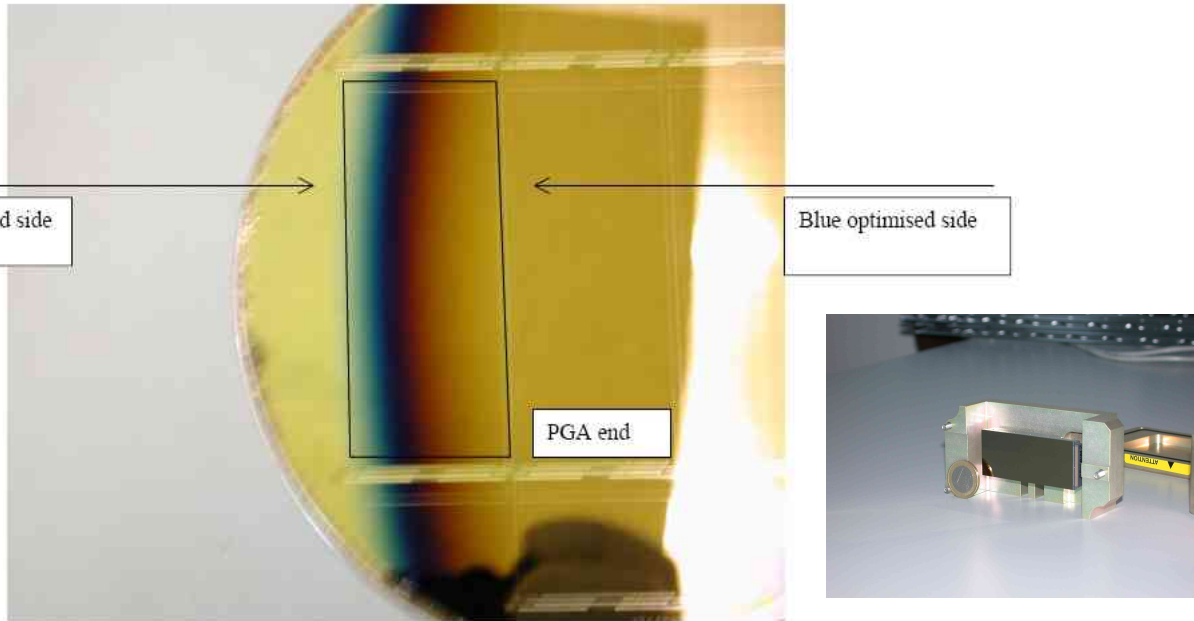
Con



# HRF: slicer



# Graded AR coating: Anti-fringing



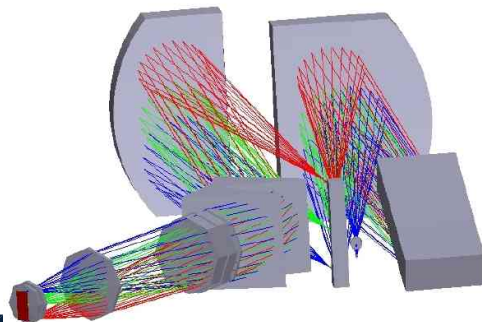
## ***Hermes spectrograph room: tripple isolation***



- 1) Room 1:  $\Delta T$  +/- 0.1 degree (working)
- 2) Room 2:  $\Delta T$  +/- 0.01 degree
- 3) Room 3: passive (installed around table)
- 4) Absolute pressure controle



# ***Spectrograph room now***



# Software: control + user gui

Instrument Control GUI

## HERMES

**HERMES**

Source selector: Thorium-Argon-Neon, Neon, Thorium-Argon, Red halogen, Blue halogen

Transmission: 60.0%

WRF NDF

WRF shutter: Closed

Calibration shutter: Closed

ADC

Fibre selector: Out, LRF, HRF

Fibre mask: All open, HRF open, All closed, LRF open

Guiding NDF: Transmission LRF=55.0% HRF=50.0%

Guiding camera: Default

HERMES

HERMES shutter: Closed, Open

Science CCD

Exposure meter: Off, On

Status: Idle, Freq: 663 Hz, Counts: 245 K, Sigma: 60, Time: 456 s

WRF: LRF, HRF

Elevation: Closed, 64° - 90°, 43° - 64°, 30° - 43°, <30°

LEDs: Off, On

Mode: LRF\_FF

Current status: HRF\_OBJ

Change to: LRF\_FF

Capture flat fields with the LRF

adc: CLOSED, cal\_shutter: OPEN, fibre\_mask: LRFOPEN, fibre\_selector: LRF

Lamps: Th-Ar-Ne: OFF, Ne: WARMING UP (6m58s / 15m), Th-Ar: WARMING UP (6m57s / 15m), Red halogen: OFF, Blue halogen: OFF

Neutral Density Filters: Guiding NDF: HRF transmission 50.0%, WRF NDF: Transmission 60.0%

ADC: Current status: Elevation 64 deg - 90 deg, Set position: Closed

Exposure meter: Observer: Allowed, Denied, Hermes mode: Allowed, Denied, Exposure meter: On, Off, Flux protection: No fault

Advanced: Enable advanced tab, Disable advanced tab, (Re)initialize the instrument, Shutdown the instrument

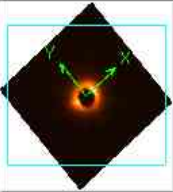
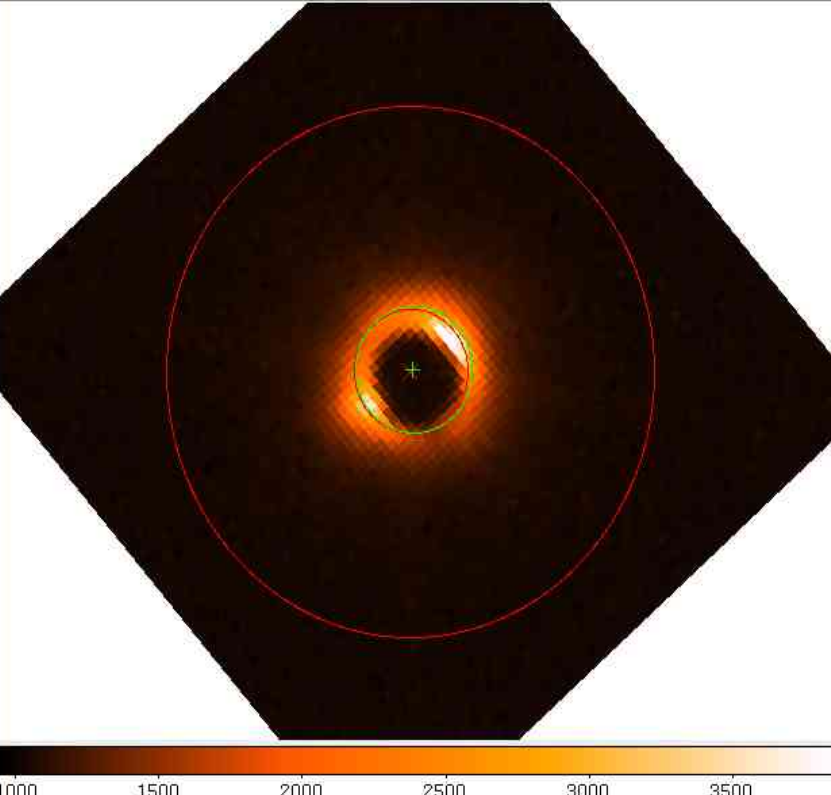
UT	Process	Level	Description
2010-05-06 22:03:21.217	Hermes IC	INFOL4	Getting the status of the calibration lamps
2010-05-06 22:03:21.214	Hermes IC	INFOL4	The Thorium-Argon lamp is now switched ON
2010-05-06 22:03:20.569	Hermes IC	INFOL4	Getting the status of the calibration lamps
2010-05-06 22:03:20.566	Hermes IC	INFOL4	The Neon lamp is now switched ON
2010-05-06 22:03:18.168	Hermes IC	INFOL4	Switching the Thorium-Argon lamp ON
2010-05-06 22:03:17.513	Hermes IC	INFOL4	Switching the Neon lamp ON



### SAOImage Mercator Guiding

File Edit View Frame Bin Zoom Scale Color Region WCS Analysis Help

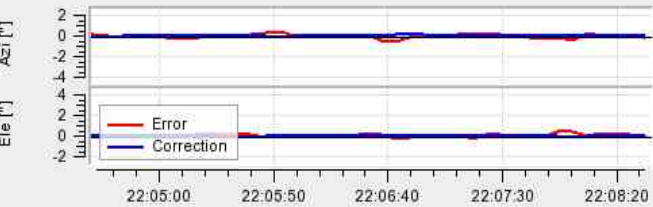
File: stdin  
 Object: Guide star  
 Min: 946 Max: 3870  
 Value:   
 Detector: X Y  
 Image: X Y  
 Frame1: Zoom: 6.000 Angle: 41.200

0000 1500 2000 2500 3000 3500


### Guiding GUI

#### Azimuth, Elevation



Azimuth: Error: -0.08" Corr: 0.00"  
 Elevation: Error: 0.02" Corr: -0.05"

#### Seeing




FWHM: 1.06"

#### Guiding signal and Hermes PMT

Display on graph:  
 Guiding signal  
 Background  
 PMT - frequency  
 PMT - S/N

Observation	V = 9.93, B-V = -0.17, HRF	Change
Max pixel ADU	R: 2238, V: 2269, B: 2231	
S/N	R: 26, V: 31, B: 25	
Aimed S/N	R: 50, V: 50, B: 50	Change
Remaining exp.	R: 689s, V: 433s, B: 772s	

Signal: 4045999  
 Background: 3682861  
 Efficiency: 84.7%



Guiding Settings Detector control Procedures

Status: **IDLE**

Settings: **Default settings**

Instrument: **HERMES** Change...

Guiding detector: **BUSY**

Status: **BUSY**

Integration: **2 sec** 3 sec remaining

Readout: **0 px** 6400 px remaining

Guiding sequence

Start guiding... Stop guiding

MEROPE HERMES

- Set the guiding NDF transmission: **Finished** Transm. 50.0%
- Selection of the target star
  - 2a. Continuously take images: **2 frames acquired** Exp time 5 s
  - 2b. Suggest target star: **Suggestion found** This star Confirm
  - 2c. Confirm target star: **Finished** Timeout Stop timeout
- Center the star on the fiber: **HRF finished** Enable Yes Y N
- Run the guiding loop
  - 4a. Get a windowed image: **Acquiring image...** Exp time 5 s
  - 4b. Find centroid: **Finished** Radius 30 px
  - 4c. Adjust the telescope: **Busy** Enable Yes Y N

UT	Process	Level	Description
2010-05-06 22:08:32.219	Hermes Guiding CCD	INFOL4	Starting exposure (5.0s, OBJ, object 'Guide star', programme -1)
2010-05-06 22:08:32.215	Guiding	INFOL3	Sending offsets azimuth=0.00" elevation=-0.05" to the old TCS ...
2010-05-06 22:08:32.214	Guiding	INFOL3	Centroid found at x=1214.53 y=451.61
2010-05-06 22:08:25.926	Hermes Guiding CCD	INFOL4	Starting exposure (5.0s, OBJ, object 'Guide star', programme -1)

## Scheduler GUI

STATUS	T	I	G	E	Instrument	Obs-mode	R.A.	Dec.	Object	Progr.	Misc.	mV	B-V	S/N (V)	Exp Time	#	
FINISHED	X	X	X	X	HERMES	HRF_OBJ	10 <sup>h</sup> 37 <sup>m</sup> 33.5"	-13° 23' 4.3"	HD92055	10	every 20 days	4.93	2.740	200	60.0	1	blueN last obs: 2455322
FINISHED	X	X	X	X	HERMES	HRF_OBJ	9 <sup>h</sup> 13 <sup>m</sup> 50.1"	+14° 12' 39.1"	HD79319	10	every 20 days	8.57	1.670	40	200.0	1	blueJ-R last obs: 2455322
FINISHED	X	X	X	X	HERMES	HRF_OBJ	11 <sup>h</sup> 38 <sup>m</sup> 44.9"	+45° 6' 30.2"	HD101177	1		6.45	0.570	10	40.0	1	RV stand
BUSY	X	X	X	X	HERMES	HRF_OBJ	10 <sup>h</sup> 38 <sup>m</sup> 55.4"	+10° 3' 48"	V*DNLeo	10	every 20 days	9.93	-0.170	50	800.0	1	pagb last obs: 2455295
	X	X	X	X	HERMES	HRF_OBJ	15 <sup>h</sup> 35 <sup>m</sup> 28"	+43° 28' 49.1"	YY Boo	26	every night 2 exposures well ...	11.9	0.000	20	600.0	2	!!!!!! observed at
	X	X	X	X	HERMES	HRF_OBJ	10 <sup>h</sup> 56 <sup>m</sup> 57.7"	+15° 16' 46.3"	HIP53522	10	every 20 days	10.11	1.290	50	900.0	1	R:ORB last obs: 2455296
	X	X	X	X	HERMES	LRF_WRF_OBJ_TH	15 <sup>h</sup> 8 <sup>m</sup> 51.5"	+2° 20' 35.9"	WASP-24	6		11.3	0.800	50	2000.0	1	!!!! Every night two times LRF+WRF!
	X	X	X	X	HERMES	HRF_OBJ	11 <sup>h</sup> 7 <sup>m</sup> 26.3"	+24° 3' 12.6"	PG1104+243	10	every 10 days	11.29	0.000	30	1200.0	1	sdOB+G last obs:
									STOP								Stop the observations
	X	X	X	X	HERMES	HRF_OBJ	11 <sup>h</sup> 0 <sup>m</sup> 48.5"	+40° 42' 10.5"	HIP53832	10	every 20 days	10.0	1.170	50	820.0	1	R:ORB last obs: 2455297
	X	X	X	X	HERMES	HRF_OBJ	11 <sup>h</sup> 14 <sup>m</sup> 36.6"	+33° 40' 26.8"	PG1111+339	10	every 30 days	12.52	0.000	99	2700.0	1	sdB+G last obs:
	X	X	X	X	HERMES	HRF_OBJ	11 <sup>h</sup> 15 <sup>m</sup> 23"	-11° 35' 17.5"	V*UUCrt	10	every 20 days	7.3	1.500	50	120.0	1	M:ORB last obs: 2455287
	X	X	X	X	HERMES	HRF_OBJ	11 <sup>h</sup> 18 <sup>m</sup> 30.2"	+52° 41' 41.9"	DZUMa	10	every 10 days	12.0	0.000	20	2700.0	1	disk last obs:
	X	X	X	X	HERMES	HRF_OBJ	11 <sup>h</sup> 22 <sup>m</sup> 49"	+43° 28' 57.1"	HD98839	10	every 50 days	4.99	0.980	200	60.0	1	Ba:ORB last obs: 2455294
	X	X	X	X	HERMES	HRF_OBJ	11 <sup>h</sup> 23 <sup>m</sup> 18.5"	-21° 38' 33"	HE1120-2122	10	every 20 days	11.93	1.510	30	1800.0	1	CEMP:SB? last obs:
	X	X	X	X	HERMES	HRF_OBJ	11 <sup>h</sup> 35 <sup>m</sup> 42.4"	-14° 35' 36.6"	CGCS3066	10	every 20 days	8.73	1.080	50	300.0	1	J-Disk last obs: 2455292
	X	X	X	X	HERMES	HRF_OBJ	11 <sup>h</sup> 35 <sup>m</sup> 42.8"	-14° 35' 36.4"	HD 100764	10	every 30 days	8.73	1.080	30	600.0	1	R - IR excess last obs:
	X	X	X	X	HERMES	HRF_OBJ	11 <sup>h</sup> 37 <sup>m</sup> 54.4"	-1° 9' 3.1"	HD101079	10	every 30 days	8.2	1.020	50	200.0	1	Ba:ORB last obs: 2455292
	X	X	X	X	HERMES	HRF_OBJ	11 <sup>h</sup> 49 <sup>m</sup> 48.3"	-8° 17' 20.3"	IRAS11472-0800	10	every 20 days	11.7	0.500	20	2700.0	1	disk last obs: 2455298
	X	X	X	X	HERMES	HRF_OBJ	11 <sup>h</sup> 55 <sup>m</sup> 56.4"	+9° 50' 49.9"	BD+10.2357	10	every 10 days	8.87	0.000	50	800.0	1	sdO+A last obs: 2455291
	X	X	X	X	HERMES	HRF_OBJ	12 <sup>h</sup> 7 <sup>m</sup> 10.7"	+12° 59' 7.8"	HD105262	10	every 20 days	7.09	0.010	80	200.0	1	pagb last obs: 2455287
	X	X	X	X	HERMES	HRF_OBJ	12 <sup>h</sup> 12 <sup>m</sup> 52.1"	+33° 54' 2.5"	HZ19	10	every 30 days	12.72	0.000	20	2700.0	2	sdOB+F last obs: 2455308
	X	X	X	X	HERMES	HRF_OBJ	12 <sup>h</sup> 24 <sup>m</sup> 32"	-28° 18' 55.4"	V420Hya	10	every 20 days	10.1	0.900	50	900.0	1	S:symbio last obs: 2455293
	X	X	X	X	HERMES	HRF_OBJ	12 <sup>h</sup> 24 <sup>m</sup> 49"	+31° 2' 1.8"	HD108078	10	every 20 days	7.55	1.140	70	200.0	1	K:SB+abun last obs: 2455287
	X	X	X	X	HERMES	HRF_OBJ	12 <sup>h</sup> 37 <sup>m</sup> 23.1"	+25° 3' 59"	Feige66	10	every 30 days	10.51	0.000	40	2700.0	1	sdOB last obs: 2455297
	X	X	X	X	HERMES	HRF_OBJ	12 <sup>h</sup> 43 <sup>m</sup> 56.7"	+61° 5' 35.9"	V*SUMa	10	every 20 days	8.87	2.430	50	600.0	1	M:ORB last obs: 2455320
	X	X	X	X	HERMES	HRF_OBJ	12 <sup>h</sup> 44 <sup>m</sup> 7"	+2° 44' 38"	BD+03.2688	10	every 20 days	10.5	1.000	50	1200.0	1	mdBa last obs: 2455292
	X	X	X	X	HERMES	HRF_OBJ	12 <sup>h</sup> 46 <sup>m</sup> 37.1"	+47° 22' 20.8"	V*BYCVn	10	every 20 days	7.94	1.400	50	300.0	1	M:ORB last obs: 2455308
	X	X	X	X	HERMES	HRF_OBJ	12 <sup>h</sup> 52 <sup>m</sup> 58.4"	-14° 37' 14.7"	V*VWCrv	10	every 50 days	7.59	1.600	50	250.0	1	M:ORB last obs: 2455308
	X	X	X	X	HERMES	HRF_OBJ	12 <sup>h</sup> 53 <sup>m</sup> 32.4"	-22° 52' 22.8"	V*LWHya	10	every 30 days	9.69	0.830	50	800.0	1	Abell-35 last obs:
	X	X	X	X	HERMES	HRF_OBJ	12 <sup>h</sup> 54 <sup>m</sup> 21.2"	-9° 32' 20"	V*psiVir	10	every 50 days	4.81	1.580	50	30.0	1	M:ORB last obs: 2455284

Insert row(s) after the selection




Empty row(s)




Duplicate

Insert calibrations



Multiple...



Single ...

Special actions



Insert stop



Insert wait...

Move/Cut/copy/paste/delete



Append/save/save as file



Scheduler settings

- Stop after each line
- Beep after each line
- Autofocus at each pointing

Actions during AUTO sequence

 T: Telescope pointing  I: Instrument / filter wheel  G: Guiding (+acquisition)  E: Exposure

Object

 RA  Dec  V mag  B-V 

 Object  Programme ID  S/N in V 
Comments 

Observation

 Merope  Hermes

 Obs mode  Exp time  # 

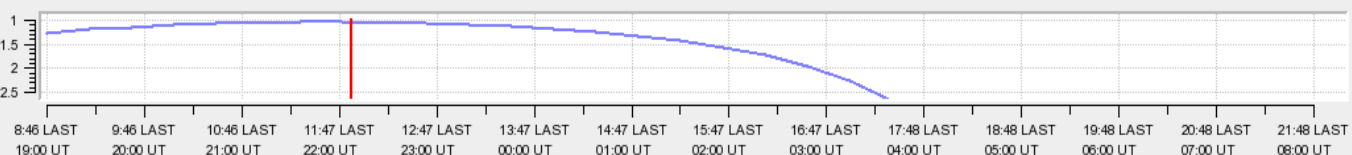
Sequence

Telescope Thermal focus finished

Instrument ADC has been set

Guiding OK

Exposing Exposure(s) busy



AUTO sequence

START

STOP

MANUAL sequence

Point telescope

Set instrument

Start guiding

Start exposure(s)

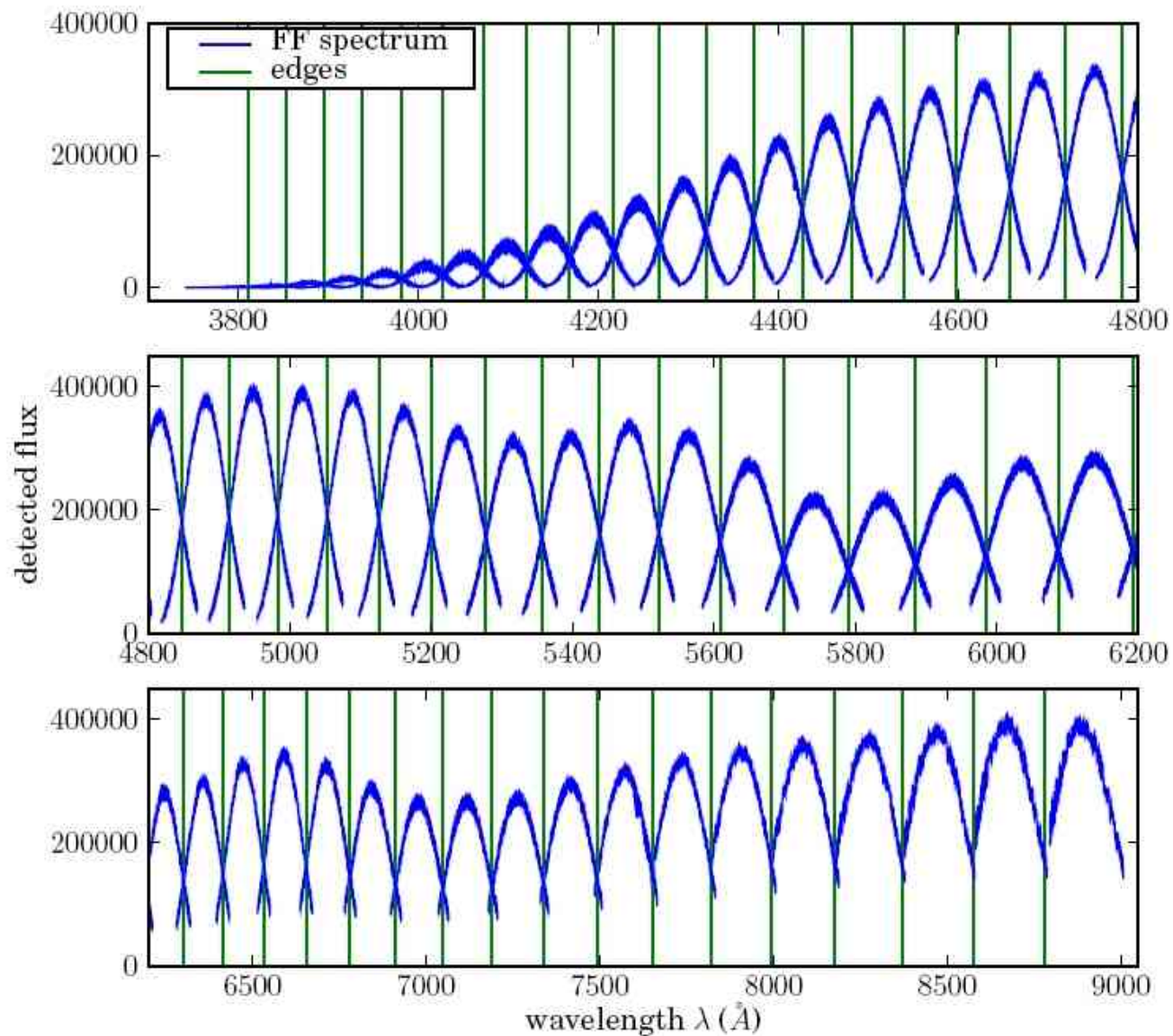
# ***WP900: Data Reduction System***

**Herman Hensberge, Yves Fremat, Louis Dumortier, Alain Jorissen,  
Sophie Van Eck, Christos Siopis, Hans Van Winckel  
WP that ran during whole project (and is still running)**

**Take maximal advantage of stability of spectrograph**

- good instrumentmodel !**
- detailed algorithms**
- coded in python**
- quality control during observations, fine detailed reduction/night**
  
- order tracing, bias correction, background mapping/correction,  
extraction with/w.o. clipping, wavelengthcalibration, merging,  
CCF routines**

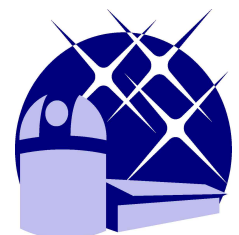




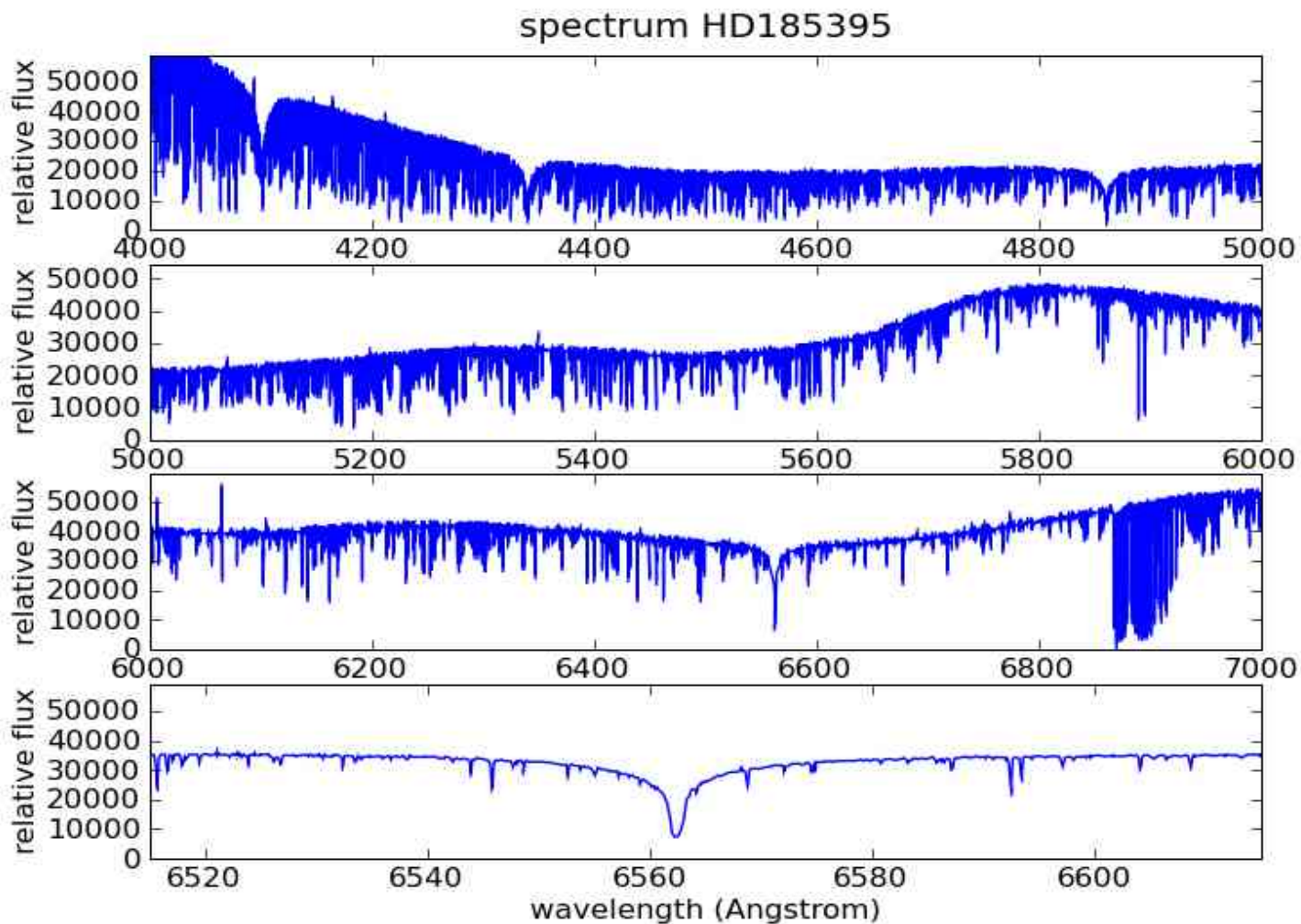
- blaze profiles  
so S/N is very  
dependent on the  
distance from the  
blazewavelength

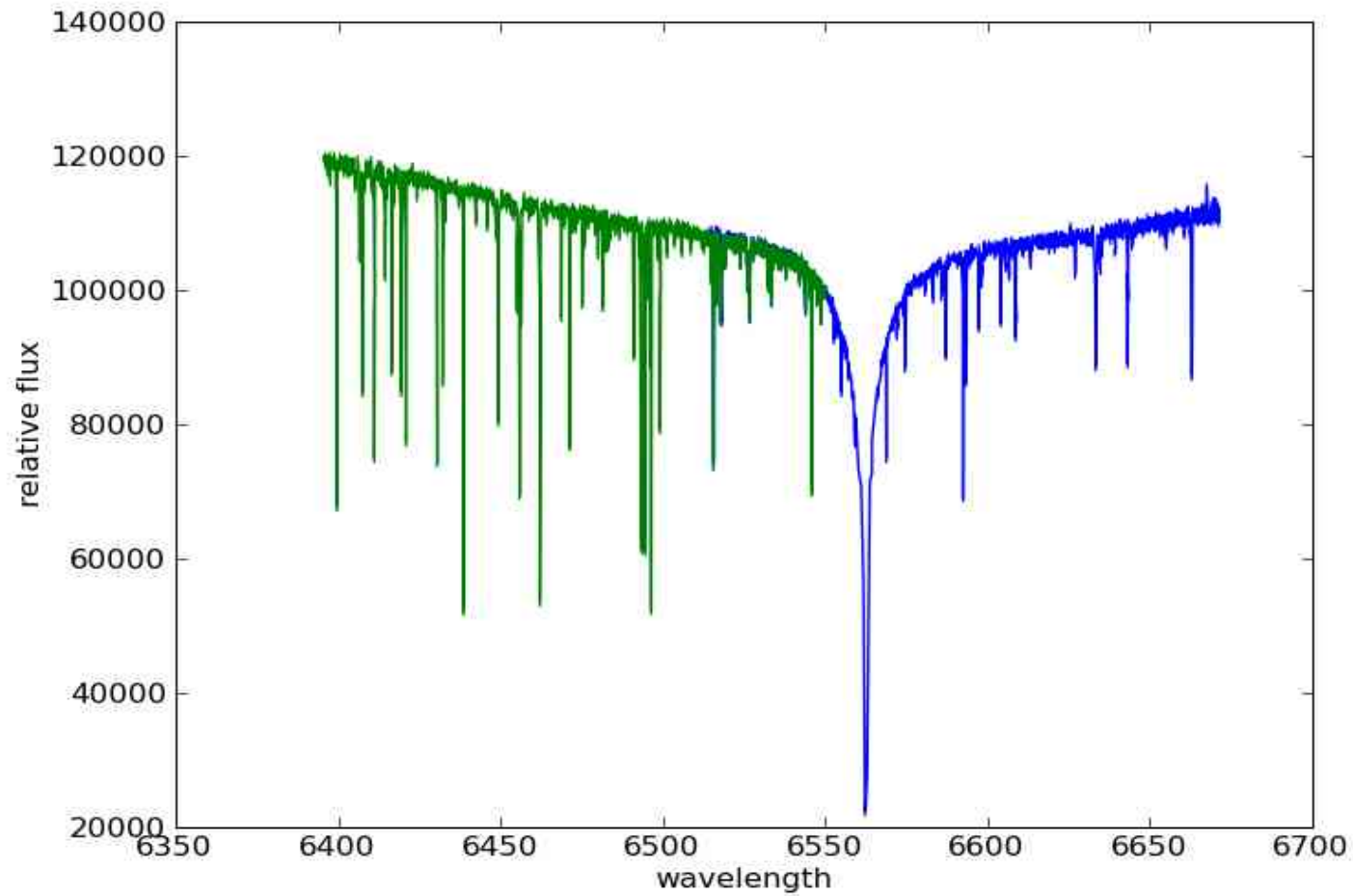
- reddest orders not  
overlap.

Orders: 40 to 94



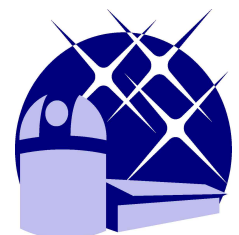
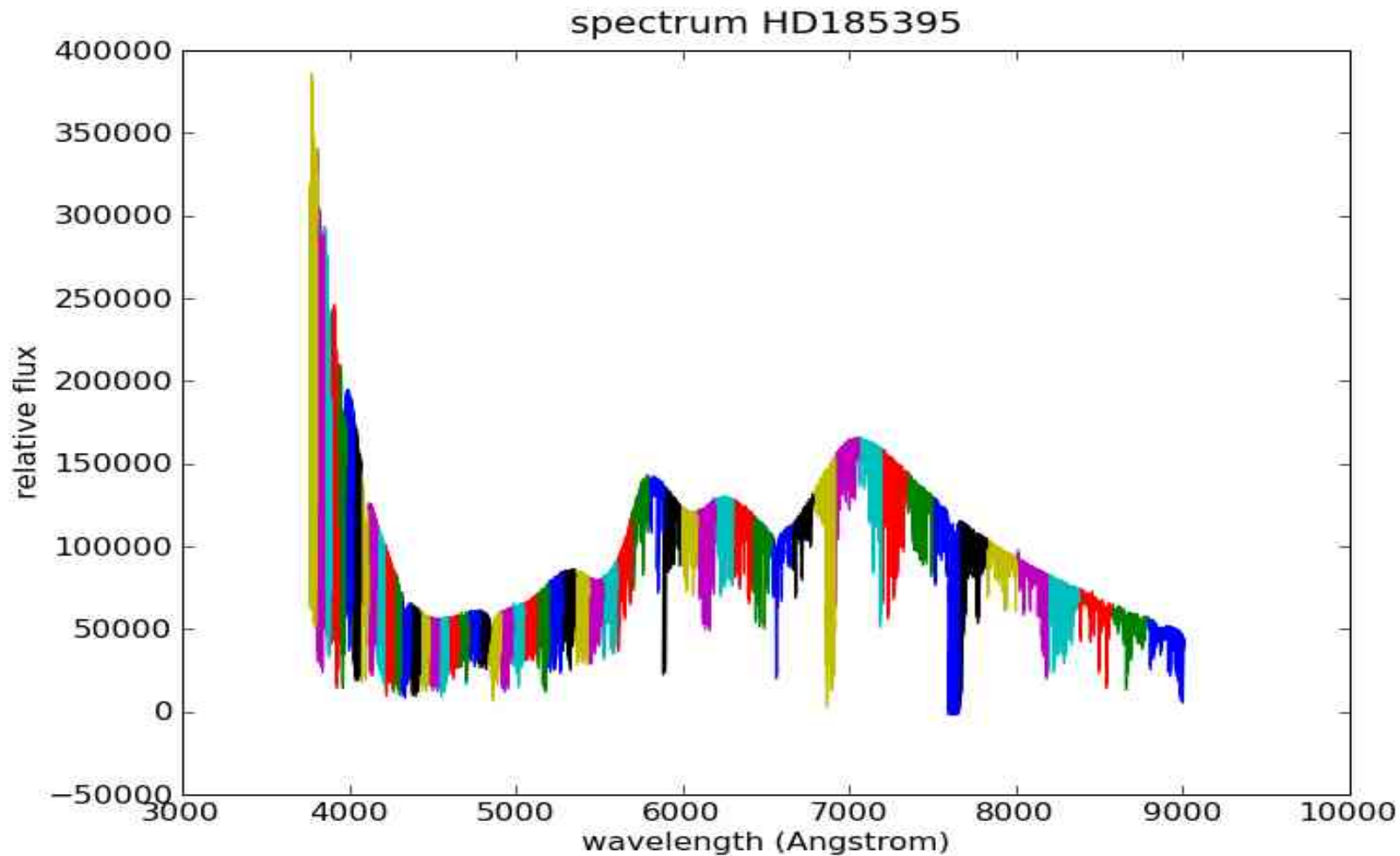
# ***DRS: Merged Spectrum 1D***





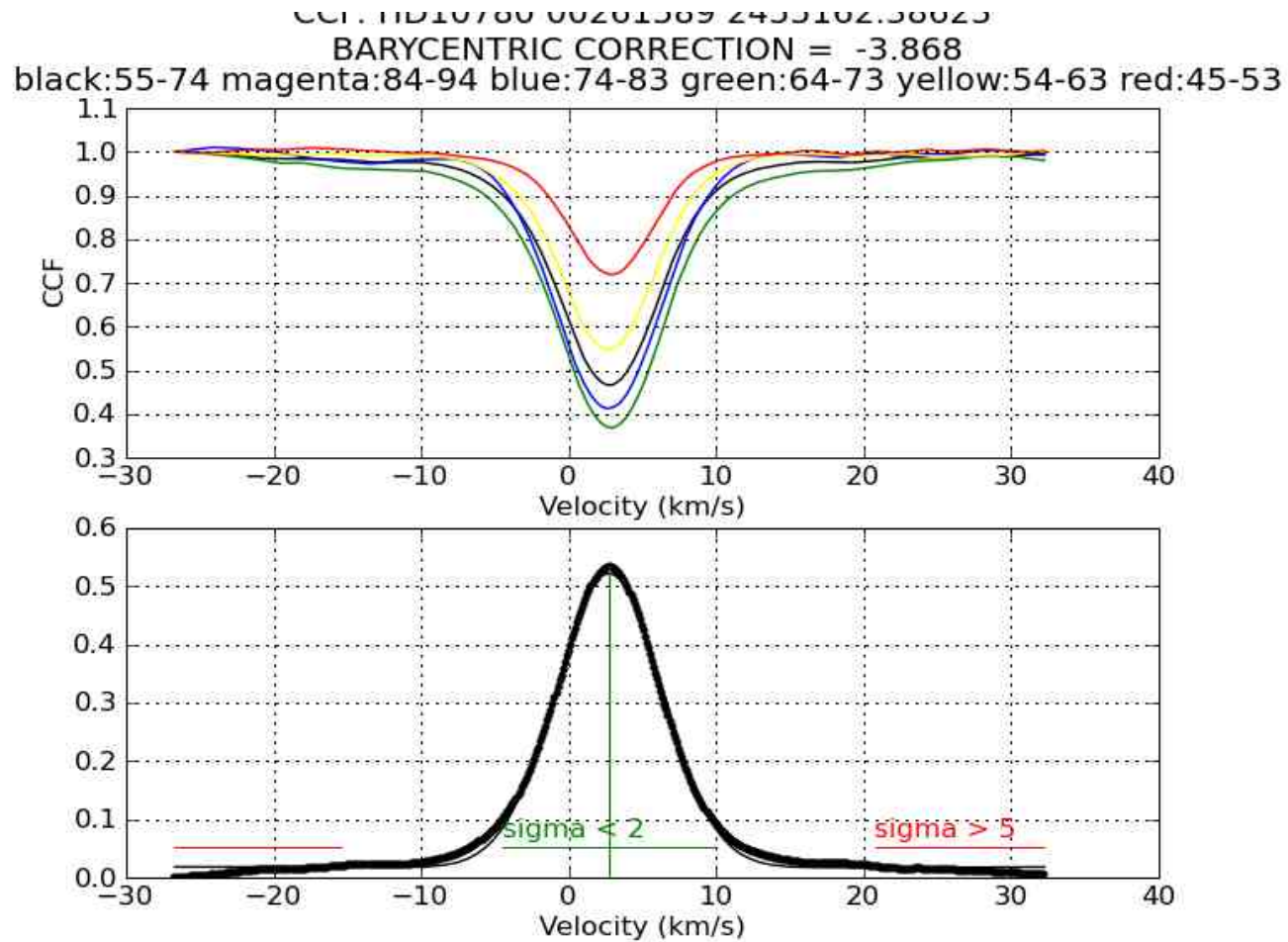
# ***Slope = star + filtered FF lamp combination***

*FF normalisation is by division of median of whole extracted FF*



# Spectral Masks: HD 10780 (K0V) $v_{\text{helio}}$ : 2.7 km/s

file 00261589 of 20091126





# Efficiency of Mercator-HERMES

$\lambda$ nm	LRF	HRF	F/N+Slic.	4 mirrors	Echelle	CrossD.	Camera	Spectro.	CCD	Total LR	Total HR
380	0.49	0.70	0.90	0.60	0.60	0.85	0.75	0.21	0.75	0.08	0.11
400	0.51	0.71	0.90	0.83	0.60	0.86	0.80	0.31	0.80	0.13	0.18
450	0.51	0.73	0.90	0.91	0.61	0.86	0.85	0.37	0.85	0.16	0.23
500	0.52	0.74	0.90	0.95	0.62	0.88	0.85	0.40	0.90	0.19	0.26
600	0.52	0.75	0.90	0.93	0.63	0.88	0.85	0.39	0.85	0.17	0.25
700	0.52	0.75	0.90	0.92	0.64	0.89	0.85	0.40	0.80	0.17	0.24
800	0.52	0.75	0.90	0.90	0.65	0.89	0.85	0.40	0.60	0.12	0.18
900	0.52	0.75	0.90	0.89	0.65	0.89	0.85	0.39	0.30	0.06	0.09

Table 9. Estimated efficiency in low-and high-resolution mode.

Example:

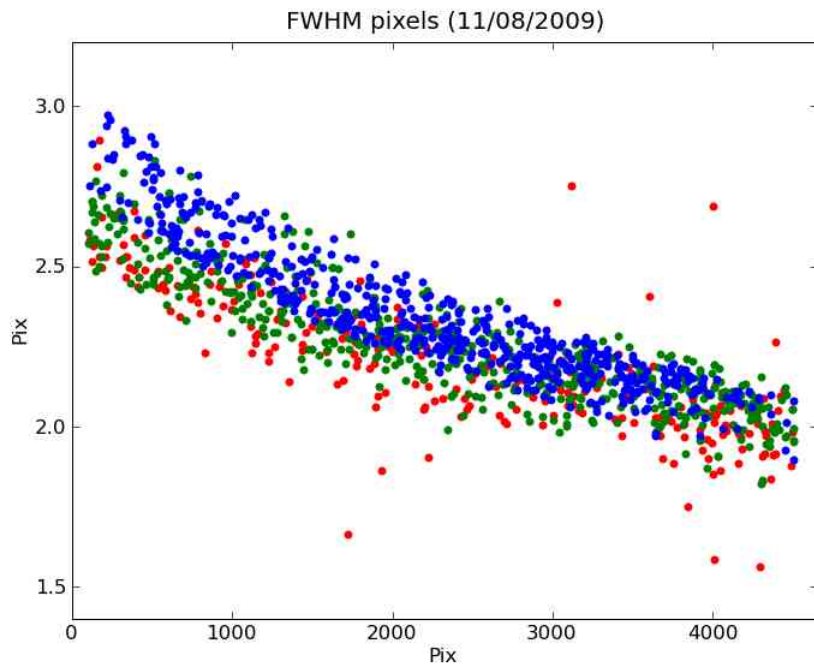
Good seeing 1.2 arcsec

10 min exposure , 9<sup>th</sup> magn star: S/N 110 per resolution element for the HRF  
S/N 74 per pixel for the HRF

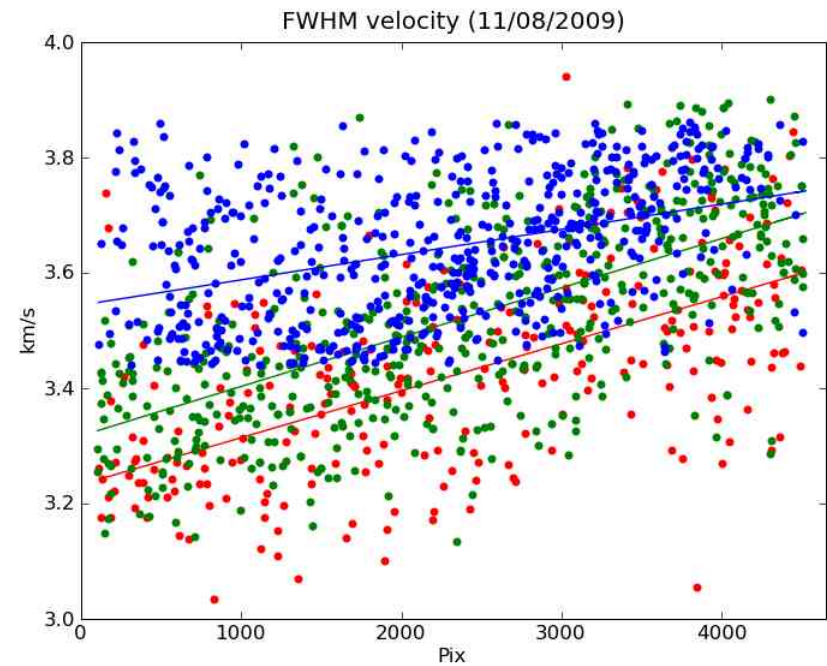
60 min exposure, 14<sup>th</sup> magn star: S/N 40 per resolution element for the HRF  
S/N 29 per pixel for the HRF



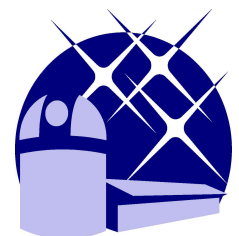
# Image Quality



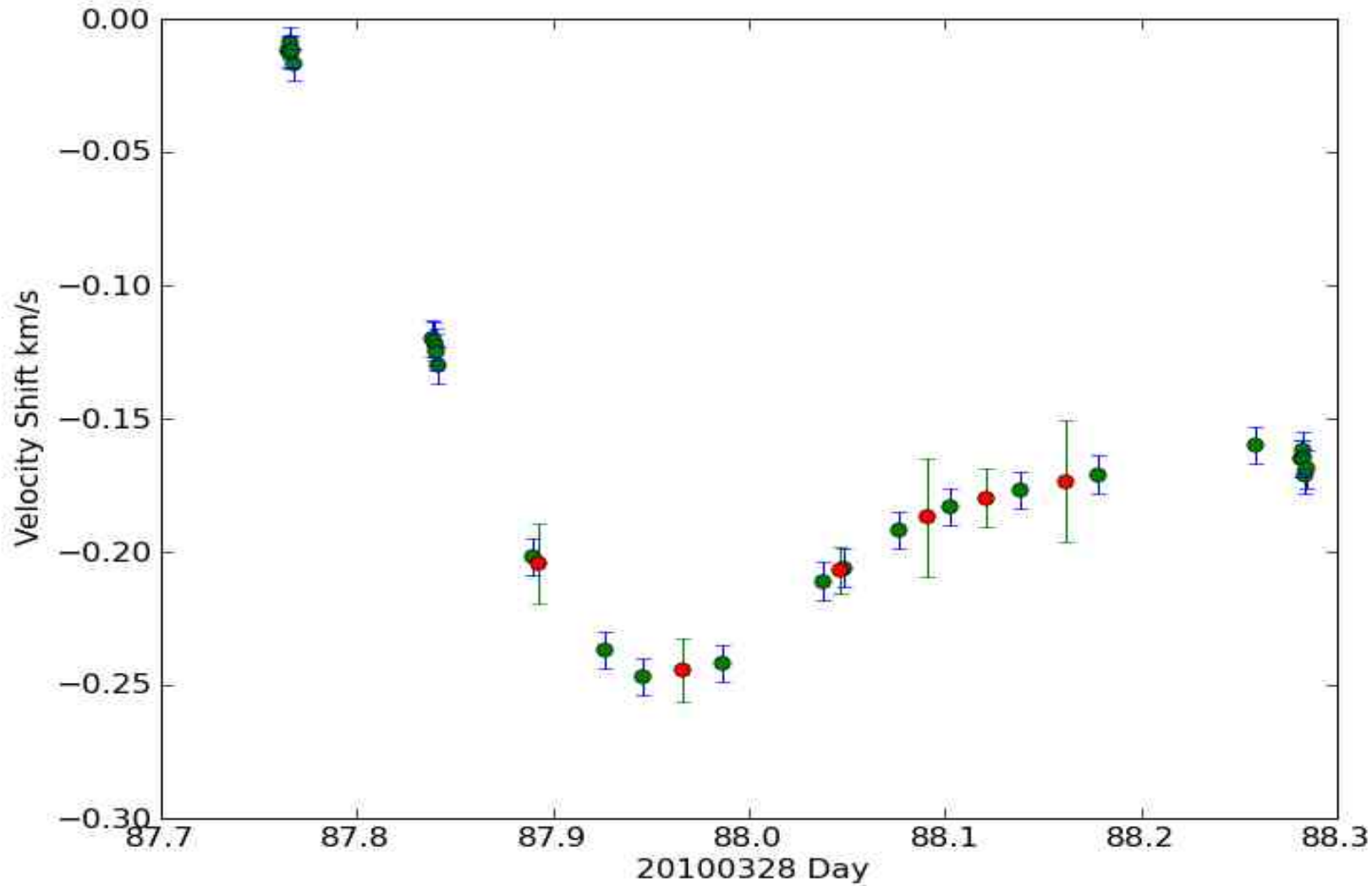
**HRF pixel space.  
FWHM of a resolution element**



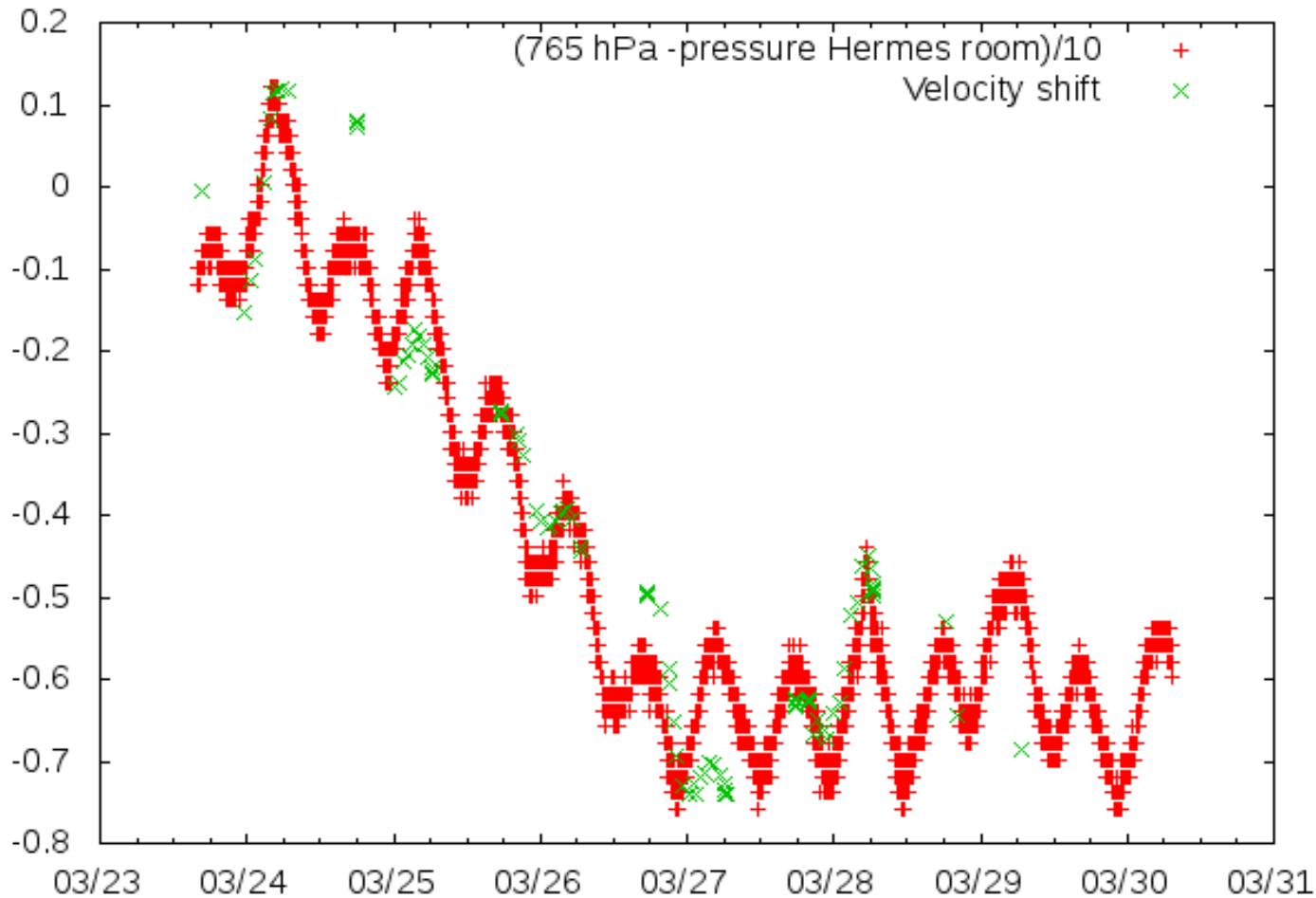
**Velocity space**



# Stability of Hermes

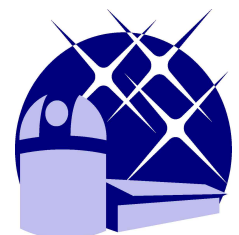


Velocity shift and atmospheric pressure in Hermes room



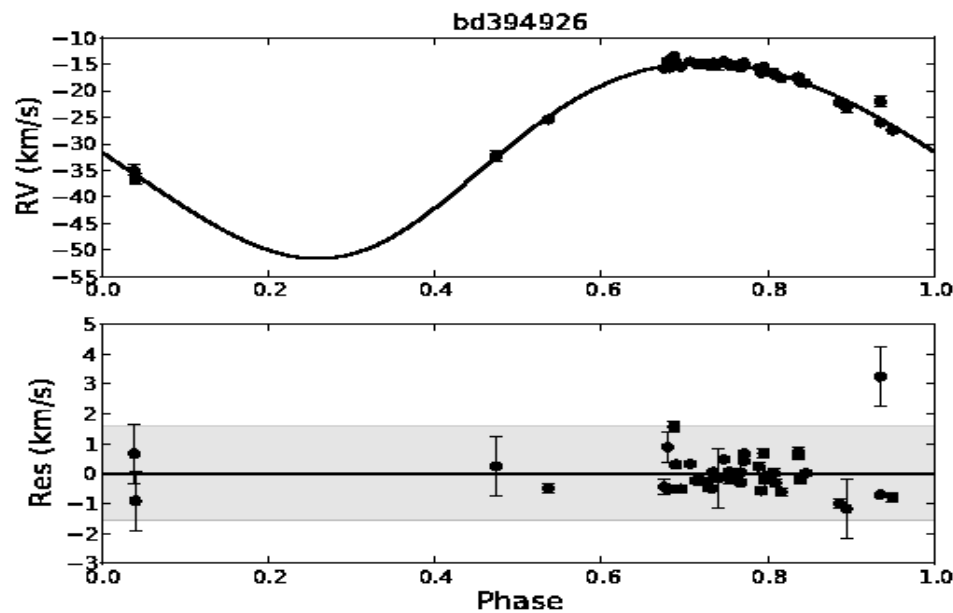
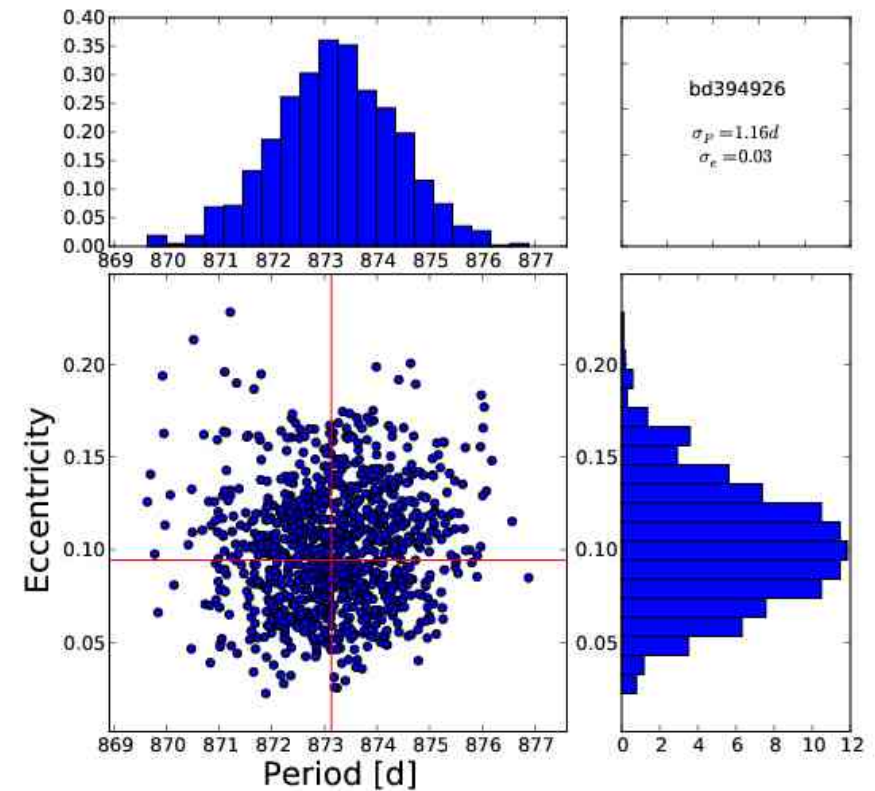
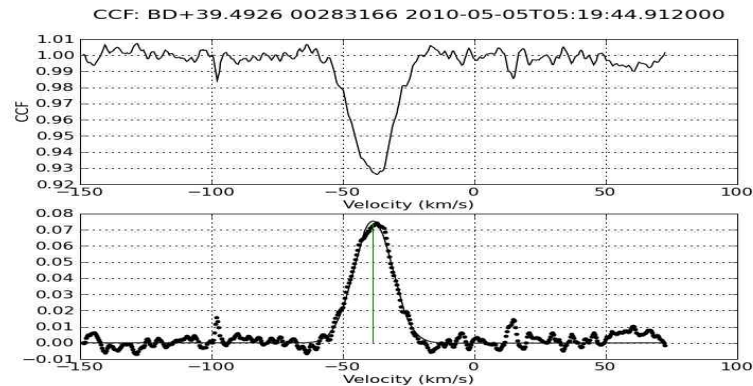
**For high-accurate vr  
Work use the LRF**

**Noise in limited LRF  
Experiments till now  
2 m/s**



# Some illustrative results: **BD+39.4926**

Very depleted post-AGB binary  $[Fe/H] = -3.2$ . Binary Orbit 873 days



# Hyades stream: evaporated population / thin disc resonance effect ?

Pompeia et al.

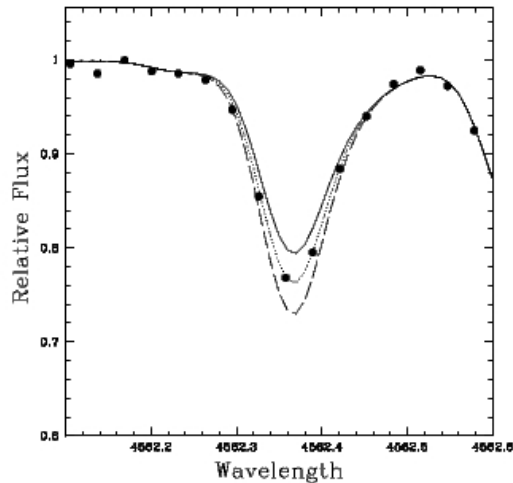
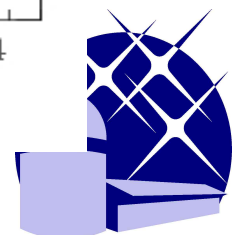
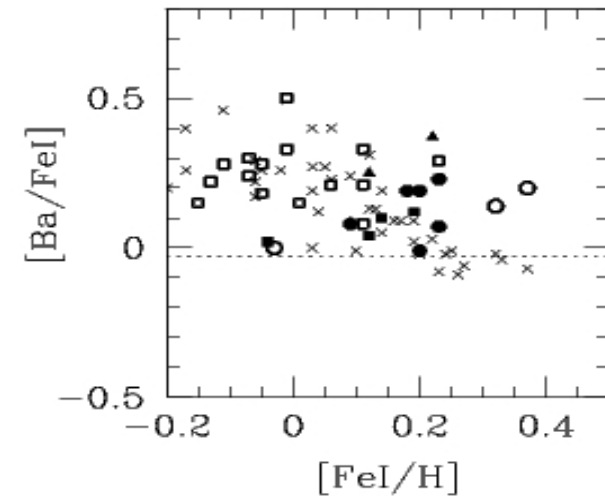
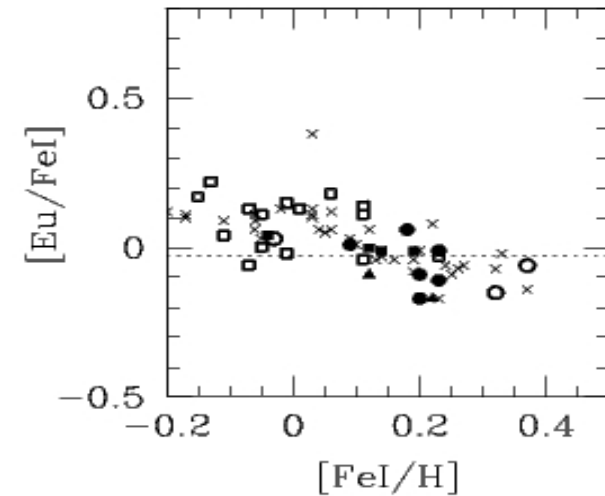


Fig.7. Example of line synthesis for the Ce II line at 456.2 nm in HD 171067. The dots depict the observed spectrum, the dashed line shows the synthetic spectrum for  $[\text{Ce}/\text{Fe}] = 0.35$  dex, the dotted line for  $[\text{Ce}/\text{Fe}] = 0.25$  dex and the solid line for  $[\text{Ce}/\text{Fe}] = 0.15$  dex.

Resolution + spectral coverage ideal  
For abundance studies of bright stars.



# ***Hermes project***

After:

- many hundreds of mails, skype sessions, discussions
- detailed optical, mechanical design figs.+ (electronic) control  
(very much instrument of Gert Raskin)
- hundreds of ATBs and purchased items (administration...)
- thousands of lines of codes,

***Science exploitation at full swing !***

- *semester allocation*
- *pooled observation*
- *DB+ more automatic Scheduling*
- *DRS fully relative*



Contact Group Meeting 25/05/2010

